# The Fermi-GBM 3-year X-ray Burst Catalog

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#### ABSTRACT

The Fermi Gamma Ray Burst Monitor (GBM) is an all sky gamma-ray monitor well known in the gamma-ray burst community. Although GBM excels in detecting the hard, bright extragalactic GRBs, its sensitivity above 8 keV and all-sky view make it an excellent instrument for the detection of rare, short-lived Galactic transients. In March 2010, we initiated a systematic search for transients using GBM data. We conclude this phase of the search by presenting a 3 year catalog of 1084 X-ray bursts. Using spectral analysis, location and spatial distributions we classified the 1084 events into 752 thermonuclear X-ray bursts, 267 transient events from accretion flares and X-ray pulses, and 65 untriggered gamma-ray bursts. All thermonuclear bursts have peak blackbody temperatures broadly consistent with photospheric radius expanison (PRE) bursts. We find an average rate of 1.4 PRE bursts per day, integrated over all Galactic bursters within about 10 kpc. These include 33 and 10 bursts from the ultra-compact X-ray binaries 4U 0614+09 and 2S 0918-549, respectively. We discuss these recurrence times and estimate the total mass ejected by PRE bursts in our Galaxy.

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Subject headings: type I X-ray burst, Neutron Stars, Accreting Pulsars

#### 1. Introduction

Rare, unpredictable and transient astronomical phenomena are difficult to observe due to their very own nature, yet they often lead to exciting astrophysical discoveries. At any wavelength, the most efficient way of detecting rare transients is to maximize the observed field of view (FoV). The high-energy (X-ray and Gamma-ray) sky can vary rapidly, on timescales much shorter than hours. If we are interested in *short-lived* rare transient phenomena (seconds to minutes long), the most relevant observational capability is the *instantaneous* FoV. Even though serendipitous detections occur, neither pointed narrow FoV instruments nor all-sky monitors based on scanning techniques are well suited to catch such short and rare events.

The Gamma-ray burst monitor (GBM) onboard the *Fermi* observatory has an instantaneous FoV of about 75% of the sky (Meegan et al. 2009) and is sensitive to photon energies down to 8 keV. Even though it was designed to detect and characterize Gamma-ray bursts (GRBs), these characteristics make GBM a unique instrument to detect rare, short and bright X-ray bursts (XRBs). In March 2010, we initiated a systematic search for XRBs using *Fermi*-GBM data (Section 2). In the first three years, this search has yielded 752 thermonuclear X-ray bursts (tXRBs; Secs. 1.1 & 3.1), 267 transient events from accretion flares and X-ray pulses (aFXPs), as well as 65 untriggered long gamma ray bursts (uGRBs). We present here the *Fermi*-GBM 3-year X-ray Burst Catalog and summarize its main results with an emphasis on tXRBs.

# 1.1. The rare and most energetic thermonuclear bursts

The accreted matter in neutron star low-mass X-ray binaries (NS-LMXBs) piles up on the neutron star surface, reaching regions of increased density and becoming fuel for thermonuclear reactions. When ignition conditions are met at the bottom of the accreted shell, unstable reactions trigger a thermonuclear runaway that quickly burns the pile of fuel, generally a mix of hydrogen (H), helium (He) and heavier elements ("metals"). This cyclic phenomenon has been observed for four decades in what is know as thermonuclear (type I X-ray) bursts (Grindlay et al. 1976; Belian et al. 1976).

The main parameter which sets the frequency or recurrence time of thermonuclear X-ray Bursts (tXRBs) is the mass accretion rate per unit area,  $\dot{m}$  (Fujimoto et al. 1981;

Bildsten 1998). The main reason is simple:  $\dot{m}$  sets the rate at which fuel is replenished between tXRBs. However, other factors play an important role in the tXRB recurrence time, including composition and the thermal state of the NS envelope. In particular, at the lowest  $\dot{m}$  (near or below 1% of the Eddington limit) the heat flux from the NS crust can critically influence the ignition conditions for tXRBs. Thus we can potentially use low- $\dot{m}$  tXRBs to constrain the internal properties of NSs (Cumming et al. 2006). However, because recurrence times of low- $\dot{m}$  tXRBS are of the order of weeks to months, they are extremely difficult to measure with pointed or scanning X-ray detectors. GBM has opened a new window to these events, and it is yielding the first accurate measurements of their recurrence times (Linares et al. 2012).

# 2. The Fermi-GBM X-ray Burst Monitor

GBM is an all sky monitor whose primary objective is to extend the energy range over which gamma-ray bursts are observed in the Large Area Telescope (LAT) on *Fermi* (Meegan et al. 2009). GBM consists of 12 NaI detectors with a diameter of 12.7 cm and a thickness of 1.27 cm and two BGO detectors with a diameter and thickness of 12.7 cm. The NaI detectors have an energy range from 8 keV to 1 MeV while the BGOs extend the energy range to 40 MeV. The GBM flight software was designed so that GBM can trigger on-board in response to impulsive events, when the count rates recorded in two or more NaI detectors significantly exceed the background count rate on at least one time-scale from 16 ms to 4.096 s in at least one of four energy ranges above 25 keV. The lower energy and longer time-scales are inaccessible to the on-board triggering algorithms owing to strong variations in background rates that are incompatible with a simple background modeling needed for automated operation on a spacecraft. Between 25 and 50 keV, only the shortest time-scales are probed on-board (under 128 ms). We report here on our search of GBM continuous data for impulsive events that are too long and too spectrally soft to trigger on-board.

GBM has three continuous data types: CTIME data with nominal 0.256-second time resolution and 8-channel spectral resolution used for event detection and localization, CSPEC data with nominal 4.096-second time resolution and 128-channel spectral resolution which is used for spectral modeling, and Continuous Time Tagged Event (CTTE) data with time stamps ( $2\mu$ s precision) on individual events at full 128-channel spectral resolution that was made available November 2012. The NaI CTIME and CSPEC data from 8-50 keV are used in the following analysis.

#### 2.1. Data Selection

The Fermi-GBM X-ray Burst Monitor relies on daily inspection of CTIME channel 1 (12-25 keV) data and began operations in 2010 March 12. The CTIME data are rebinned to a minimum of 0.25 second time bins to adjust intervals of high resolution data initiated by instrument triggers. NaI detector rates, from all 12 detectors and channels 0-2 (8-50 keV), are automatically filtered removing phosphorescence events, times of high total rates, times near the SAA and intervals of rapid spacecraft slews. An empirical background model is fit to the detector rates in each channel (0-2) and each detector. The background model has terms to account for bright sources and their Earth occultations plus a quadratic spline model to account for the low frequency trends of the remaining background (below  $\sim$ 1 mHz). The background model is visually compared to the rates in the energy band between 12 and 25 keV with a time resolution of  $\sim$  8.2 seconds. Transient events that rise above the background model are saved by manually selecting the corresponding time intervals. Source rates and background rates for the first three energy bands (8 - 50 keV) along with mid-times of these manually-selected time intervals are recorded. Between March 2010 and March 2013, the search resulted in 5093 selected events.

type I X-ray bursts, the softest population of events likely to be detected, are expected to have a blackbody spectrum with a temperature between about 0.5 and 3 keV. Due to the gradual rollover in the expected photon spectrum between 12 and 25 keV and the steep drop in effective area in CTIME channel 0 ( $\sim$  8-12 keV) data (Meegan et al. 2009), channel 1 (12-25 keV) is the most sensitive channel to these XRBs. The choice of 8.2 second timing resolution for channel 1 data is a compromise between the desire to maximize our sensitivity to these events and the time demands of this labor-intensive process, and limits the minimal detectable burst duration to around 10 seconds. Background count rate variations over the Fermi orbit, caused both by changes in geomagnetic latitude and varying spacecraft attitude, prevent visual identification of very long bursts. Our search is thus sensitive to bursts and flares with durations in the 10 – 1000 s range.

#### 2.2. Localization

Localization of our events of interest utilizes the angular response of the NaI detectors to reconstruct the most likely arrival direction based on the differences in background-subtracted count rates recorded in 12 NaI detectors that have different sky orientations. The method is adapted from the method used for GBM GRB localization (Connaughton et al. 2015), with a cruder background fitting method. We use data between 12 – 50 keV and the model rates more suitable for sources with softer energy spectra: galactic transients

(power-law with index = -2), solar flares (power-law with index = -3), and type I XRBs (blackbody with temperature = 4 keV). This process yields a localization and a 68% statistical uncertainty radius (assuming a circular uncertainty region),  $\sigma$ . We also determine a goodness-of-fit parameter,  $\chi^2$ , of the localization. Other parameters of interest include a rough event duration, a list of detectors with an angle between source and detector normal less than 60°, the net count rates in these detectors, and hardness ratios derived from count rates in different energy channels. If the event localizes to within 10° of the centroid of the solar disk or is less than  $3\sigma$  from the Sun position then the event is rejected, as are events with localizations clearly (beyond the statistical uncertainty) beneath the Earth's horizon.

If the net count rates of the two brightest detectors are inconsistent with a single source direction then the event is rejected. Such events may occur during a particle shower within or near the spacecraft and are not associated with an astronomical source. An additional check is performed to eliminate particle events which originate in the spacecraft. These events appear to have a hard spectra and thus might be initially classified as uGRBs but unlike GRBs their light curves in the 50-300 keV range are very similar for all 12 NaI detectors. This produces a poor  $\chi^2$  in the localization fit and we use a cut-off in  $\chi^2$  of 1000 to reject these particle events, more tolerant than reported in (Connaughton et al. 2015) because the quality of the background fits over the low-energy channel data analyzed here is more variable, and even real astrophysical events may produce localizations with large  $\chi^2$  values. All other events are considered XRB candidates. Once the events are localized, they are searched for temporal and spatial coincidence with GBM and Swift triggered GRBs. If the XRB candidate event locates within  $3\sigma$  of a triggered GRB and the XRB candidate event mid-time occurs within 150% of the T-90 duration of the GRB trigger time then the XRB event is considered a triggered GRB and rejected. After these filtering steps there are 2253 events remaining of the original manually selected sample of 5093. The vast majority of rejected events were identified as solar flares.

### 2.3. Spectral Analysis

Response matrices for each XRB candidate event are created from a response model constructed from simulations incorporating the Fermi spacecraft mass model into GEANT4 (Agostinelli and et (2003)). CSPEC data are used for spectral analysis in RMFIT, a forward-folding spectral analysis software often used in GBM gamma ray burst studies.<sup>1</sup> Through localization and visual inspection (see Section 3.2) many of the events were identified with

<sup>&</sup>lt;sup>1</sup>https://gamma-wiki.mpe.mpg.de/GBM/RMFITPublicReleasePage

Sco X-1 and Vela X-1 (aFXPs) and spectral analysis was not necessary for identification. We did; however, performed spectral analysis on a few of these in order to aid in the association of those events in which identification was not apparent. Blackbody and power-law models are fit to all of the remaining data, the former because it is physically motivated for tXRBs and the latter because it is a simple model that can be used to fit a variety of events, and may be useful to classify their spectral hardness even if the model does not fully describe the data.

In the course of our spectral analysis, we identified fits for which the residuals of the unfolded spectrum for different detectors were inconsistent with each other. This is evidence for a bad localization which we attributed to poor background fits in one or more detectors. We selected background time intervals before and after the source time interval, as is done for GRB localization by the GBM team. We fit the selections with a polynomial (usually a quadratic but occasionally a higher order polynomial is necessary to fit the data) for each detector and for each channel between 0 and 2 (8-50 keV). The event is selected and the fitted backgrounds subtracted. A new localization is performed and the event is labeled as before. Subsequent localizations almost invariably provided an improved localization  $\chi^2$  and smaller error. This was most often due to the previous background fit including the source and reducing the residual rates in each detector in a non-uniform manner thus producing an erroneous localization and resulting in poor detector responses.

Weak events (85) in which spectral analysis was not possible were rejected. With these rejected events and events that were reclassified as solar due to the new localization, there remained 1084 XRB candidates. Figure 1 shows the results of the spectral modeling of these XRB candidates. The top panel is a histogram of the resulting temperatures from blackbody fits while the lower panel is a histogram of the indices from power-law fits. The histograms are fit to a model that consists of multiple gaussians. The power-law distribution is well fit with two gaussians with  $\chi^2 = 85$  with 101 DOF. The blackbody distribution required three gaussians for a fit with  $\chi^2 = 87$  with 56 DOF.

The energy spectra of tXRBs is expected to be a blackbody with temperature between 0.5 and 3.0 keV. Our results for 4U 0614+09 (Linares et al. 2012) and Figure 1 suggest that GBM is sensitive to bursts with temperatures at the high end of this range. The characteristics of event 10032800979 (Table 5), which has been identified as a tXRB from 4U 0614+09 (Linares et al. 2012), was used to demonstrate our ability to recover the temperature using the spectral analysis approach described above. 1000 simulated data sets were created for the NaI detectors N0, N1, N9 and NA for the brightest 4.1 second bin which has an energy flux (10-100 keV) of  $(7.57 \pm 0.33)E^{-8}$  erg s<sup>-1</sup> cm<sup>-2</sup> and another 1000 data sets for a 4.1 second bin in the tail of the burst which has an 10-100 keV energy flux of  $(3.12\pm0.29)E^{-8}$ 

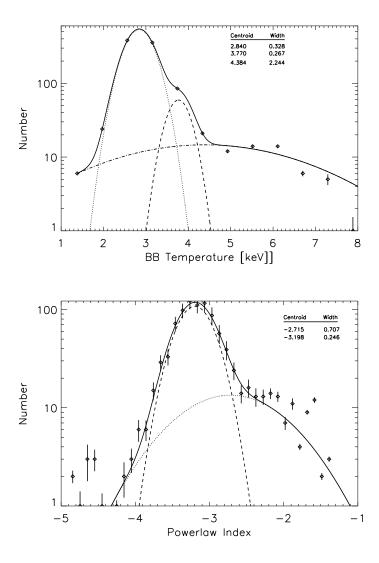


Fig. 1.— Top: Distribution of the temperatures in keV from the blackbody fits to the XRB event spectra. Diamonds are the data points. The solid line is the total model and the dashed and dotted lines are the gaussian components of the total model. Bottom: Distribution of the index from the power-law fits to the XRB event spectra. Diamonds are the data points. The solid line is the total model and the dashed and dotted lines are the gaussian components of the total model.

erg s<sup>-1</sup> cm<sup>-2</sup>. The best fit temperatures for this burst in these time bins are  $3.28 \pm 0.12$  keV and  $3.00 \pm 0.24$  keV respectively. A blackbody spectrum with a temperature of 3.0 keV is used to simulate the data. The simulated data are fit to a blackbody spectrum resulting in the best fit spectral temperature centered on  $3.0 \pm 0.16$  keV for the brightest interval and

 $3.0\pm0.3$  keV for the weak interval assuming the temperatures are normally distributed. The resulting temperature distributions and fits to a gaussian function are shown in Figure 2. These results indicate that any systematic error in the spectral analysis is not dominated by

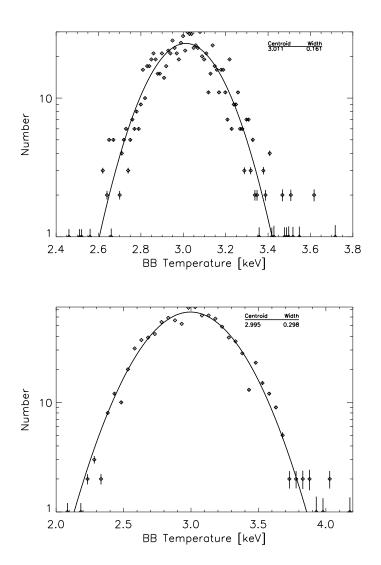


Fig. 2.— Top: Distribution of the temperatures in keV from the blackbody fits to the simulated data from the peak of event 10032800979. Diamonds are the data points. The solid line is the best fit gaussian. Bottom: Distribution of the temperatures in keV from the blackbody fits to the simulated data from the tail of event 10032800979. Diamonds are the data points. The solid line is the best fit gaussian.

the lack of spectral sensitivity in GBM at these energies and fluxes.

We checked if scattering off the Earth's limb was a possible contributor to the systematic

error in the spectral analysis by checking the proximity of the events to the Earth's limb. There were only 23 events that were within 100 seconds of the Earth's limb and only 2 less than 20 seconds. These last two events had a blackbody temperature of  $3.3 \pm 0.2$  keV and  $2.9 \pm 0.2$  keV and were within 10 degrees of the Galactic center. We do not expect such limb events to be a source of systematic error in our catalog.

# 2.4. Temporal Analysis

Temporal analysis of XRB events include the calculation of event duration, rise times, and decay times and was performed after classification (see Section 3) was finished. Due to the nature of the aFXPs (see Section 3.2), these events were excluded from the temporal analysis. Durations for these events are taken from the time interval of the original event selection. Since this analysis requires detailed visual inspection of the light curve, these events underwent additional scrutiny to ensure that aFXPs did not contaminate the remains categories.

For each event where durations are calculated, light curves for all detectors are visually inspected in the 12-25 keV energy band and background regions are selected. The background is fit to a polynomial (usually a quadratic but occasionally a higher order polynomial is necessary to fit the data). The background fit is then subtracted from the light curve. The detectors, in which signal is evident, are selected and the first three energy channels (8-50 keV) are added together and displayed as a single light curve. The peak intensity of the light curve  $(t_{peak})$  is selected. The times at 25% of peak during the rise  $(t_{25})$ , 90% of peak during the rise  $(t_{90})$  and 10% of the peak along the decay  $(t_{10})$  are calculated. As in Galloway et al. (2008), the rise time  $t_{rise}$  is the time for the intensity to rise from 25% to 90% of its peak value, the duration of the event is defined as  $t_{10}$  -  $t_{25}$  and the decay time  $(t_{decay})$  is defined as  $(t_{10} - t_{peak})$ .

Figure 3 shows the duration distribution for the three categories separated by color. The durations for each category show considerable overlap and are not used to distinguish between categories.

### 3. X-ray Burst Catalog

The three year XRB catalog contains 1084 events occurring between MJD 55267 and 56347 (2010 March 12 - 2013 February 24) which are classified into three categories: the tXRBs, the aFXPs, and the uGRBs. Clear distinctions between the three categories is

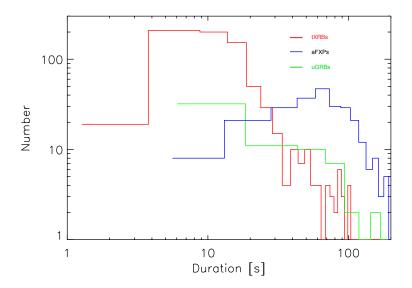


Fig. 3.— Duration distribution for the three categories of events in the catalog. The red curve is the duration distribution for the tXRBs, the green curve is the duration distribution for the uGRBs, and the blue curve is the duration distribution for the aFXPs.

not possible; therefore, we make the following quantitative effort. First, the aFXPs are categorized based on location, visual inspection of the light curve, and spectral analysis (see Section 3.2). Second, the tXRBs are categorized using spectral analysis alone and then the uGRB events are categorized based on spectral analysis and location.

The XRB events are from a wide variety of sources and their spectra is expected to be just as varied. The tXRBs are expected to have a blackbody spectrum (0.5 - 3 keV) while many of the aFXPs and uGRBs are expected to have non-thermal specta which may be modeled, in part, by a power-law. Although the power-law spectral model is generally not a good choice for all three categories of events, it serves well as an indicator of spectral properties for which all categories may be compared. We used spectral results from 32 events that we confidently associate with 4U 0614+09 from this work and Linares et al. (2012) to compare the spectral fit results from a blackbody and power-law model (see Figure 4). There is a tight correlation between the blackbody temperature and the index from a power-law fit justifying our sole use of the power-law in spectral comparisons. This correlation, when considering all events, is tight up to 4 keV (index = -2.5) after which there is considerable scatter in the blackbody temperature.

We choose to be inclusive with our category of tXRBs and use a cut-off in spectral index of -2.5 (4 keV). Any event which is not an aFXP and has a spectral index that is

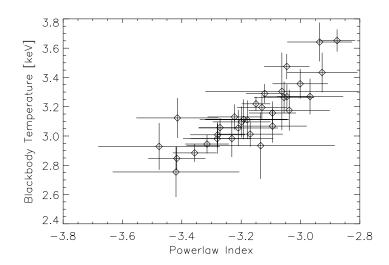


Fig. 4.— Blackbody temperature distribution for the XRB candidates associated by location and spectral shape with 4U 0614+09 show a close correlation with the index of the power-law fits to the same events.

consistent (1 $\sigma$ ) with being softer than -2.5 (< 4 keV) is categorized as a tXRB (See the red distribution in the right panel of Figure 5). This exceeds, by a good margin, the theoretical maximum temperature for type I bursts (Boutloukos et al. (2010)). The left panel of Figure 5 shows the distribution of power-law indices for all XRB events. The softer distribution has a centroid of  $-3.2 \pm 0.25$  and, being the softer distribution, is expected to contain the tXRBs. The spectral index cut-off of -2.5 represents a  $3\sigma$  departure from the centroid thus validating our choice.

The uGRBs are expected to be isotropically distributed across the sky while the tXRBs are mostly at the Galactic center. If we assume GBM uniform exposure, the power-law index cutoff that maximizes the source distribution isotropy can also be used to distinguish these two categories. Using the Rayleigh Test, the maximum isotropy ( $\chi^2 = 1.8/3$  dof) occurs for those events whose power-law index is consistent ( $1\sigma$ ) with being greater than -2.43 thus again validating our choice of -2.5 as a spectral index discriminator between the uGRBs and the tXRBs. Three uGRB events had a spectral index between -2.43 and -2.5 and could arguably be placed in the tXRB classification and they were 10101041428, 11100350666, and 12062078172.

Figure 6 shows the location of all the events in Galactic coordinates with the categories distinguished by color and symbol. The purple diamonds are the tXRBs and there is a large number distributed around the Galactic center which is consistent with the distribution of

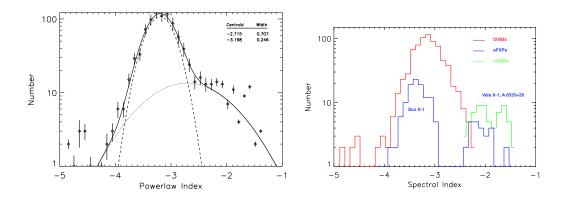


Fig. 5.— Distribution of the spectral index from power-law fits to the XRB candidate spectra. Left: Diamonds are the data points for all the XRB events. The solid line is a model fit to the data and the dashed and dotted lines are the two gaussian components of the total model. Right: Separation of indices by class of event. The red curve is the index distribution for the tXRBs. The blue curve is the index distribution for the aFXPs while the green curve is the index distribution for the uGRBs. Contributions from Sco X-1 and Vela X-1 (both aFXPs) are marked.

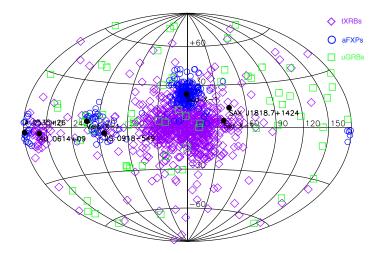


Fig. 6.— Centroids of the localization of all events in Galactic coordinates. The purple diamonds are the locations of the tXRBs. The blue circles are the location of the aFXPs, the green squares are the uGRBs. The error circles for the localization are generally larger than their symbols.

the known type I XRBs. There is a smaller cluster of events consistent with the location of 4U 0614+09. The aFXPs are shown as blue circles and are largely in three clusters centered

around A 0535+26, Vela X-1 and Sco X-1. The green squares are the uGRBs which are distinguished by their isotropic distribution. The classification scheme is summarized in table 1.

Table 1: Source Classification Summary

Category	Number Events	Selection	Properties
	Events	Process	d
aFXPs	267	Location, Visual Inspection	Periodic, Continuous flares
		Spectral: (Sco X-1, $\Gamma < -3$ )	
tXRBs	752	Spectral $(\Gamma < -2.5)$	Galactic; $\overline{kT} = 3.2 \pm 0.3 \text{ keV}$
uGRBs	65	Spectral ( $\Gamma > -2.5$ ); Isotropic	Hard; Extragalactic

#### 3.1. Thermonuclear X-ray Bursts (tXRBs)

The largest category of events in our catalog are soft and their spectra are well fit using a simple blackbody model with temperature in the  $\sim 2-5$  keV range, largely consistent with the spectral properties of thermonuclear bursts from accreting neutron stars (e.g., Swank et al. 1977). They also show a spatial distribution consistent with the  $\sim 100$  known thermonuclear burst sources ("bursters"; see Figure 7), strongly concentrated towards the Galactic bulge region. For the bursts that are bright enough, time resolved spectroscopy reveals cooling along the tail of the burst, the unequivocal signature of tXRBs. All of our tXRBs associated with 4U 0614+09 are bright enough for verification via time resolved spectroscopy, including those reported by Linares et al. (2012). In the most energetic bursts from 2S 0918-549 we also detect cooling along the decay (Section 4.1.2).

We detect in total 752 tXRB candidates with 375 bright enough for time resolved spectral analysis. Their average blackbody temperature is 3.2 +/- 0.3 keV. This value is consistent with the highest temperature measured during photospheric radius expansion (PRE) bursts, when the photosphere is thought to reach the neutron star surface at the end of the Eddington-limited phase (the so-called "touch-down"; Lewin et al. 1993; in't Zand 2005; Kuulkers et al. 2010). The properties of the full tXRBs sample are presented in table 5, including morphology and spectral parameters. Light curves for these events are given in Appendix C, labeled by burst ID.

GBM location errors are typically larger than a few degrees and occasionally tens of degrees (Connaughton et al. 2015). Since the majority of known bursters are within  $\sim$ 20 degrees of the Galactic center, individual identification of tXRBs is limited to those located

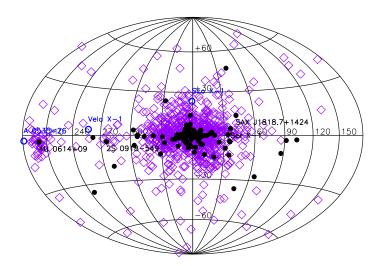


Fig. 7.— The purple diamonds are the locations of the tXRBs. The black filled circles are the locations of the known type I bursters. Those that we have multiple associations for are labeled in black. A few of the aFXP sources are labeled in blue for comparison.

sufficiently far from the Galactic bulge. Due to this limitation intrinsic to the GBM location accuracy, we make no attempt to associate events within this central distribution. Instead, we focus on those sources which are more than 30 degrees from Sag A\*. Out of 103 known bursters, this leaves 26 systems that we attempt to associate with our tXRB events. Furthermore, we use MAXI 2-20 keV weekly light curves (Matsuoka et al. 2009) in an attempt to determine if a given burster was active at the time of the tXRB (see below). We place the 26 bursters far from the Galactic bulge into one of the following four categories.

- If the source is close enough to be detected in MAXI but has not flared within our catalog time period, the source is considered always off and we remove it from consideration. Only Cen X-4 is in this category.
- There are 6 bursters that are below the 10σ detection threshold in MAXI but have always shown persistent emission whenever they have been observed with pointed X-ray detectors. All but one (4U 1323-62, with an orbital period of 2.9 hr) are confirmed or candidate ultra-compact X-ray binaries (UCXBs: orbital periods shorter than 1 hr; see in't Zand et al. 2007): 4U 0513-40 (in the globular cluster NGC 1851), 4U 1246-58, 4U 1915-05 (dipper), 4U 2129+12 (M15-X2 in the globular cluster M15) and 2S 0918-549 (discussed in detail in Sec. 4.1.2). They are assumed to be persistently accreting at a low rate, and considered a candidate for association with all events. Their mass

accretion rates are below 5% of the Eddington limit (in't Zand et al. 2007), which explains the low persistent flux detected by MAXI together with their distances  $\gtrsim$ 5 kpc.

- There are 8 sources whose transient or persistent activity can be monitored with MAXI: when actively accreting they are detected above the 10σ threshold. We only consider these sources as possible associations to our events if the source exceeds 10σ threshold on the week of the event. These sources are 4U 0614+09 (persistent atoll and UCXB candidate; discussed in detail in Sec. 4.1.1), EXO 0748-676 (quasi-persistent transient, in quiescence since 2008), GS 0836-429 (transient, outburst in July 2012), 4U 1254-69 (persistent atoll dipper), Cir X-1 (peculiar atoll/Z), Ser X-1 (persistent), Aql X-1 (canonical atoll transient with typically one or two outbursts per year) and Cyg X-2 (persistent Z source) (see, e.g., Galloway et al. 2008, and references therein). EXO 0748-676 has not shown activity in MAXI during our search period, thus in practice this burster is treated as off.
- The remaining category contains 11 sources with no available MAXI weekly light curves. This category includes some of the so-called "burst-only sources" (Cornelisse et al. 2002b) as well as faint transients in which there is only one known outburst with which the source was discovered. Swift-BAT daily light curves for these sources, when available, do not provide a clear distinction between quiescent and active periods. These sources are: MAXI J1421-613 (outburst in January 2014, i.e., after catalog), UW Crb (peculiar persistently faint "accretion disk corona" source, known since 1990 Hakala et al. 2005), IGR J17062-6143 (persistently faint at <1% of the Eddington luminosity since its discovery in 2006 Degenaar et al. 2013, and references therein), SAX J1818.7+1424, SAX J1324.5-6313 and SAX J2224.9+5421 (Cornelisse et al. 2002b), Swift J185003.2-005627 (faint transient active in May-June 2011), MXB 1906+00, XB 1940-04, XTE J2123-058 and 4U 2129+47. They are considered for association with the tXRBs, even though their activity and mass accretion rate history are often ill constrained.

An association list is generated for each tXRB using the following criteria. If the event location is within  $2\sigma$  of a burster in the association list above, then that event is associated with the source. If more than three sources are associated with an event, then the event has a large location error, and all associations are rejected as spurious. All associations are listed in the table in ascending order of distance (in  $\sigma$  given in parenthesis) from the source. If only one source locates within  $2\sigma$  of an event then it is listed in bold type, and we consider this a robust association. Out of the total of 752 tXRBs, 685 have no associations and 29 have non-unique associations. We find unique associations for 54 tXRBs, with eight known

bursters. For this reduced sample we can assess the origin of the bursts, and their properties are summarized in Table 2.

# 3.2. Accretion Flares and X-ray pulses

Accretion powered events such as those originating from Sco X-1, A0535+26 and Vela X-1 are identified once the their location and spectra are know (see Figure 8). Sco X-1 events have soft emission (PL index < -3) and are generally well localized to Sco X-1's position. These events are usually part of a longer flaring episode that is distinctive in GBM channel 1 (12-25 keV) data. Events from Vela X-1 and A0535+26 are typically part of a chain of

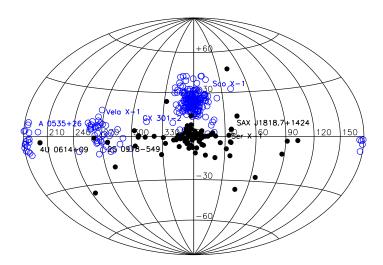


Fig. 8.— The blue circles indicate the locations of the aFXPs. The black filled circles are the locations of the known type I bursters. Those that we have associations for are labeled in black.

pulsations that are identified in the CTIME data due to the dominate harmonic of their characteristic spin periods of 103.3 s and 283.5 s respectfully as well as their harder spectra with a typical power-law index in excess of -2.5. The events associated with A 0535+26 coincide with a giant flare from A 0535+26 which occurred in February 2011 (Camero-Arranz et al. 2011).

The aFXPs are summarized in table 3. The columns are as follows: ID is the time of the midpoint, in UTC, of the event selection identified by YYMMDDTTTTT where YY indicates the last two digits of the year, MM the month, DD the day, and TTTTT is the time in seconds from the start of the day. Peak is the time (UTC) of the peak count rate for

the event measured in seconds since MJD 55267. The RA and Dec is the GBM location and the Error is the statistical error on the location. Association is the source which is associated with the event. The light curves for these events are in Appendix A and identified by ID.

# 3.3. Untriggered GRBs

The uGRBs are hard events that are selected due to their isotropic distributed on the sky which implies an extragalactic origin (see Figure 9). In principle extragalactic bursts could

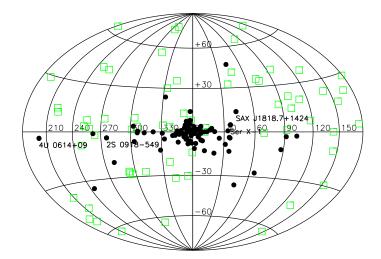


Fig. 9.— The green squares indicate the locations of the uGRBs. The black filled circles are the locations of the known type I bursters. Those that we have multiple associations for are labeled in black.

arise from sources other than GRBs but given the broad range of spectral and temporal behavior exhibited by GRBs, we use the term uGRB to denote our whole extragalactic population of bursters. Their spectra are well fit with a Band function (Band et al. (1993)) or power-law with an exponential cut-off function that is typical of GRBs. Parameters from the spectral fits using the Band function were typically not well constrained and are not reported. The spectral results for the power-law and power-law with exponential cut-off are summarized in Table 4. The first three columns are the same as the aFXPs. The next three columns (4-6) are the results of spectral fitting using a power-law with a exponential cut-off parameterized as Epeak. A '-' in these column denotes that the spectral parameters could not be constrained and these results are left out of the table. The fifth column is Epeak in keV, the sixth column (Comp Flux) is the energy flux [erg cm<sup>-2</sup>s<sup>-1</sup>] from 10-1000 keV, and

the seventh column (Comp Flnc) is the energy fluence [erg cm<sup>-2</sup>] form 10-1000 keV. The next three columns (8-10) are the results of the spectral fitting using a power-law model. The eight column is the power-law index, while the ninth and tenth columns are the energy flux and fluence from 10-100 keV. The last 4 columns are results from the temporal analysis discussed in detail in Section 2.4 and include the rise time (Rise), fall time (Fall), duration (Duration), and a column labeled Structure describing the temporal structure of the event. The Structure column contains an 'S' if the light curve is single peaked or an 'M' if the light curve is multi-peaked. If an event is multi-peaked, the rise time and fall time that is calculated may no longer represent a true rise or fall time for the event since the peak of the event could occur on any of the multiple peaks. The light curves for these events are in Appendix B and identified by ID.

#### 4. Discussion

We have uncovered a large catalog of untriggered bursts in the GBM data that reflect the power of GBM as an all-sky monitor of diverse astrophysical phenomena in the hard X-ray energy band. Despite the difficulties inherent in uncovering these bursts in the background-limited GBM detectors, and the limitations imposed by GBM's coarse source localization, we identified at least three distinct classes of events: untriggered GRBs, accretion-powered flares and X-ray pulsations from known sources, and thermonuclear type I X-ray bursts.

Our source classification relied strongly on spectral modeling and, particularly for the aFXPs, location. Classification from spectral analysis was complicated by the overlapping distributions of spectral parameters among the different classes. We used the spatial distribution of source locations on the sky to verify our choice for the spectral hardness cut-off for the events assigned to the uGRB sample by verifying that the hardness cut-off maximized the isotropy of the spatial distribution.

The tXRBs are the primary science driver for this catalog and we discuss them in depth in Section 4.1. The distribution of temperatures from the blackbody spectral fits of the tXRBs is shown in Figure 10. The temperature distribution has a hard tail that extends beyond 6 keV prompting speculation that there was a fourth, unknown, category of XRBs. Monte Carlo analysis performed in Section 2.3 indicates that GBM has sufficient spectral sensitivity to accurately measure the spectral temperature down to 3 keV yet there are 62 tXRBs whose spectral temperature exceed 4.0 keV and none are associated with a known type I source. Furthermore, these events are distributed along the Galactic plane and concentrated at the Galactic center. A few are weak and may be explained by poor background subtraction while a few may be soft GRBs with chance location along the Galactic plane. We find no

evidence of a bimodal distribution in spectral temperature or fluence thus we conclude that a fourth 'unknown' category is unwarranted with the current data set. We will revisit this when more data has been analyzed.

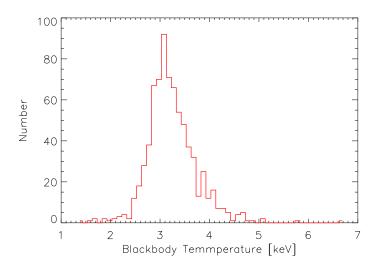


Fig. 10.— The figure shows the distribution of blackbody temperatures for the tXRBs. The distribution is not symmetric and has a hard tail that extends to 6 keV.

The aFXPs were a byproduct of our XRB search since we have dedicated programs to study them (The GBM Pulsar Project<sup>2</sup> and The Earth Occultation Project<sup>3</sup>). Nevertheless; the aFXPs in the catalog provide a unique opportunity to observe these sources in rare, bright states that would normally require a targeted observation.

The brightest (other than the Sun) recurring source GBM observes from 8-50 keV is Sco X-1 and we intentionally attempted to avoid this source since our focus was on tXRBs, nevertheless; Sco X-1 dominates the aFXP category due to its numerous flares. Its persistent nature makes background subtraction difficult and this occasionally leads to poor localization. Luckily, its soft spectrum (index  $\sim -3.5$ ) makes this source relatively easy to identify. The other aFXPs are magnetically dominated accretion powered neutron stars with a harder spectrum (index  $\sim -2$ ). Again, none of the aFXPs were intentionally targeted by our efforts but bright pulsations from these sources occasionally mimic XRBs in the 12-25 keV band and only careful follow-up review of these events reveal the train of pulses that help identify these sources.

<sup>&</sup>lt;sup>2</sup>http://gammaray.msfc.nasa.gov/gbm/science/pulsars.html

<sup>&</sup>lt;sup>3</sup>http://heastro.phys.lsu.edu/gbm

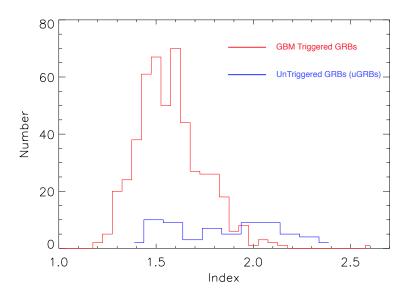


Fig. 11.— The red curve shows the histogram of the spectral index of comparable GBM triggered long GRBs while the blue curve is the histogram of the spectral index of the uGRBs in the XRB catalog

The uGRBs are either GBM sub-threshold trigger events or events which occur when triggering is disabled (rare occurrence). The sub-threshold events are an interesting population of GRBs which might include intrinsically weak, distant, or off-axis GRBs whose detection has consequences for population synthesis studies and future gravitational wave experiments optimized to detect rotating collapsars (Ott et al. 2011) and will be explored in future work. Figure 11 shows the spectral index distribution (in red) of the GBM triggered GRBs during the XRB catalog period whose duration (T90) is greater than four seconds. Overlaid (in blue) on the triggered distribution are the uGRB's spectral index distribution. It is reasonable from the figure to claim that most of the uGRBs are a sub-threshold continuation of the triggered GRB population.

#### 4.1. GBM's view on thermonuclear bursts

With an instantaneous FoV covering 75% of the sky, GBM offers an unprecedented coverage of most Galactic bursters. Due to its sensitivity at energies above ~8 keV, GBM detects only the hottest phases of the hottest type I X-ray bursts: the touch-down phase of PRE bursts (as shown quantitatively in the simulations presented in Linares et al. (2012)) Thus our GBM X-ray burst monitor is a "PRE burst monitor" with an excellent observing duty cycle (50%, only interrupted by Earth occultations and SAA passages).

Figure 14 shows a histogram of energy flux (10–100 keV) for the tXRBs from the black-body spectral fits. The flux distribution was fit ( $\chi^2 = 101/81$  dof) with a gaussian with the centroid at  $3.1 \times 10^{-8}$  erg cm<sup>-2</sup> s<sup>-1</sup> and a standard deviation of  $1.2 \times 10^{-8}$  erg cm<sup>-2</sup> s<sup>-1</sup>. The faintest tXRB in the catalog has a flux of  $(3.4\pm1.0)\times10^{-9}$  erg cm<sup>-2</sup> s<sup>-1</sup>, which gives an estimate of the absolute flux limit in our catalog. Due to the strongly variable X-ray background at 8–50 keV; however, the minimum detectable flux can vary strongly.

PRE bursts reach the Eddington limit, which for a  $1.4 \,\mathrm{M}_{\odot}$  neutron star is in the range  $[1.6\text{--}3.8] \times 10^{38} \,\mathrm{erg \, s^{-1}}$  (depending of the radius and composition of the photosphere; see, e.g., Lewin et al. 1993; Kuulkers et al. 2003). In order to test if the flux distribution is consistent with thermonuclear bursts from the Galactic bulge, we adopt a fiducial Eddington luminosity of  $L_{Edd} = 2.5 \times 10^{38} \,\mathrm{erg \, s^{-1}}$ , and show in Figure 14 the 10–100 keV fluxes corresponding to  $L_{Edd}$  at a distance of 8 kpc and 10 kpc (horizontal lines labelled  $L_{Edd,8}$  and  $L_{Edd,10}$  and). At least three factors contribute to the observed flux scatter: i) bursters have a range of distances, ii) different systems can have different  $L_{Edd}$  (due to differences in neutron star mass, radius or photospheric composition), and iii) even in a given burster the peak luminosity of PRE bursts show significant scatter (Galloway et al. 2008).

We thus conclude, from the flux distribution shown in Figure 14, that our tXRB sample is consistent with a population of Eddington-limited PRE bursts coming from a mix of bursters around the Galactic bulge region. Moreover, because only a handful of tXRBs have fluxes lower than that corresponding to an Eddington-limited burst at 10 kpc (yet several known bursters are farther than that), we estimate that our catalog is limited to PRE bursts occurring within  $\sim$ 10 kpc. The fluence distribution, on the other hand, shows that for an assumed distance of 8 kpc, most bursts have energies between  $10^{39}$  erg and  $10^{40}$  erg (see horizontal lines labelled E39 and E40 in Fig. 14), although the range of fluences is wide with about two orders of magnitude. The duration of the tXRBs in the GBM band also spans a wide range, between  $\sim$ 5 s and  $\sim$ 500 s. The observed distributions of fluence and duration are not bimodal, both in the full tXRB sample and in the two low- $\dot{m}$  bursters presented in Sections 4.1.1 and 4.1.2. This indicates that the longest and most energetic thermonuclear bursts, sometimes referred to as "intermediate/long bursts", are an extreme case of normal burst ignition.

We use hereafter a bolometric correction factor of  $f_{bolo}=1.9$  to convert from 10–100 keV to bolometric burst flux and fluence, which we derive using a typical kT<sub>bb</sub>=3 keV spectrum. Moreover, due to the high background rate and lack of sensitivity below 8 keV, GBM only detects the peak of tXRBs, where the temperature is highest. To take this into account (i.e., to include an estimate of the energy radiated during the burst tail), we use a "band correction factor" of  $f_{band}=1.3$  to convert from GBM fluences (8–50 keV) to a more standard

(2–50 keV) energy band. This band correction was calculated by Linares et al. (2012) using simulated GBM lightcurves of bursts observed with the RXTE-PCA.

Table 2: GBM bursts with associated bursters

Burster	Nr.Bursts	D(kpc)	$L(10^{38} \text{ erg/s})$	$E(10^{39} \text{ erg})$	dur(s)	rise(s)	<kT $>$ (keV)
4U 0614+09	33	$3.2^{a}$	0.3-1.7	0.4-6.1	6.0-51.3	1.2-8.6	3.2
2S 0918-549	10	$5.0^{b}$	0.7 - 1.7	1.0 - 17.0	9.7 - 75.6	2.5 - 38.9	3.1
SAX J1818.7+1424	4	$9.4^{c}$	2.2 - 4.3	8.8 - 25.7	20.1 - 90.9	13.5 - 62.0	3.5
UW Crb	2	$5^d$	1.0-1.0	1.4-1.8	10.6 - 13.7	2.4 - 4.9	3.2
IGR J17062-6143	2	$5^e$	0.9 - 1.0	2.0 - 2.6	16.0 - 22.0	6.9 - 10.2	3.6
XB 1940-04	1	$8^f$	4.4 - 4.4	28.9 - 28.9	50.1 - 50.1	11.8-11.8	3.9
Ser X-1	1	$8.4^{g}$	2.8 - 2.8	19.4 - 19.4	52.6 - 52.6	39.9 - 39.9	3.1
MAXI J1421-613	1	$7^h$	2.0-2.0	13.8-13.8	53.9-53.9	2.3-2.3	4.8

<sup>&</sup>lt;sup>a</sup>(Kuulkers et al. 2010)

# 4.1.1. 4U 0614+09

The burster and UCXB candidate 4U 0614+09 has been extensively studied by most X-ray missions, and is known to accrete persistently at a rate close to 1% of the Eddington limit. Due to its location far from other bursters and its proximity (Kuulkers et al. 2010 measured a distance of d=3.2 kpc, which we adopt in this work), it is an ideal source to study thermonuclear bursts at low accretion rates. During the first year of the Fermi-GBM X-ray burst monitor we detected 15 bursts from 4U 0614+09 (Linares et al. 2012).

Our three year catalog includes 33 tXRBs from 4U 0614+09 detected by GBM between March 2010 and March 2013. This is the same number of bursts detected from 4U 0614+09 with 9 different instruments over the course of 15 years (1992–2007, Kuulkers et al. 2010), which shows the drastic improvement in detection efficiency gained by GBM. Given GBM's 50% observing duty cycle, we measure a burst recurrence time of  $t_{rec}$ =17±2 d (1 $\sigma$  Poissonian

 $<sup>^{</sup>b}$ (in't Zand 2005) 4.0-5.3 kpc

 $<sup>^{</sup>c}$ (Cornelisse et al. 2002b) <9.4 kpc

 $<sup>^{</sup>d}$ (Hakala et al. 2005) >5-7 kpc

<sup>&</sup>lt;sup>e</sup>(Degenaar et al. 2013)

f(Murakami et al. 1983) unknown distance

g(Cornelisse et al. 2002a) 7.7-10.0 kpc

 $<sup>^{</sup>h}$ (Serino et al. 2015) <7 kpc

uncertainty), 5 d longer than, but consistent with, the results of Linares et al. (2012). The closest burst pair we find is only 1.4 d apart, on 2012-06-18/20, the shortest wait time between thermonuclear bursts measured from this source to date. The bolometric and band-corrected burst energies from 4U 0614+09 span more than an order of magnitude, between  $[0.4-6.1]\times10^{39}$  erg, and show no evidence of bimodality, as shown in Table 2 and Figure 12.

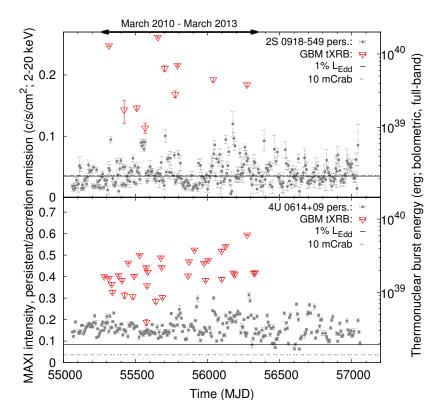


Fig. 12.— Red triangles show the time and radiated energy (right axis scale, band-corrected; see text) of the tXRBs from 4U 0614+09 (bottom) and 2S 0918-549 (top) detected by GBM during the period covered by this catalog (shown with the black arrow). The X-ray intensity is shown in the same panels (gray small circles, left axis), tracing the mass accretion rate history of each burster. The solid and dashed horizontal lines show the corresponding 1%  $L_{Edd}$  and 10 mCrab levels, respectively.

### 4.1.2. 2S 0918-549

Before our GBM campaign, 7 thermonuclear bursts had been reported from the UCXB candidate 2S 0918-549, between 1996 and 2004 (in't Zand 2005, we use the same distance of d=5 kpc throughout this work). This burster is analogous to 4U 0614+09 in many ways:

both are candidate UCXBs persistently accreting at a very low rate, and without detected hydrogen or helium lines in the optical spectrum (Nelemans et al. 2004). The inferred mass accretion rate in 2S 0918-549 is about two times lower,  $\sim 0.5\%$  of the Eddington limit (see Figure 12). In three years, we detect 10 tXRBs from 2S 0918-549, yielding a recurrence time of  $t_{rec}=56\pm12$  d. The closest pair of bursts was detected in August 2011, only  $\sim 16$  d apart.

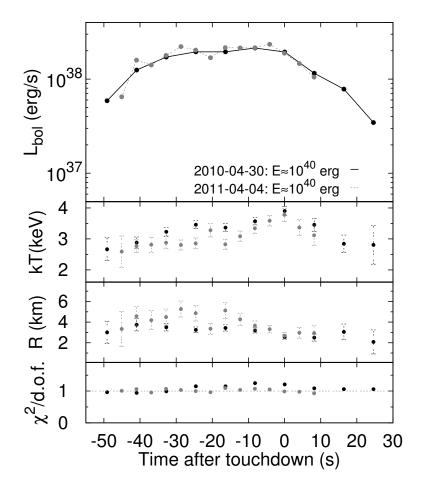


Fig. 13.— Luminosity and spectral evolution of the two most energetic bursts from 2S 0918-549, from top to bottom panels: bolometric luminosity, blackbody temperature, apparent emmiting radius and reduced chi squared. The maximum temperature and highest signal-to-noise in the GBM detectors coincides with the so-called "touch-down", when the neutron star photosphere is thought to reach the surface after expanding and contracting.

Two of the bursts, shown in Figure 13 are consistent with the so-called "long bursts" (in't Zand 2005; Cumming et al. 2006; Chenevez et al. 2008). These were detected on 2010-04-30 and 2011-04-04, with durations in the GBM band of 66 s and 76 s, respectively. They have energies above  $10^{40}$  erg and their duration in the full 2–50 keV band is likely more

than ten times longer than that in the GBM band, i.e., several tens of minutes (the band-corrected duration is more uncertain than the total energy, see Linares et al. 2012, for an estimate). Both the burst durations and energies show a continuous distribution in the range [10-76] s and  $[0.1-1.7]\times10^{40}$  erg, respectively (energies are bolometric and band-corrected, see Sec.3.1).

#### 4.1.3. Other bursters and the integrated Galactic tXRB sample

The remaining associations, 11 tXRBs detected from the direction of six other bursters, are presented in Table 2. Some of these events are faint and have large location errors (Table 5), which together with the low number of tXRBs per burster makes the association uncertain. These include: i) four events from the direction of SAX J1818.7+1424 (detected on 2010-07-01, 2010-07-11, 2010-10-02 and 2011-01-28); ii) two events from the direction of the high Galactic latitude burster UW Crb (on 2011-11-03 and 2011-12-31) and two from the direction of IGR J17062-6143 (on 2010-07-19 and 2011-04-29); iii) one event associated to XB 1940-04 (2011-10-20), one to Ser X-1 (2010-05-31), and one tXRB from the direction of MAXI J1421-613 (on 2011-10-16; note that this source was discovered in outburst in January 2014).

It is also worth discussing which bursters are missing from the association list. Most notoriously, we do not detect any tXRB from 4U 1246–58 in our three-year catalog. In a study of this burster and UCXB candidate accreting persistently below 1% of the Eddington rate, in't Zand et al. (2008) found 7 PRE bursts, all but two with long durations, and a distance of 4.3 kpc. The corresponding burst rate between 1996 and 2008 was  $12\pm6$  d (in't Zand et al. 2007). In contrast, our non-detection of GBM bursts from 4U 1246–58 between 2010 and 2013 implies a 95% lower limit on the recurrence time  $t_{rec}>186$  d, significantly longer than that measured by in't Zand et al. (2007). Thus our results suggest that a drastic change in the burst properties of this burster took place between 2008 and 2010, which might be linked to the long-term decay of its persistent emission already noted in in't Zand et al. (2008).

Two other UCXB bursters are probably too distant to be detected with the GBM X-ray burst monitor: 4U 0513-40 (8.2–11 kpc according to Galloway et al. 2008) and 4U 2129+12 (X-2 in M15, 10.4 kpc away according to Harris 1996). 4U 1915-05 is also close to our detection limit (6.8–8.9 kpc according to Galloway et al. 2008) and is a high inclination "dipper" UCXB, which may explain why no bursts are detected by GBM in the present catalog. The rest of UCXBs and low mass accretion rate bursters are too close to the extended Galactic bulge region to be resolved by GBM, but are included in the total Galactic

rate measured and discussed below.

We show in Figure 14 the distribution of blackbody temperature, flux, fluence and duration, in the full sample of tXRBs. The vast majority of tXRBs come from the Galactic "extended bulge" region (489 locate to within 30° of Sag-A). Fluence and duration are clearly correlated, showing that the most energetic thermonuclear bursts are also the longest, as expected given the physical (Eddington) limit on the burst luminosity.

The total of 752 tXRBs detected in our three-year catalog, correcting for the 50% observing duty cycle, implies a total rate of  $1.37\pm0.04$  thermonuclear bursts per day  $(1-\sigma)$  Poissonian uncertainty). This represents the average over three years of all bursters within the reach of the GBM X-ray burst monitor, which we estimate below corresponds to distances  $\lesssim 10$  kpc (Section 4). Due to GBM's broad sky coverage, this constitutes an unprecedented measurement of the total Galactic thermonuclear burst rate, which we discuss in Section 4.

On average, the GBM bursts in 2S 0918-549 ( $t_{rec}$ =56±12 d;  $\langle E \rangle$ =6×10<sup>39</sup> erg) are more energetic and less frequent than those from 4U 0614+09 ( $t_{rec}$ =17±2 d;  $\langle E \rangle$ =2×10<sup>39</sup> erg). This is qualitatively explained by ignition models, given that 2S 0918-549 accretes at a rate about two times lower than 4U 0614+09 (Cumming et al. 2006). Lower  $\dot{m}$  implies a colder neutron star envelope, a longer fuel accumulation time and a higher ignition depth. However, ignition models still have problems to reproduce quantitatively these recurrence times and burst energies (Kuulkers et al. 2010; Linares et al. 2012). Assuming Solar metallicity, the pure helium ignition models from Cumming and Bildsten (2000) require large amounts of deep crustal heating to reproduce the recurrence times that we measure in 2S 0918-549 and 4U 0614+09: more than 3 MeV per accreted nucleon (see Figure 7 and further discussion in Linares et al. 2012). Having two low- $\dot{m}$  bursters with robust GBM measurements of recurrence times, we can place the first meaningful constraints on the  $t_{rec}$ - $\dot{m}$  relation at  $\dot{m}/\dot{m}_{Edd} \sim 1\%$ . The measured  $t_{rec}$  and  $\dot{m}$  in 2S 0918-549 and 4U 0614+09 are not consistent with a linear relation, and suggest a steeper  $t_{rec} \propto \dot{m}^{[1.7-1.8]}$  relation.

We find a total Galactic rate of 1.4 PRE bursts per day, out to about 10 kpc from the Sun and averaged over the three years of our catalog (Section 4.1.3). During PRE bursts the neutron star atmosphere can be pushed by radiation forces up to hundreds or thousands of kilometers above the surface, and small but significant amounts of nuclear burning ashes may be ejected (Weinberg et al. 2006). To conclude, we roughly estimate the total mass ejected by the PRE bursts uncovered by GBM, by adding their bolometric- and band-corrected fluences and assuming they are all at 8 kpc (see Sec. 3.1 and discussion above). This yields a total radiated energy of  $8\times10^{42}$  erg,  $1.6\times10^{43}$  erg after correcting for the 50% observing duty cycle. For a nuclear energy release of 1.6-4.4 MeV per nucleon, this translates into  $[4-11]\times10^{24}$  gr of burned fuel. For a fraction of ejected mass of  $10^{-4}-10^{-2}$  Weinberg et al.

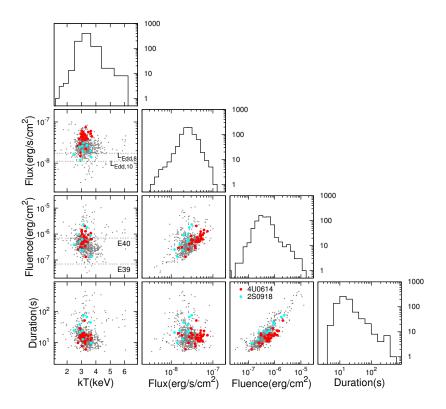


Fig. 14.— Temperature, flux (10–100 keV), fluence (10–100 keV) and duration distribution of the full tXRB sample, as measured by GBM. The bursts from 4U 0614+09 and 2S 0918-549 are shown in red and blue, respectively. Horizontal dashed lines show, for comparison and from top to bottom: GBM flux corresponding to an Eddington luminosity  $L_{Edd}=2.5\times10^{38}$  erg s<sup>-1</sup> at a distance of 8 kpc; GBM fluence corresponding to a bolometric energy of  $10^{40}$  and  $10^{39}$  erg at 8kpc. Histograms of all four parameters are also shown.

(2006), this implies a total of  $[4\times10^{20}-10^{23}~{\rm gr}]$  ejected during three years. With the above assumptions, we are able to place direct observational constraints on the amount of mass ejected into the interstellar medium by PRE bursts in our Galaxy (within 10 kpc of the Sun):  $10^{-13}-10^{-11}~{\rm M}_{\odot}~{\rm yr}^{-1}$ . Whether or not this contributes significantly to the Galactic abundances of any elements (proton-rich isotopes have received particular attention in the context of thermonuclear bursts; see Weinberg et al. 2006, and references therein), remains a subject for future studies.

# Acknowledgements

- M.L. was supported by the Spanish Ministry of Economy and Competitiveness under the grant AYA2013-42627.
- This work was also supported by NASA Fermi-GI grant nr. NNX11AO19G (PI: Linares).
- This research has made use of the MAXI data provided by RIKEN, JAXA and the MAXI team.

Table 3:: GBM Accretion Powered Events

	Table 3:: G	BM Accret	ion Powered	d Events	
ID	Peak	Ra	Dec	Error	Association
	s	degrees	degrees	degrees	_
10033115145	01656753	237.6	-19.0	6.6	ScoX-1
10033122117	01663724	253.6	-10.7	6.4	Sco X-1
10041602529	03026531	239.0	-19.0	1.1	ScoX-1 ScoX-1
10041602579 10041602801	03026532 03026803	$240.2 \\ 250.2$	-12.1 -21.8	$\frac{1.6}{2.1}$	ScoX-1 ScoX-1
10041602801	03034062	245.7	-13.8	2.1	Sco X-1
10041610234	03034236	245.7	-18.5	6.1	Sco X-1
10042668136	03956138	246.9	-12.1	4.0	ScoX-1
10042684127	03972145	248.2	-25.8	3.2	ScoX-1
10051456430	05499632	245.2	-16.0	5.0	Sco X-1
10051644231	05660233	249.9	-4.6	4.2	Sco X-1
10052213157	06147420	244.8	-21.4	1.9	ScoX-1
10052218183	06152598	232.6	-17.1	2.0	ScoX-1
10060237774	07122576	239.8	-17.6	4.5	ScoX-1
10060465574	07323177	244.4	-18.3	2.9	Sco X-1
10061439954 10061538029	08161556 08246031	253.0	-19.5	$\frac{3.6}{7.7}$	ScoX-1 ScoX-1
10061538029	08246031	245.4 $129.4$	-19.6 -47.5	13.8	VelaX-1
10062820324	09351541	168.1	-47.5 -61.9	6.2	GX 301-2
10070233600	-	255.6	19.9	7.9	ScoX-1
10071614574	10900977	149.1	-45.8	17.2	VelaX-1
10071614633	10901035	180.0	-70.6	1.7	GX301-2
10071644045	10930444	240.6	-16.6	1.5	ScoX-1
10072224159	11428961	248.6	-14.1	6.6	ScoX-1
10072266106	11470908	245.9	-13.6	1.2	ScoX-1
10072280879	11485681	224.8	-20.0	10.8	ScoX-1
10072468369	11645971	251.1	-17.7	1.0	ScoX-1
10072476948	11654550	226.8	-7.7	8.5	ScoX-1
10072477354	11654953	249.9	-7.1	3.3	ScoX-1
10072528476	11692478	238.1	-21.2	2.9	ScoX-1
10081808487	13746089	256.9	-33.6	13.8	ScoX-1
10082043388 10082313670	13953819 14183272	$246.0 \\ 248.4$	-17.8 -15.0	$\frac{2.6}{2.0}$	ScoX-1 ScoX-1
10082313070	14184189	240.5	-13.0	5.4	ScoX-1
10082858304	14659906	253.7	-4.0	6.1	ScoX-1
10082030304	14871875	236.6	-13.3	4.8	ScoX-1
10083138505	14899307	245.1	-14.9	4.2	ScoX-1
10090701973	15467575	137.4	-39.1	13.1	VelaX-1
10091451311	16121713	248.7	-12.8	3.8	ScoX-1
10091614419	16257621	244.2	-13.1	3.1	ScoX-1
10091614545	16257747	242.3	-24.4	2.6	ScoX-1
10091833221	16449223	249.7	-21.9	3.5	ScoX-1
10091845340	16461331	248.7	-15.5	5.2	ScoX-1
10091950467	16552864	268.5	-10.7	5.2	ScoX-1
10091962088	16564490	254.8	-19.2	$\frac{5.0}{2.3}$	ScoX-1
10092002644 10092236271	16591447 16797869	144.7 $234.2$	-41.0 -17.7	6.5	VelaX-1 ScoX-1
10092269394	16830997	248.0	-3.4	7.3	ScoX-1
10092282005	16843607	234.4	-16.1	3.4	ScoX-1
10092282489	16844084	246.5	-13.1	2.2	ScoX-1
10092340871	16888874	274.1	-10.6	11.6	ScoX-1
10092451782	16986185	109.4	-32.7	13.1	VelaX-1
10092565936	17086738	141.6	-40.6	6.7	VelaX-1
10092623999	17131202	243.8	-12.0	2.2	ScoX-1
10092753639	17247242	134.8	-55.7	8.5	VelaX-1
10100482935	17881338	238.4	-14.3	6.4	ScoX-1
10100507171	17891973	242.4	-17.7	5.9	ScoX-1
10100548695	17933498	245.9	-21.1	3.3	ScoX-1
10100565184 10100629742	17949987	251.7	-5.4 14.1	6.3	ScoX-1 ScoX-1
10100629742	18000944 18001160	254.4 $242.1$	-14.1 -5.8	$12.5 \\ 8.4$	ScoX-1 ScoX-1
10100629938	18011100	239.7	-3.8 -19.4	$\frac{6.4}{3.4}$	ScoX-1
10100722670	18080272	133.1	-31.5	9.2	VelaX-1
10100722070	18093880	133.3	-35.8	1.1	VelaX-1
10100841167	18185170	144.8	-36.1	5.4	VelaX-1
10100981367	18311770	151.5	-53.0	14.2	GX 301-2
10101606400	18841603	238.9	-23.6	3.1	ScoX-1
10101633706	18868908	243.6	-13.3	4.3	ScoX-1
10103004298	20049100	246.2	-16.3	3.0	ScoX-1
10110554438	20617646	244.9	-16.6	1.1	ScoX-1
10111206222	21174224	242.5	-10.0	5.0	Sco X-1
101111349861	21304264	128.5	-54.4	11.9	Vela X-1
10111373491	21327893	140.1	-40.3	9.0	Vela X-1
10112917548 10120155195	$\begin{array}{c} 22654350 \\ 22864797 \end{array}$	246.0 $130.1$	-11.0 -46.2	2.2 4.6	ScoX-1 VelaX-1
10120100100	22004101	100.1	10.2	2.0	. 01011-1

Table 3:: GBM Accretion Powered Events (continued from previous page)

10120524820   2318022   239.0   -14.6   3.6   ScoX-1     10121110242   2385747   247.4   -16.3   2.1   ScoX-1     10121312945   23859347   245.1   -19.0   3.1   ScoX-1     1012131312945   23859347   245.1   -19.0   3.1   ScoX-1     101213131295   23859347   245.1   -19.0   3.1   ScoX-1     101213131295   23859347   245.1   -20.4   2.4   Sco X-1     10121411613   23944416   243.8   -23.5   2.7   Sco X-1     10121431820   23944416   243.8   -23.5   2.7   Sco X-1     10121431820   23946410   243.8   -23.5   2.7   Sco X-1     10121435821   23968624   244.8   -16.0   4.4   Sco X-1     10121431820   23014004   243.5   -13.7   1.6   Sco X-1     10121662433   24168035   243.4   -16.1   4.2   Sco X-1     10121662433   24168035   243.4   -16.1   4.2   Sco X-1     1012166243   24175266   246.3   -12.3   5.9   Sco X-1     10121783224   22475226   246.3   -12.3   5.9   Sco X-1     10121783224   22475226   246.3   -12.3   5.9   Sco X-1     10121856716   24355118   2427   -18.5   3.7   Sco X-1     10121264742   2430176   239.7   -18.5   3.7   Sco X-1     10122067452   24518655   245.2   -12.0   4.7   Sco X-1     10122067452   24550675   244.2   -25.4   4.7   Sco X-1     10122067452   24550675   244.2   -25.4   4.7   Sco X-1     10122079568   24550771   248.0   -14.4   4.0   Sco X-1     10122109310   24560912   227.9   -16.1   5.8   Sco X-1     10122109310   24560912   227.9   -16.1   5.8   Sco X-1     1012210930   24560912   227.9   -16.1   5.8   Sco X-1     1012210930   24560912   227.9   -16.1   5.8   Sco X-1     10122109310   24560912   227.9   -16.1   5.8   Sco X-1     10122109310   24560912   227.9   -16.1   5.8   Sco X-1     10122109310   24560912   227.9   -16.1   5.8   Sco X-1     1012210930   24560912   247.1   -16.3   3.3   Sco X-1     1012210930   24560912   247.1   -16.3   3.6   Sco X-1     1012210930   24560912   247.1   -16.3   3.6   Sco X-1     1012210930   24560912   247.1   -16.3   3.6   Sco X-1     101210930   24560912   247.1   -16.5   Sco X-1     101210930   24560912   247.1   -16.5   Sco X-1     1012109	ID	Peak	Ra	Dec degrees	Error	Association
10121111024   23884626   228.8   -21.4   7.8   ScoX-1   10121312945   23850347   245.1   -19.0   3.1   ScoX-1   10121312945   23850347   245.1   -19.0   3.1   ScoX-1   10121316959   23861294   147.6   -53.6   6.4   Veln X-1   101213161953   23908355   245.1   -20.4   2.4   Sco X-1   10121406699   23933948   242.3   -17.3   2.9   Sco X-1   1012141613   23944416   243.8   -23.5   2.7   ScoX-1   10121418705   23961507   239.3   -14.7   2.3   ScoX-1   10121481202   24014004   243.5   -13.7   1.6   Sco X-1   1012166342   24171944   242.1   -19.4   4.5   Sco X-1   1012166342   24171944   242.1   -19.4   4.5   Sco X-1   10121783224   24275226   246.3   -12.3   5.9   Sco X-1   1012178324   2427528   246.3   -12.3   5.9   Sco X-1   10121856716   24335118   242.4   -14.8   1.7   ScoX-1   10121856716   24335118   242.4   -14.8   1.7   ScoX-1   10122045130   24496332   252.9   -15.6   2.4   Sco X-1   10122067452   24518655   245.2   -12.0   4.7   Sco X-1   10122067452   24518655   245.2   -12.0   4.7   Sco X-1   10122067552   24530467   244.2   -25.4   4.7   Sco X-1   10122067956   2453047   244.3   -15.1   3.6   ScoX-1   10122108793   24546414   241.7   -16.2   2.6   ScoX-1   10122108793   2454612   247.7   24.3   -15.1   3.6   ScoX-1   10122108793   2454612   247.7   -16.3   3.3   Sco X-1   10122108793   2454612   247.7   -16.3   3.6   ScoX-1   10122108793   2454612   247.7   -16.3   3.8   Sco X-1   10122108793   2	10120524820					ScoX-1
10121310842   23857247   247.4   -16.3   2.1   ScoX-1   10121317896   23864294   147.6   -53.6   6.4   Vela X-1   10121317896   23864294   147.6   -53.6   6.4   Vela X-1   101214361953   2399855   245.1   -20.4   2.4   Sco X-1   10121428705   23934167   239.3   -14.7   2.3   ScoX-1   10121428705   23961507   239.3   -14.7   2.3   ScoX-1   10121428705   23961507   239.3   -14.7   2.3   ScoX-1   10121435822   23968624   244.8   -16.0   4.4   Sco X-1   10121662433   24168035   243.4   -16.1   4.2   Sco X-1   10121662433   24168035   243.4   -16.1   4.2   Sco X-1   10121662432   24168035   243.4   -16.1   4.2   Sco X-1   10121772298   24264301   241.1   -13.9   2.9   Sco X-1   10121783281   24275286   246.9   -15.4   3.2   Sco X-1   10121783281   24275286   246.9   -15.4   3.2   Sco X-1   10121861774   24340176   239.7   -18.5   3.7   Sco X-1   101220573642   2456855   245.2   -12.0   4.7   Sco X-1   10122067452   24518655   245.2   -12.0   4.7   Sco X-1   10122067452   24518655   245.2   -12.0   4.7   Sco X-1   10122097956   24530467   244.2   -25.4   4.7   Sco X-1   10122109310   24546912   227.9   -16.1   5.8   Sco X-1   10122109300   24546912   227.9   -16.1   5.8   Sco X-1   10122109300   24569202   247.1   -16.3   3.3   Sco X-1   10122109300   24567202   247.1   -16.3   3.3   Sco X-1   10122109300   24567202   247.1   -16.3   3.3   Sco X-1   10122109300   24567202   247.1   -16.3   3.6   Sco X-1   10122109300   24567202   247.1   -16.3   3.8   Sco X-1   10122109300   24567202   247.1   -16.3   3.8   Sco X-1   10122109300   24567202   247.1   -16.3   3.8   Sco X-1   10122159600   24597202   247.1   -16.3   3.8   Sco X-1   10122159600   24597202   247.1   -16.3   3.8   Sco X-1   10122159600   24597202   247.1   -16.3   3.8   Sco X-1   1012113980   2658686   248.0   -18.8   5.9   Sco X-1   10101217880   26199433   251.5   -13.6   6.2   Sco X-1   10101217880   26199433   251.5   -13.6   6.2   Sco X-1   10101216767   2828586   243.1   -13.6   6.2   Sco X-1   10101246717   2935179   244.8   -1.7   -1.6   2.4   Sc						
10121312945   23859347   245.1   1-19.0   3.1   Sco.X-1   10121361953   23998355   245.1   -20.4   2.4   Sco.X-1   10121406666   23939498   242.3   -17.3   2.9   Sco.X-1   1012141613   23944416   243.8   -23.5   2.7   Sco.X-1   10121428705   23961507   239.3   -14.7   2.3   Sco.X-1   10121438522   23968624   244.8   -16.0   4.4   Sco.X-1   1012146323   24168035   243.4   -16.1   4.2   Sco.X-1   10121666342   24171944   242.1   -19.4   4.5   Sco.X-1   10121666342   24171944   242.1   -19.4   4.5   Sco.X-1   10121783224   24275226   246.3   -12.3   5.9   Sco.X-1   10121783234   24275226   246.9   -15.4   3.2   Sco.X-1   10121861761   24335118   242.4   -14.8   1.7   Sco.X-1   10121861774   24340176   239.7   -18.5   3.7   Sco.X-1   10122053642   24504844   248.9   -15.2   3.0   Sco.X-1   10122053642   24504844   248.9   -15.2   3.0   Sco.X-1   10122053642   24504844   248.9   -15.2   3.0   Sco.X-1   10122079265   24530467   244.2   -25.4   4.7   Sco.X-1   10122079265   24530467   244.2   -25.4   4.7   Sco.X-1   10122108793   24546142   241.7   -16.2   2.6   Sco.X-1   10122109250   24547627   244.3   -11.5   3.6   Sco.X-1   10122109250   24547627   244.3   -15.1   3.6   Sco.X-1   10122150930   24569220   247.1   -16.3   3.3   Sco.X-1   10122150930   24569220   247.1   -16.3   3.3   Sco.X-1   10122150930   24569202   247.1   -16.3   3.3   Sco.X-1   10122150930   24569202   247.1   -16.3   3.3   Sco.X-1   10122150930   24569202   245.6   -18.3   2.8   Sco.X-1   10122150930   24569202   245.6   -18.3   2.8   Sco.X-1   10122150930   2456920   245.6   -18.3   2.8   Sco.X-1   10121210925   24580484   244.9   -14.6   5.6   Sco.X-1   10121210920   24590493   251.5   -13.6   6.2   Sco.X-1   10121210920   24590493   251.5   -13.6   6.2   Sco.X-1   10121206000   24597020   247.2   -16.1   5.8   Sco.X-1   101216060   24597020   245.6   -18.3   2.8   Sco.X-						
10121317896						
10121416696   23939498   242.3   -17.3   2.9   Sco X-1   1012141613   2394416   243.8   -23.5   2.7   Sco X-1   10121435822   23968624   244.8   -16.0   4.4   Sco X-1   101214815202   24014004   243.5   -13.7   1.6   Sco X-1   10121662433   24168035   243.4   -16.1   4.2   Sco X-1   10121662433   24168035   243.4   -16.1   4.2   Sco X-1   10121662433   24264301   241.1   -13.9   2.9   Sco X-1   10121772298   24264301   241.1   -13.9   2.9   Sco X-1   10121773224   2427526   246.3   -12.3   5.9   Sco X-1   10121783581   24275584   246.9   -15.4   3.2   Sco X-1   10121861774   24340176   239.7   -18.5   3.7   Sco X-1   1012166174   24340176   239.7   -18.5   3.7   Sco X-1   10122045130   24496332   252.9   -15.6   2.4   Sco X-1   10122057452   24518655   245.2   -12.0   4.7   Sco X-1   10122067452   24518655   245.2   -12.0   4.7   Sco X-1   10122067452   24518655   245.2   -12.0   4.7   Sco X-1   10122079568   24530771   248.0   -14.4   4.0   Sco X-1   10122109310   24546912   227.9   -16.1   5.8   Sco X-1   10122109310   24546912   227.9   -16.1   5.8   Sco X-1   10122107950   24597202   247.1   -16.3   3.3   Sco X-1   10122170751   24608364   244.9   -14.6   5.6   Sco X-1   10122252616   24677229   240.8   -17.0   2.0   Sco X-1   101122167676   24597202   247.1   -16.3   3.8   Sco X-1   101122167676   24698364   244.9   -14.6   5.6   Sco X-1   101122167676   24698364   244.9   -14.6   5.6   Sco X-1   10112167676   24698364   244.9   -14.6   5.6   Sco X-1   10112167676   24698364   244.9   -14.6   5.6   Sco X-1   10111240860   26479262   241.1   -16.7   2.4   Sco X-1   10111240860   26479262   241.1   -16.7   2.4   Sco X-1   10111240810   26479262   241.1   -16.7   2.4   Sco X-1   10111240860   26479262   241.1   -16.7   2.4   Sco X-1   10111240860   26479262   241.1   -16.6   7.5   Sco X-1   10111240860   26479262   241.1   -16.7   2.4   Sco X-1   1011240860   26479262   241.1   -16.6   7.5   Sco X-1   10102036764   2893865   242.4   -19.8   5.9   Sco X-1   10102036764   2893865   242.4   -19.7   -1.3   Sco						
10121411613   23944416   243.8   2-23.5   2.7   Sco X-1   10121435822   23968624   244.8   -16.0   4.4   Sco X-1   101214635822   24014004   243.5   -13.7   1.6   Sco X-1   10121666343   24168035   243.4   -16.1   4.2   Sco X-1   10121666342   24171944   242.1   -19.4   4.5   Sco X-1   10121783224   24275226   246.3   -12.3   5.9   Sco X-1   10121783224   24275226   246.3   -12.3   5.9   Sco X-1   10121783234   24275226   246.3   -12.3   5.9   Sco X-1   10121856716   24335118   242.4   -14.8   1.7   Sco X-1   10121861774   24340176   239.7   -18.5   3.7   Sco X-1   10122053642   24504844   248.9   -15.2   3.0   Sco X-1   10122053642   24504844   248.9   -15.2   3.0   Sco X-1   10122079265   24530467   244.2   -25.4   4.7   Sco X-1   10122079265   24530467   244.2   -25.4   4.7   Sco X-1   10122079265   24530467   244.2   -25.4   4.7   Sco X-1   10122108793   24546414   241.7   -16.2   2.6   Sco X-1   10122109310   24546912   227.9   -16.1   5.8   Sco X-1   10122109310   24546912   227.9   -16.1   5.8   Sco X-1   10122159600   24597202   247.1   -16.3   3.3   Sco X-1   10122176751   24608364   244.9   -14.6   5.6   Sco X-1   1012176752   24608364   244.9   -14.6   5.6   Sco X-1   1011219762   2458164   252.4   -19.8   5.9   Sco X-1   1011219762   26458164   252.4   -19.8   5.9   Sco X-1   10111204860   26479262   241.1   -16.7   2.4   Sco X-1   10111204860   26479262   241.1   -16.7   2.4   Sco X-1   10111204860   -10.6   -10	10121361953	23908355	245.1	-20.4	2.4	Sco X-1
10121428705   23961507   239.3   -14.7   2.3   ScoX-1   10121481202   24014004   243.5   -13.7   1.6   Sco X-1   10121662432   24171944   242.1   -19.4   4.5   Sco X-1   10121662432   24171944   242.1   -19.4   4.5   Sco X-1   10121772298   24264301   241.1   -13.9   2.9   Sco X-1   10121783224   24275226   246.3   -12.3   5.9   Sco X-1   10121783324   24275264   246.3   -12.3   5.9   Sco X-1   10121783581   24275584   246.9   -15.4   3.2   Sco X-1   10121861774   24340176   239.7   -18.5   3.7   Sco X-1   10122045130   24496332   252.9   -15.6   2.4   Sco X-1   10122045130   24496332   252.9   -15.6   2.4   Sco X-1   10122067452   24518655   245.2   -12.0   4.7   Sco X-1   10122079568   24530771   248.0   -14.4   4.0   Sco X-1   10122079568   24530771   248.0   -14.4   4.0   Sco X-1   1012210930   24546912   227.9   -16.1   5.8   Sco X-1   1012210930   24546912   227.9   -16.1   5.8   Sco X-1   1012210930   24546912   227.9   -16.1   5.8   Sco X-1   10122159600   24597202   247.1   -16.3   3.3   Sco X-1   10122170751   24608364   244.9   -14.6   5.6   ScoX-1   10122210930   24567229   240.8   -17.0   2.0   ScoX-1   101022079030   24597202   247.1   -16.3   3.3   Sco X-1   10122219030   24568144   241.9   -14.6   5.6   ScoX-1   10122219030   24597202   247.1   -16.3   3.3   Sco X-1   10122170751   24608364   244.9   -14.6   5.6   ScoX-1   10122253216   24677229   240.8   -17.0   2.0   ScoX-1   101029030   26199433   251.5   -13.6   6.2   Sco X-1   10101240860   26479262   241.1   -16.7   2.4   Sco X-1   10101240860   26479262   241.1   -16.7   2.4   Sco X-1   1010138800   26582802   245.8   -16.6   7.1   ScoX-1   1010138800   26582802   245.8   -16.6   7.1   ScoX-1   10101366166   -   140.7   62.7   10.3   Sco X-1   10102036744   28406916   225.6   -15.1   2.8   VelaX-1   10202665284   2868686   127.9   -47.0   6.3   VelaX-1   1020266774   28936569   240.3   -17.5   2.8   VelaX-1   1020266774   2893693   241.6   -15.1   2.8   VelaX-1   1020266774   2893693   241.6   -15.1   2.8   VelaX-1   1020266921   303892	10121406696	23939498	242.3	-17.3	2.9	Sco X-1
10121485822   23968624   244.8   -16.0   4.4   Sco X-1   10121662433   24168035   243.4   -16.1   4.2   Sco X-1   10121666342   2417194   242.1   -19.4   4.5   Sco X-1   10121666342   24171944   242.1   -19.4   4.5   Sco X-1   10121783224   24275226   246.3   -12.3   5.9   Sco X-1   10121783224   24275226   246.3   -12.3   5.9   Sco X-1   10121856716   24335118   242.4   -14.8   1.7   ScoX-1   10121856716   24335118   242.4   -14.8   1.7   ScoX-1   10121857174   24340176   239.7   -18.5   3.7   Sco X-1   10122053642   24504844   248.9   -15.2   3.0   Sco X-1   10122053642   24504844   248.9   -15.2   3.0   Sco X-1   10122079265   24530467   244.2   -25.4   4.7   Sco X-1   10122079265   24530467   244.2   -25.4   4.7   Sco X-1   10122079369   24546414   241.7   -16.2   2.6   ScoX-1   10122108793   24546414   241.7   -16.2   2.6   ScoX-1   10122109301   24546412   227.9   -16.1   5.8   Sco X-1   10122109302   2454627   244.3   -15.1   3.6   ScoX-1   10122159600   24597202   247.1   -16.3   3.3   Sco X-1   10122159600   24597202   247.1   -16.3   3.3   Sco X-1   10122159600   24597202   247.1   -16.3   3.6   ScoX-1   10122253216   24677229   240.8   -17.0   2.0   ScoX-1   10102170751   24608364   244.9   -14.6   5.6   ScoX-1   1010121878   26409303   245.6   -18.3   2.8   Sco X-1   10101219762   26458164   252.4   -19.8   3.0   Sco X-1   10101219762   26458669   238.0   -17.5   2.3   Sco X-1   10102036741   28406916   225.6   -15.1   2.8   VelaX-1   10102036771   28308180   240.3   -17.5   2.3   Sco X-1   10102036771   28308180   240.3   -17.5   2.3   Sco X-1   10102036771   28308180   240.3   -17.5   2.3   Sco X-1   1020205377   28308180   240.3   -17.5   2.8   VelaX-1   1020206754   2935699   92.1   23.8   91.   A 0535+26   1022204940   2982942   80.7   23.8   91.   A 0535+26   1022204940   2982939371   255.6   4.4   8.8   Sco X-1   102217667	10121411613	23944416	243.8	-23.5	2.7	Sco X-1
10121481202	10121428705	23961507	239.3	-14.7	2.3	ScoX-1
10121666343	10121435822	23968624	244.8			
10121666342						
10121772928   24264301   241.1   -13.9   2.9   Sco X-1   10121783381   2427526   246.9   -15.4   3.2   Sco X-1   10121856716   24335118   242.4   -14.8   1.7   ScoX-1   10121856716   24335118   242.4   -14.8   1.7   ScoX-1   10122045130   24496332   252.9   -15.6   2.4   Sco X-1   10122056362   24504844   248.9   -15.2   3.0   Sco X-1   10122079265   24530467   244.2   -25.4   4.7   Sco X-1   10122079265   24530467   244.2   -25.4   4.7   Sco X-1   10122079266   24530771   248.0   -14.4   4.0   Sco X-1   10122108793   24546414   241.7   -16.2   2.6   ScoX-1   10122109301   24546412   22.79   -16.1   5.8   Sco X-1   1012210930   24546412   22.79   -16.1   5.8   Sco X-1   1012210931   24546414   241.7   -16.3   3.3   Sco X-1   10122109755   24547627   244.3   -15.1   3.6   ScoX-1   10122170751   24608364   244.9   -14.6   5.6   ScoX-1   10122253216   24677229   240.8   -17.0   2.0   ScoX-1   10122170751   24608364   244.9   -14.6   5.6   ScoX-1   10102917328   26199433   251.5   -13.6   6.2   Sco X-1   10101917328   2619930   245.6   -18.3   2.8   Sco X-1   10101240860   26479262   241.1   -16.7   2.4   Sco X-1   10111240860   26479262   241.1   -16.7   2.4   Sco X-1   1011138800   26582802   245.8   -16.6   7.1   ScoX-1   101013890   26582802   245.8   -16.6   7.1   ScoX-1   1010135800   26582802   245.8   -16.6   7.1   ScoX-1   1010203666   -						
10121783224   24275226   246.3   -12.3   5.9   Sco X-1   10121856716   24335118   242.4   -14.8   1.7   ScoX-1   10121861774   24340176   239.7   -18.5   3.7   Sco X-1   101220574510   24466332   252.9   -15.6   2.4   Sco X-1   10122057452   24518655   245.2   -12.0   4.7   Sco X-1   10122079265   24530467   244.2   -25.4   4.7   Sco X-1   10122079265   24530467   244.2   -25.4   4.7   Sco X-1   10122079368   24530771   248.0   -14.4   4.0   Sco X-1   10122109393   24546414   241.7   -16.2   2.6   ScoX-1   10122109310   24546912   227.9   -16.1   5.8   Sco X-1   10122109310   24546912   227.9   -16.1   5.8   Sco X-1   101221959600   24597202   247.1   -16.3   3.3   Sco X-1   10122170751   24608364   244.9   -14.6   5.6   ScoX-1   10122223216   24677229   240.8   -17.0   2.0   ScoX-1   1010293223   26192930   245.6   -18.3   2.8   Sco X-1   10101294878   26443281   248.0   -18.9   3.0   Sco X-1   1011219762   26458164   252.4   -19.8   5.9   Sco X-1   10111294878   26443281   248.0   -18.9   3.0   Sco X-1   10111338896   26558699   238.0   -17.5   2.3   Sco X-1   101013626166   -   140.7   -62.7   10.3   Sco X-1   1102036714   28406916   241.7   -10.0   2.5   Sco X-1   1102036714   28406916   225.6   -15.1   2.8   VelaX-1   1102036714   28406916   225.6   -15.1   2.8   VelaX-1   1102036714   28406916   225.6   -15.1   2.8   VelaX-1   11021441319   29330921   245.6   -15.1   2.8   VelaX-1   11021437948   29310191   239.1   -14.1   2.2   Sco X-1   11021431319   29330924   240.7   -21.4   24.8   ScoX-1   1102137394   29821797   82.3   30.5   6.4   A 0535+26   11022075663   30326458   84.3   28.9   6.8   A 0535+26   11022075663   30447438   234.4   -17.0   3.3   ScoX-1   110214778130   29620393   87.5   21.2   4.1   A 0535+26   11022060571   3037599   246.0   -12.6   4.5   ScoX-1   11022060571   3037599   246.0   -12.6   4.5   ScoX-1   1102276563						
10121783581   24275584   246.9   -15.4   3.2   Sco X-1   10121856716   24335118   242.4   -14.8   1.7   ScoX-1   10121861774   24340176   239.7   -18.5   3.7   Sco X-1   10122045130   24496332   252.9   -15.6   2.4   Sco X-1   10122067452   24518655   245.2   -12.0   4.7   Sco X-1   10122067452   24518655   245.2   -12.0   4.7   Sco X-1   10122079568   24530771   248.0   -14.4   4.0   Sco X-1   10122079568   24530771   248.0   -14.4   4.0   Sco X-1   10122108793   24546414   241.7   -16.2   2.6   ScoX-1   1012210025   24547627   244.3   -15.1   3.6   ScoX-1   1012210025   24547627   244.3   -15.1   3.6   ScoX-1   1012210025   24547627   244.3   -15.1   3.6   ScoX-1   10122170751   24608364   244.9   -14.6   5.6   ScoX-1   10122253216   24677229   240.8   -17.0   2.0   ScoX-1   10101204878   26443281   248.0   -18.9   3.0   Sco X-1   1011204878   26443281   248.0   -18.9   3.0   Sco X-1   1011219762   26458164   252.4   -19.8   5.9   Sco X-1   11011338806   26558699   238.0   -17.5   2.3   Sco X-1   11013026166   -   140.7   -62.7   10.3   Sco X-1   11020367714   28406916   240.3   -17.5   2.3   Sco X-1   1102036761481   28355684   248.8   -18.1   4.4   Sco X-1   11020367148   28355684   248.8   -18.1   4.4   Sco X-1   1102036714   28406916   225.6   -15.1   2.8   VelaX-1   1102046174   -   224.6   -9.7   1.3   ScoX-1   1102047979   28212378   255.6   -15.1   2.8   VelaX-1   1102046174   -   224.6   -9.7   1.3   ScoX-1   1102047979   28212378   243.1   -13.0   7.5   ScoX-1   11021441319   29330924   240.7   -21.4   24.8   ScoX-1   11021471380   29626933   87.5   21.2   4.1   A 05355+26   1102204940   29832650   243.1   -13.0   7.5   ScoX-1   1102147139   29330924   240.7   -21.4   24.8   ScoX-1   1102147139   29330924   240.7   -21.4						
10121856716   24335118   242.4   -14.8   1.7   Sco X-1   1012205130   24496332   252.9   -15.6   2.4   Sco X-1   10122053642   24504844   248.9   -15.2   3.0   Sco X-1   10122079265   24530467   244.2   -25.4   4.7   Sco X-1   10122079265   24530467   244.2   -25.4   4.7   Sco X-1   10122079268   24530467   244.2   -25.4   4.7   Sco X-1   1012208793   24546414   241.7   -16.2   2.6   ScoX-1   10122109310   24546912   227.9   -16.1   5.8   Sco X-1   10122109025   245466412   247.1   -16.3   3.3   ScoX-1   10122159600   24597202   247.1   -16.3   3.3   ScoX-1   10122159600   24597202   247.1   -16.3   3.3   ScoX-1   10122253216   24677229   240.8   -17.0   2.0   ScoX-1   1010292030   2619933   251.5   -13.6   6.2   Sco X-1   1010292030   2619933   251.5   -13.6   6.2   Sco X-1   1011240860   26479262   241.1   -16.7   2.4   Sco X-1   10111240860   26459262   241.1   -16.7   2.4   Sco X-1   1011333896   26558699   238.0   -17.5   2.3   Sco X-1   1011512584   26710186   241.7   -10.0   2.5   Sco X-1   101020366714   28202378   253.2   -22.4   4.7   ScoX-1   10102036714   28406916   248.8   -18.1   4.4   Sco X-1   1020366714   28406916   225.6   -15.1   2.8   Sco X-1   102036714   28406916   225.6   -15.1   2.8   Sco X-1   1020145976   28212378   253.2   -22.4   4.7   ScoX-1   102036714   28406916   225.6   -15.1   2.8   Sco X-1   1020146977   28308180   240.3   -17.5   4.7   Sco X-1   102036714   28406916   225.6   -15.1   2.8   Sco X-1   1021213389   29130191   239.1   -14.1   2.2   Sco X-1   10212441319   29330924   240.7   -21.4   24.8   ScoX-1   10212441319   29330924						
10121861774   24340176   239.7   -18.5   3.7   Sco X-1   10122053642   24504844   248.9   -15.2   3.0   Sco X-1   10122067452   24518655   245.2   -12.0   4.7   Sco X-1   10122079265   24530467   244.2   -25.4   4.7   Sco X-1   10122079268   24530771   248.0   -14.4   4.0   Sco X-1   10122108793   24546412   227.9   -16.1   5.8   Sco X-1   1012210025   24546912   227.9   -16.1   5.8   Sco X-1   1012210025   24546912   227.9   -16.1   5.8   Sco X-1   1012210025   24546912   227.9   -16.1   5.8   Sco X-1   1012210025   2454692702   247.1   -16.3   3.3   Sco X-1   10122170751   24608364   244.9   -14.6   5.6   ScoX-1   10122170751   24608364   244.9   -14.6   5.6   ScoX-1   10102917928   26199303   245.6   -18.3   2.8   Sco X-1   1010193728   26199330   245.6   -18.3   2.8   Sco X-1   1010193728   26199330   245.6   -18.3   2.8   Sco X-1   1011240860   26479262   241.1   -16.7   2.4   Sco X-1   10111240860   26479262   241.1   -16.7   2.4   Sco X-1   11011239830   265858699   238.0   -17.5   2.3   Sco X-1   1101135800   26582802   245.8   -16.6   7.1   ScoX-1   1101135800   26582802   245.8   -16.6   7.1   ScoX-1   110102045976   28212378   253.2   -22.4   4.7   Sco X-1   11020360169   28399371   255.6   4.4   8.8   Sco X-1   11021462117   29351719   247.4   -3.5   9.3   Sco X-1   1102143339   29130191   239.1   -14.1   2.2   Sco X-1   11021462117   29351719   247.4   -3.5   9.3   Sco X-1   11021462117   29351719   247.4   -3.5   9.3   Sco X-1   11021462117   2935171						
10122045130						
10122053642   24504844   248.9   -15.2   3.0   Sco X-1   10122079265   24530467   244.2   -25.4   4.7   Sco X-1   10122079568   24530771   248.0   -14.4   4.0   Sco X-1   10122109379   24546912   227.9   -16.1   5.8   Sco X-1   10122109310   24546912   227.9   -16.1   5.8   Sco X-1   10122109310   24546912   227.9   -16.1   5.8   Sco X-1   10122196000   24597202   247.1   -16.3   3.3   Sco X-1   10122170751   24608364   244.9   -14.6   5.6   ScoX-1   10122253216   24677229   240.8   -17.0   2.0   ScoX-1   10122253216   24677229   240.8   -17.0   2.0   ScoX-1   101022253216   24677229   240.8   -17.0   2.0   ScoX-1   10101204878   26443281   248.0   -18.9   3.0   Sco X-1   1011219762   26458164   252.4   -19.8   5.9   Sco X-1   1011219762   2645864   252.4   -19.8   5.9   Sco X-1   1011338806   26558699   238.0   -17.5   2.3   Sco X-1   10115584   26710186   241.7   -10.0   2.5   Sco X-1   10102046166   -2   140.7   -62.7   10.3   Sco X-1   1020366166   -2   2812378   253.2   -22.4   4.7   Sco X-1   1020366184   28355684   248.8   -18.1   4.4   Sco X-1   1020367714   28406916   225.6   -15.1   2.8   VelaX-1   102066524   28663686   247.9   -47.0   6.3   VelaX-1   102066524   28663686   227.9   -47.0   6.3   VelaX-1   1020145976   29330924   240.7   -21.4   24.8   ScoX-1   1021462117   29351719   247.4   -3.5   9.3   Sco X-1   1021778130   29620339   252.0   -18.4   3.2   Sco X-1   1021778130   29620339   252.0   -18.4   3.2   Sco X-1   1021778130   29620339   252.0   -18.4   3.2   Sco X-1   1021207677   29825629   241.0   -15.5   2.8   A0535+26   102200066   3034648   8.4   3.8   -14.5   7.3   ScoX-1   1021677799   29621801   223.2   -11.6   7.5   Sco X-1   1021778130   29620339   35.5   -1.1   -1.6   7.5   Sco X-1   1021778130   29620339   35.5   -1.1   -1.6   7.5   Sco X-1   102160066   3036458   84.3   28.9   6.8   A0535+26   10222004940   29832942   80.7   23.5   7.0   A0535+26   1022204940   29832942   80.7   23.5   7.0   A0535+26   1022203180   30373796   246.0   -12.6   4.5   Sco X-1   102266081   30474						
10122067452						
10122079568						
10122079568						
10122108793						
10122109310   24546912   227.9   -16.1   5.8   Sco X-1   10122159600   2457202   247.1   -16.3   3.3   Sco X-1   10122159600   24597202   247.1   -16.3   3.3   Sco X-1   10122170751   24608364   244.9   -14.6   5.6   ScoX-1   10122253216   24677229   240.8   -17.0   2.0   ScoX-1   1010913728   26192930   245.6   -18.3   2.8   Sco X-1   10101920230   26199433   251.5   -13.6   6.2   Sco X-1   10101920230   26199433   251.5   -13.6   6.2   Sco X-1   1011204878   26443281   248.0   -18.9   3.0   Sco X-1   1011219762   26458164   252.4   -19.8   5.9   Sco X-1   1011240860   26479262   241.1   -16.7   2.4   Sco X-1   1011333896   26558699   238.0   -17.5   2.3   Sco X-1   1011512584   26710186   241.7   -10.0   2.5   Sco X-1   11013026166   -   140.7   -62.7   10.3   Sco X-1   11020255377   28308180   240.3   -17.5   4.7   Sco X-1   11020360169   2839371   255.6   4.4   8.8   Sco X-1   11020360169   28399371   255.6   4.4   8.8   Sco X-1   11020461574   -   224.6   -9.7   1.3   ScoX-1   11020461574   -   224.6   -9.7   1.3   ScoX-1   11021443119   29330924   240.7   -21.4   24.8   ScoX-1   1102142117   29351719   247.4   -3.5   9.3   Sco X-1   1102147299   29621801   223.2   -11.6   7.5   ScoX-1   11021772999   29621801   223.2   -11.6   7.5   ScoX-1   11021771830   29926933   87.5   21.2   4.1   A.0535+26   1102204574   3037168   84.3   28.9   6.8   A.0535+26   11022600561   30326458   84.3   28.9   6.8   A.0535+26   11022600571   30328965   247.2   -14.0   6.3   ScoX-1   110226602571   30338965   74.3   -18.4   2.6   Sco X-1   11022660357   30447438   234.4   -17.0   3.3   ScoX-1   11022660351   30447438   234.4   -17.0   3.3   ScoX-1   11022660351   30447438   234.4   -17.0   3.3   ScoX-1   11022660351   30447438   234.4   -17.0   3.3   ScoX-1   11022660357   30349658   84.3   28.9   6.8   A.0535+26   1102260366   30447438   234.4   -17.0   3.3   ScoX-1   11022660357   30447438   234.4   -17.0   3.3   ScoX-1   11022660357   30447438   234.4   -17.0   3.3   ScoX-1   11022660351   30447438   234.4   -17.0   3.3						
10122159600	10122109310	24546912				
10122170751	10122110025		244.3	-15.1	3.6	ScoX-1
10122253216			247.1		3.3	
11010913728   26192930   245.6   -18.3   2.8   Sco X-1   11010920230   26199433   251.5   -13.6   6.2   Sco X-1   11011219762   26458164   252.4   -19.8   5.9   Sco X-1   11011219762   26458164   252.4   -19.8   5.9   Sco X-1   110112403800   26479262   241.1   -16.7   2.4   Sco X-1   1101133890   265858692   238.0   -17.5   2.3   Sco X-1   11011358000   26582802   245.8   -16.6   7.1   ScoX-1   11011512584   26710186   241.7   -10.0   2.5   Sco X-1   11013026166   -   140.7   -62.7   10.3   Sco X-1   11020255377   28308180   240.3   -17.5   4.7   Sco X-1   11020356481   28355684   248.8   -18.1   4.4   Sco X-1   11020360169   28399371   255.6   4.4   8.8   Sco X-1   11020360140   22866386   127.9   -47.0   6.3   VelaX-1   11020461574   -   224.6   -9.7   1.3   ScoX-1   1102193389   29130191   239.1   -14.1   2.2   Sco X-1   1102141319   29330924   240.7   -21.4   24.8   Sco X-1   1102141319   29330924   240.7   -21.4   24.8   Sco X-1   11021537226   29413229   211.0   -51.5   2.8   A 0535+26   11021771340   29620239   252.0   -18.4   3.2   Sco X-1   11021772139   29621801   223.2   -11.6   7.5   Sco X-1   11021772139   29621801   223.2   -11.6   7.5   Sco X-1   11021772130   29620333   87.5   21.2   4.1   A 0535+26   11022004940   2982569   92.1   23.8   9.1   A 0535+26   11022600056   30326458   84.3   28.9   6.8   A 0535+26   1102260056   30326458   84.3   28.9   6.8   A 0535+26   1102260571   30328965   247.2   -14.0   6.3   Sco X-1   11022610994   30337396   77.3   27.5   2.6   A 0535+26   1102260056   30326458   84.3   28.9   6.8   A 0535+26   1102260056   303464015   84.2   22.6   3.3   A 0535+26   11022603180   30502382   87.8   23.3   6.0   A 0535+26   11022705663   30418465   84.2   22.6   3.3   A 0535+26   11022803180   30502382   87.8   23.3   6.0   A 0535+26   11022803180   30502381   38.4   25.6   8.9   A 0535+26   11022803180   3050						
11010920230   26199433   251.5   -13.6   6.2   Sco X-1   11011204878   26443281   248.0   -18.9   3.0   Sco X-1   11011219762   26458164   252.4   -19.8   5.9   Sco X-1   11011240860   26479262   241.1   -16.7   2.4   Sco X-1   11011333896   26558699   238.0   -17.5   2.3   Sco X-1   11011358000   265582602   245.8   -16.6   7.1   ScoX-1   11011512584   26710186   241.7   -10.0   2.5   Sco X-1   11013026166   -   140.7   -62.7   10.3   Sco X-1   11020145976   28212378   253.2   -22.4   4.7   ScoX-1   11020360169   28399371   255.6   4.4   8.8   Sco X-1   11020360169   28399371   255.6   4.4   8.8   Sco X-1   11020367144   28406916   225.6   -15.1   2.8   VelaX-1   11020461574   -   224.6   -9.7   1.3   ScoX-1   11020981064   28938650   243.1   -13.0   7.5   ScoX-1   11021441319   29330924   240.7   -21.4   24.8   Sco X-1   11021461217   29351719   247.4   -3.5   9.3   Sco X-1   1102177299   29621801   223.2   -11.6   7.5   Sco X-1   1102177299   29621801   223.2   -11.6   7.5   Sco X-1   1102177299   29621801   223.2   -11.6   7.5   Sco X-1   110217778130   29626933   87.5   21.2   4.1   A 0535+26   1102204794   29832942   80.7   23.5   7.0   A 0535+26   1102204794   29832942   80.7   23.5   7.0   A 0535+26   1102204940   29832942   80.7   23.5   7.0   A 0535+26   11022610994   30337396   77.3   27.5   2.6   A 0535+26   11022610994   30337396   77.3   27.5   2.6   A 0535+26   11022751213   30464015   84.2   22.6   3.3   A 0535+26   11022751213   30464015   84.2   22.6   3.3   A 0535+26   11022803186   30502371   89.1   21.7   6.8   A 0535+26   11022803186   30502371   89.1   21.7   6.8   A 0535+26   11022803186   30502371   89.1   21.7   6.8   A 0535+26   11022803186   30502381   83.4   25.6   8.9   A 0535+26   11022803186   30502387   87.8   23.3   10.4   4.7   A 0535+26   11022803186   3						
11011204878						
11011219762						
11011240860   26479262   241.1   -16.7   2.4   Sco X-1   11011333896   26558699   238.0   -17.5   2.3   Sco X-1   11011358000   26582802   245.8   -16.6   7.1   ScoX-1   11011512584   26710186   241.7   -10.0   2.5   Sco X-1   11013026166   -						
11011333896						
11011358000   26582802   245.8   -16.6   7.1   ScoX-1     11011512584   26710186   241.7   -10.0   2.5   Sco X-1     11013026166   -						
11011512584   26710186   241.7   -10.0   2.5   Sco X-1     11013026166   -						
11013026166						
11020145976   28212378   253.2   -22.4   4.7   Sco X-1     1102055377   28308180   240.3   -17.5   4.7   Sco X-1     11020316481   28355684   248.8   -18.1   4.4   Sco X-1     11020360169   28399371   255.6   4.4   8.8   Sco X-1     11020367714   28406916   225.6   -15.1   2.8   VelaX-1     11020461574   -   224.6   -9.7   1.3   ScoX-1     11020665284   28663686   127.9   -47.0   6.3   VelaX-1     11020981064   28938650   243.1   -13.0   7.5   ScoX-1     1102143389   29130191   239.1   -14.1   2.2   Sco X-1     11021441319   29330924   240.7   -21.4   24.8   ScoX-1     11021452172   29351719   247.4   -3.5   9.3   Sco X-1     11021537226   29413229   211.0   -51.5   2.8   A 0535+26     11021771436   29620239   252.0   -18.4   3.2   Sco X-1     11021772999   29621801   223.2   -11.6   7.5   Sco X-1     11021778130   29626933   87.5   21.2   4.1   A 0535+26     1102103454   -   243.8   -14.5   7.3   ScoX-1     11022013794   29821797   82.3   30.5   6.4   A 0535+26     11022024940   29832942   80.7   23.5   7.0   A 0535+26     1102204940   29832942   80.7   23.5   7.0   A 0535+26     1102260056   30326458   84.3   28.9   6.8   A 0535+26     11022602571   30328965   247.2   -14.0   6.3   ScoX-1     11022610994   30337396   77.3   27.5   2.6   A 0535+26     11022602571   30328965   247.2   -14.0   6.3   ScoX-1     11022734635   30447438   234.4   -17.0   3.3   ScoX-1     11022734635   30448465   84.0   25.1   5.4   A 0535+26     11022734635   30447438   234.4   -17.0   3.3   ScoX-1     11022734635   30447438   234.4   -17.0   3.3   ScoX-1     11022803186   30502371   89.1   21.7   6.8   A 0535+26     11022803186   30502383   83.4   25.6   8.9   A 0535+26     11022803186   30502383   83.4   25.6   8.9   A 0535+26     11022803186   30502383   83.4   25.6   8.9   A 0535+26     11022803186   30502381   83.4   25.6   8.9   A 0535+26     11022803186   30502381   83.4   25.6   8.9   A 0535+26     11022803186   30502383   83.4   25.6   8.9   A 0535+26     11022803186   30502371   89.1   21.7   6.8   A 0535+26     11022803		-				
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	11020145976	28212378				ScoX-1
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	11020255377	28308180	240.3	-17.5	4.7	Sco X-1
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	11020316481	28355684	248.8	-18.1	4.4	Sco X-1
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	11020360169	28399371	255.6	4.4	8.8	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$		28406916				
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$\begin{array}{cccccccccccccccccccccccccccccccccccc$		29821797				
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	11022017627	29825629	92.1			A $0535+26$
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$\begin{array}{cccccccccccccccccccccccccccccccccccc$				25.6		
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	11022809694	30508897		27.5		A $0535+26$
11030101205 30586807 251.3 -6.2 4.2 Sco X-1						
11030173543 30659145 85.7 25.2 4.4 A 0535+26						
	11030173543	30659145	85.7	25.2	4.4	A 0535+26

Table 3:: GBM Accretion Powered Events (continued from previous page)

					1 -6 -7
ID	Peak s	Ra degrees	Dec degrees	Error degrees	Association
11030179508	30665111	242.4	-19.8	2.6	Sco X-1
11030181519	30667121	243.5	-10.6	2.0	Sco X-1
11030205018	30677020	235.4	1.8	11.6	Sco X-1
11030215270	30687273	248.0	-5.0	2.7	ScoX-1
1030218823	30690825	239.2	-31.0	2.9	ScoX-1
11030220961	30692963	81.7	28.0	6.8	A $0535+26$
1030274909	30746911	241.9	-14.0	1.2	Sco X-1
1030311905	30770307	240.6	-7.5	3.8	Sco X-1
1030323917	30782319	133.0	-34.1	7.8	Vela X-1
1030342788	30801190	81.1	32.7	5.7	A $0535+26$
1030348871	30807273	82.3	16.6	8.3	A 0535+26
1030383047	30841450	98.2	13.8	21.3	A 0535+26
1030408722	30853548	81.3	25.5	2.9	A 0535+26
11030448088	30892890	81.9	20.8	8.0	A 0535+26
1031382048	31704450	239.1	-21.2	5.8	Sco X-1
1033057363	33148565	135.1	-61.1	20.8	VelaX-1
1042652025	35476028	242.0	-30.0	3.0	ScoX-1
1042652314	35476316	267.3	-8.3	6.4	Sco X-1
1042652484	35476486	247.6	-4.2	3.2	ScoX-1
1042841815	35638617	246.4	-4.2	4.9	ScoX-1
1042869962	35666764	237.7	-24.4	8.5	ScoX-1
1042871462	35668264	236.2	-11.7	6.6	ScoX-1
1050379535	36108337	133.8	-49.4	2.7	VelaX-1
1050748517	36422919	241.4	-13.2	5.0	ScoX-1
1050847498	36508301	249.4	-22.0	3.1	ScoX-1
11051129424	36749427	247.9	-11.7	2.9	ScoX-1
11051135466	36755469	247.5	-10.2	6.4	ScoX-1
11051157743	36777745	245.7	-19.0	2.1	ScoX-1
11051248887	36855289	244.3	-16.2	1.8	ScoX-1
11051270190	36876616	247.3	-15.5	3.0	ScoX-1
1051342842	36935644	239.1	-9.8	1.8	ScoX-1
1052301263	37758066	244.1	-27.1	3.3	ScoX-1
1060753139	39105941	139.8	-41.4	3.0	VelaX-1
11060847684	39186871	247.6	-18.2	2.1	ScoX-1
1061006652	39318654	245.1	-17.2	2.2	ScoX-1
1061200969	39485771	244.9	-15.7	3.5	ScoX-1
1061861410	40064613	242.5	-39.0	7.8	ScoX-1
1062071061	40247063	131.9	-42.6	1.9	VelaX-1
1062165368	40327771	132.5	-36.8	4.0	VelaX-1
1062300765	40435968	248.2	-19.6	21.5	ScoX-1
1070525196	41497199	136.0	-43.9	4.9	VelaX-1
1070667454	41625872	254.4	-17.7	5.1	ScoX-1
1070672799	-	260.9	-12.2	6.9	ScoX-1
1070713658	41658460	235.7	-18.0	4.1	ScoX-1
1070841776	41772978	240.8	-18.4	6.4	ScoX-1
11073150555	43768952	248.6	-9.0	2.0	ScoX-1
11082026427	45472829	239.2	-10.6	2.8	ScoX-1
11082448227	45840241	251.9	-13.2	2.1	ScoX-1
1082711166	46062369	142.1	-49.2	5.4	VelaX-1
1090313504	46669507	256.4	-17.9	4.2	ScoX-1
1090313304	46670801	243.5	-23.0	2.7	ScoX-1
.1090314799	47520795	243.5	-23.0 -15.4	3.4	ScoX-1
1091300793	47669864	255.3	-15.4 -6.7	6.3	ScoX-1
1091463462	48298426		-6.7 -31.5		ScoX-1 ScoX-1
111092200824	48298426	257.4		3.5	ScoX-1 ScoX-1
		243.5	-13.1	2.0	C 37 1
11100215761	49177363	242.6	-17.1	3.6	ScoX-1
11100230702	49192304	229.9	-15.0	5.4	ScoX-1
11100232599	-	243.0	-8.5	2.6	ScoX-1
11100242670	49204272	250.2	-20.8	1.0	ScoX-1
11100244189	49205791	232.5	-4.7	6.0	ScoX-1
1100268246	49229849	240.9	-23.9	3.3	ScoX-1
1101636463	50407665	256.3	-16.0	12.4	ScoX-1
1101636670	50407872	250.2	-2.0	7.4	ScoX-1
11101712910	50470512	239.6	-8.6	6.9	ScoX-1
11101802197	-	246.7	-9.6	9.0	ScoX-1
11101832878	50576880	250.3	-17.1	7.2	ScoX-1
1101833068	50577070	239.8	3.6	7.0	ScoX-1
11101833179	50577181	237.1	-4.3	3.7	ScoX-1
11101908312	50638714	253.1	-19.6	12.8	ScoX-1
1101909692	50640094	238.7	-28.4	9.4	ScoX-1
11101978676	50709079	242.2	-6.9	4.2	ScoX-1
11102078044	50794846	266.7	-6.4	19.6	ScoX-1
11102073044	50860556	158.0	-64.0	16.3	VelaX-1
11102197939	51503441	133.4	-50.5	6.2	VelaX-1
11102909039	51503729	143.0	-38.2	8.7	VelaX-1
	01000128	140.0	-00.2	0.1	v C1421-1

Table 3:: GBM Accretion Powered Events (continued from previous page)

ID	Peak	Ra	Dec	Error	Association
	s	degrees	degrees	degrees	
11102978257	51572660	249.0	-6.5	2.8	ScoX-1
11103170079	51737281	227.2	-22.9	2.5	ScoX-1
11103178586	51745788	228.0	-19.7	1.7	ScoX-1
11110437755	52050557	234.7	-13.3	3.1	ScoX-1
11111232722	52736724	247.7	-8.2	2.0	ScoX-1
11111303703	52794105	150.4	-34.2	1.5	ScoX-1
11111304282	52794118	246.3	-20.2	2.1	ScoX-1
11111410647	52887698	250.6	-17.3	1.0	ScoX-1
11111432972	52909667	253.2	-16.4	2.0	ScoX-1
11111433027	52909768	246.1	-17.1	2.5	ScoX-1
11111625957	53075559	244.5	-17.0	3.9	ScoX-1
11111631676	53081277	234.7	-28.2	9.1	ScoX-1
11112357690	53712093	250.6	-14.2	6.1	ScoX-1
11120462617	54667419	134.9	-35.0	2.9	VelaX-1
11120505963	54697165	129.0	-43.8	8.0	VelaX-1
11120785503	54949505	131.3	-36.0	14.7	VelaX-1
11120933602	55070404	245.8	-14.4	3.7	ScoX-1
11121128322	55237924	248.0	-4.3	5.5	ScoX-1
11122861321	56739723	250.8	-11.9	3.4	ScoX-1
12012975965	59519167	272.0	-16.5	5.2	ScoX-1
12020914747	60408349	247.9	-6.4	7.0	ScoX-1
12052729734	69754537	240.9	-15.5	4.0	ScoX-1
12052801679	69812881	228.9	-10.3	10.3	ScoX-1
12052870492	69881695	235.6	2.3	7.3	ScoX-1
12071250340	73749543	248.9	-11.1	4.8	VelaX-1
12080904347	76122750	247.2	-23.2	10.6	ScoX-1
12082260763	77302365	233.1	-18.5	2.2	ScoX-1
12082941689	77888091	145.9	-43.1	4.9	VelaX-1
12091611473	79413076	242.2	-25.5	5.4	ScoX-1
12091715060	79503045	248.1	-12.0	3.2	ScoX-1
12091727300	79515302	248.3	-14.8	2.8	ScoX-1
12100530195	81073397	137.0	-40.0	3.0	VelaX-1
12100620316	81149918	139.7	-48.8	5.8	VelaX-1
12100818648	81321051	253.5	-16.0	3.0	ScoX-1
12110858521	84039323	137.3	-34.2	3.0	VelaX-1
12110883345	84064147	163.0	-54.4	1.5	VelaX-1
12112714160	85636573	243.0	-26.7	4.2	ScoX-1
12120774013	86560416	120.3	-51.9	27.6	VelaX-1
12122404206	87959408	247.2	-17.6	2.2	ScoX-1
12122851174	88351976	252.9	-37.2	6.4	ScoX-1
13010978715	89416317	253.2	-15.1	1.0	ScoX-1
13021437595	-	247.7	-12.1	4.6	ScoX-1

Table 4:: GBM Untriggered GRB Events

E	Peak	Z Z	Dec.	Error	AsonA	Comp Flux	Comp Fluc	PI. Indev	PI. Flux	PI. Flac	Rise	Fall	Duration	Structure
į	s s	3	3	(sigma)	keV	$10-8 \text{ erg cm}^{-2} s^{-1}$	$10^{-7} \mathrm{erg \ cm^{-2}}$		10 <sup>-8</sup> erg cm <sup>-2</sup> s <sup>-1</sup>	$10^{-7} \text{ erg cm}^{-2}$	sec	sec	sec	
10031206566	00006572	94.6	71.7	6.7	$48148.250000 \pm 7221517.$	$7.17 \pm 0.90$	$20.5 \pm 2.5$	$-1.675 \pm 0.064$	$7.82 \pm 0.85$	$22.3 \pm 2.4$	9.38	10.67	21.34	S
10040364545	01965350	67.2	-13.4	8.6	$63.8 \pm 10.$	$5.09 \pm 0.56$	$14.5 \pm 1.6$	$-1.577 \pm 0.099$	$13.3 \pm 2.6$	$38.2 \pm 7.5$	3.48	21.87	27.25	S
10041528253	02965835	261.6	47.8	2.1	$24.2 \pm 2.0$	$2.28 \pm 0.23$	$17.7 \pm 1.8$	$-2.17 \pm 0.15$	$4.28 \pm 0.96$	$33.2 \pm 7.4$	28.38	95.49	125.40	S
10041565751		272.7	-19.1	4.9	$92.0 \pm 12$ .	$7.99 \pm 0.72$	$22.8 \pm 2.0$	+	$15.3 \pm 1.4$	$43.8 \pm 4.2$	7.65	9.79	19.38	S
10062847601	3 6878789	358.8	69.1	1.0	$125.6\pm7.0$	$15.91 \pm 0.62$	$136.3 \pm 5.3$	$-1.492 \pm 0.020$	$28.55 \pm 0.96$	$244.6 \pm 8.2$	24.75	63.99	101.82	S
10071023190		8.662	-45.0	12.5	$69.9 \pm 8.5$	+	$10.50 \pm 0.97$	$-1.725 \pm 0.070$	$9.4 \pm 1.0$	$19.3 \pm 2.1$	9.29	5.32	15.33	M
10071069939		6.68	-84.0	5.4	$62.6 \pm 11.$	$3.94 \pm 0.44$	$9.6 \pm 1.0$	+	$6.94 \pm 0.89$	$16.9 \pm 2.1$	5.52	8.74	14.66	M
10071369978		290.2	68.3	15.9	$46.5 \pm 15.$	$3.46 \pm 0.63$	$7.0 \pm 1.3$	+	$5.18 \pm 0.99$	$10.5 \pm 2.0$	14.30	7.95	23.36	M
10072541217		313.8	8.92	1.2	$112.7 \pm 18.$	$13.3 \pm 1.0$	$65.4 \pm 5.3$	$-1.664 \pm 0.033$	$21.6 \pm 1.2$	$105.9 \pm 6.1$	17.89	23.18	42.92	M
10081624089		_	37.1	3.7	$50.0 \pm 4.9$	+	$22.5 \pm 1.3$	$-1.910 \pm 0.042$	$10.11 \pm 0.71$	$37.1 \pm 2.6$	25.64	22.73	50.41	M
10081718364	13669576		-22.7	12.9	$33.9 \pm 5.7$	$3.67 \pm 0.52$	$4.49 \pm 0.64$	$-1.94 \pm 0.16$	$3.49 \pm 0.37$	$4.27 \pm 0.45$	6.15	4.65	11.34	S
10090270542		41.6	26.4	9.9	$84.1 \pm 13$ .	$5.43 \pm 0.46$	$33.2 \pm 2.8$	$-1.626 \pm 0.061$	$7.15 \pm 0.74$	$43.8 \pm 4.5$	66.63	21.37	95.72	M
10092473690		184.4	45.0	7.4	$51.3 \pm 15.$	$3.43 \pm 0.50$	$19.6 \pm 2.8$	+	$4.60 \pm 0.54$	$26.2 \pm 3.1$	8.48	63.28	80.96	M
10092635620			56.9	9.4	$153.3 \pm 62.$	$2.82 \pm 0.66$	$23.0 \pm 5.4$	$-1.497 \pm 0.096$	$4.77 \pm 0.71$	$38.9 \pm 5.8$	11.47	8.34	24.78	S
10092955416			-14.8	6.1	$90.3 \pm 6.8$	$7.80 \pm 0.45$	$117.8 \pm 6.8$	$-1.563 \pm 0.037$	$14.86 \pm 0.97$	$224.3 \pm 14.$	71.76	82.78	164.98	S
10100209542	17635152	129.4	-37.7	7.8	$35.5 \pm 15.$	$2.57 \pm 0.77$	$2.10 \pm 0.63$	$-1.95 \pm 0.16$	$4.6\pm1.2$	$3.8 \pm 1.0$	9.78	20.89	32.99	S
10100633055	_	318.7	50.5	3.4	$24.4 \pm 6.8$	$1.96 \pm 0.24$	$16.0 \pm 2.0$	$-2.10 \pm 0.11$	$3.32 \pm 0.53$	$27.1 \pm 4.3$	17.17	74.91	96.94	S
10110363335		238.3	-48.1	3.7	$24.8 \pm 2.0$	+	$21.8 \pm 1.5$	$-2.184 \pm 0.097$	$3.43 \pm 0.44$	$39.3 \pm 5.1$	60.81	57.27	123.74	S
10121121267		245.9	63.2	10.9	$49.3 \pm 20.$	$4.2 \pm 1.1$	$5.2 \pm 1.3$	$-1.83 \pm 0.11$	$3.33 \pm 0.24$	$4.08 \pm 0.30$	1.97	7.45	11.18	S
11011903415	27046624	291.2	55.4	10.6	$26.9 \pm 2.8$	$1.53 \pm 0.20$	$5.01 \pm 0.65$	$-2.09 \pm 0.16$	$2.92 \pm 0.69$	$9.5 \pm 2.2$	9.94	19.61	37.98	ω
11031264584	31600583	154.8	-5.5	2.8	$29.5 \pm 6.6$	$2.29 \pm 0.34$	$11.2\pm1.7$	$-1.97 \pm 0.12$	$4.77 \pm 0.98$	$23.4 \pm 4.8$	9.73	17.64	28.15	S
11031969722	32210533	259.2	45.6	6.6	$46.8 \pm 6.1$	$1.86 \pm 0.26$	$6.84 \pm 0.97$	$-1.64 \pm 0.10$	$5.22 \pm 0.96$	$19.2 \pm 3.5$	12.43	25.86	39.78	S
11041830315	34763063	122.2	30.1	13.5	$169.22 \pm 106.$	$5.9 \pm 1.7$	$13.0 \pm 3.8$	$-1.576 \pm 0.072$	$8.4 \pm 1.1$	$27.6 \pm 3.7$	2.26	70.03	73.11	M
11051735845		282.6	-82.8	14.0	$50.2 \pm 13.$	$8.3 \pm 1.4$	$3.42 \pm 0.60$	$-1.74 \pm 0.10$	$19.3 \pm 3.7$	$7.8\pm1.5$	1.05	5.72	7.06	S
11052339922			46.5	7.5	$134.7 \pm 35.$	$6.39 \pm 0.88$	+	$-1.583 \pm 0.057$	$10.3 \pm 1.0$	$37.9 \pm 3.6$	14.09	16.41	32.16	S
11052552934			-17.1	7.5	$52.1 \pm 29.$	$2.07 \pm 0.55$	$13.5 \pm 3.6$	$-2.10 \pm 0.12$	$2.25 \pm 0.36$	$18.3 \pm 3.0$	11.10	1.89	15.46	M
11062570170	40678179	270.5	-16.9	12.9	ı	I	I	$-2.22 \pm 0.17$	$1.69 \pm 0.25$	$2.76 \pm 0.42$	1.98	11.61	15.04	S
11070345736		221.6	-26.0	13.7	$136.2 \pm 42.$	$16.4 \pm 2.3$	$11.7\pm1.7$	$-1.631 \pm 0.057$	$25.3 \pm 2.3$	$18.1 \pm 1.6$	1.79	3.92	90.9	S
11081315909		129.8	-44.8	4.3	$26.9 \pm 2.7$	$2.68 \pm 0.28$	$3.28 \pm 0.34$	$-2.01 \pm 0.15$	$2.80 \pm 0.28$	$3.44 \pm 0.35$	44.76	23.37	70.78	ω
11081574970		254.9	18.2	2.1	$71.3 \pm 1.9$	$14.56 \pm 0.33$	$172.3 \pm 3.9$	$-1.714 \pm 0.018$	$24.08 \pm 0.68$	$284.9 \pm 8.1$	15.44	70.22	29.66	M
11090373623		99.1	-81.8	5.7	$18.6 \pm 11.$	$6.06 \pm 0.56$	$12.3 \pm 1.1$		$8.29 \pm 0.70$	$16.9 \pm 1.4$	2.33	10.89	13.61	M
11100751308		118.0	0.5	23.7	$89.6 \pm 60$ .	$5.7 \pm 2.3$	$9.3 \pm 3.7$	+	$3.47 \pm 0.30$	$5.67 \pm 0.49$	1.76	4.02	6.07	S
11102935106		51.9	26.7	9.2	$46.0 \pm 4.9$	$5.44 \pm 0.48$	$6.66 \pm 0.59$	+	$8.9 \pm 1.1$	$10.9 \pm 1.4$	2.96	7.44	10.99	S
11121031077		~	41.0	2.0	$101.5 \pm 2.6$	$27.20 \pm 0.46$	$155.4 \pm 2.6$	+ +	$47.10 \pm 0.78$	$269.0 \pm 4.4$	2.28	12.16	59.87	Z i
11122567818	56487021		-52.8	6.4	$218.5 \pm 50.$	$18.5 \pm 1.9$	$22.6 \pm 2.3$	₩.	$25.9 \pm 1.4$	$31.7 \pm 1.8$	80.89	14.31	82.59	M :
12011548725	58282335		-20.5	7.7	$175.7 \pm 31.$	$17.4 \pm 1.6$	$35.6 \pm 3.3$	$-1.497 \pm 0.032$	$26.8 \pm 1.4$	$54.7 \pm 2.9$	39.98	8.87	49.06	M
12012913119	59456316		-10.6	2.2	$202.6 \pm 9.8$	$73.3 \pm 2.1$	$89.6 \pm 2.5$	Н	$111.1 \pm 1.9$	$135.9 \pm 2.4$	2.79	7.69	10.95	S
12021669280		89.9	58.3	1.3	$100.0 \pm 3.5$	$12.91 \pm 0.35$	$179.1 \pm 4.9$	#	$25.18 \pm 0.74$	$349.3 \pm 10.$	76.61	167.50	255.97	M
12032279702		194.3	11.6	6.6	I	I	I	+	$17.1 \pm 3.8$	$6.9 \pm 1.5$	5.39	1.93	9.42	M
12041265725		195.2	14.2	17.0	I	I	I	$-1.83 \pm 0.13$	$8.7 \pm 2.0$	$7.1 \pm 1.6$	1.36	5.49	8.98	M
12042026776	_	106.2	-81.3	11.8	$147.4\pm7.7$	$42.8 \pm 1.6$	$52.3 \pm 1.9$	$-1.387 \pm 0.017$	$79.8 \pm 2.3$	$97.7 \pm 2.9$	0.75	7.24	8.11	M
12090414877		198.7	-6.6	12.5	$101.7 \pm 42.$	$7.9 \pm 1.5$	$9.7 \pm 1.9$	$-1.791 \pm 0.088$	$11.5\pm1.4$	$14.1 \pm 1.8$	6.10	7.18	15.63	S
12092806448	80444854		17.7	13.5	$48.2 \pm 6.1$	$5.47 \pm 0.53$	$4.47 \pm 0.44$	$-1.842 \pm 0.077$	$10.7 \pm 1.3$	$8.7\pm1.1$	1.29	86.6	11.76	M
12102005733	82344942		-10.9	8.9	$80.8 \pm 28.$	+	$12.4 \pm 2.6$	$-1.81 \pm 0.15$	$4.8 \pm 1.0$	$23.6 \pm 5.2$	24.87	6.28	32.05	M
12112083610	85101228	216.4	44.5	6.1	$108.9 \pm 13$ .	$6.17 \pm 0.46$	$35.2 \pm 2.6$	$-1.595 \pm 0.041$	$10.72 \pm 0.74$	$61.2 \pm 4.2$	20.12	29.69	58.55	S
12120241665		112.5	-35.5	21.0	$44.3 \pm 7.4$	+	$18.0 \pm 1.6$	$-2.014 \pm 0.064$	$7.10 \pm 0.69$	$26.1 \pm 2.5$	6.71	21.88	33.37	M
13010754739		126.8	-32.3	2.0	$44.2 \pm 3.2$	$8.98 \pm 0.38$	$47.6 \pm 2.0$	$-2.034 \pm 0.033$	$13.03 \pm 0.61$	$69.1 \pm 3.2$	12.47	20.69	43.06	S
13020154189	91378976	23.5	-2.4	11.5	$20.6 \pm 4.2$	$2.68 \pm 0.30$	$6.57 \pm 0.75$	$-2.23 \pm 0.13$	$4.47 \pm 0.83$	$10.9 \pm 2.0$	10.07	15.33	26.27	M

Table 5:: GBM Type 1 Events

ID	Peak	Ra	Dec	Error	Name (distance)	BB temp	BB flux	BB Flnc	PL index	PL Flux	PL Flnc	Rise	Fall	Duration S	Structure
	s				(sigma)	keV	$10-8 \text{ erg cm}^{-2} \text{ s}^{-1}$	$10^{-7} {\rm erg \ cm^{-2}}$		$10^{-8} \mathrm{erg \ cm^{-2} \ s^{-1}}$	$10^{-7} \text{ erg cm}^{-2}$	sec	sec	sec	
10032800979 (	01383378	98.4 2 8.1	4.6 -26.4	4.6	4U_0614+09	$3.09 \pm 0.10$	$3.21 \pm 0.12$ 0 710 ± 0 057	$7.88 \pm 0.31$ 6 96 $\pm$ 0.56	$-3.198 \pm 0.097$ $-5.73 \pm 0.26$	$6.46 \pm 0.35$ 0 782 ± 0 070	$7.91 \pm 0.43$ $7.66 \pm 0.68$	2.73	11.98	16.71	w w
	01454801	244.9	-46.3	6.5		3.36 ± 0.31	$1.95 \pm 0.22$	$1.59 \pm 0.17$	$-2.76 \pm 0.24$	3.00 ± 0.61	2.45 ± 0.50	1.10	5.87	7.25	o oo
	01620795	246.1	-22.8	2.9		$3.17 \pm 0.12$	$3.47 \pm 0.16$	$32.6 \pm 1.5$	$-3.08 \pm 0.12$	$4.22 \pm 0.33$	$39.6 \pm 3.1$	27.22	98.06	123.39	S
	01706787	236.0	-52.8	3.1		$3.43 \pm 0.15$	$3.39 \pm 0.18$	$4.15\pm0.22$	$-2.90\pm0.13$	+	$5.52 \pm 0.53$	1.52	7.46	11.39	Ø
	01844268	236.3	-54.0	5.0			$2.15 \pm 0.15$	+1 -	# -	# -	$4.54 \pm 0.53$	3.15	5.87	9.82	Ω.
	02535643	270.3	-23.9	13.4			$1.75 \pm 0.27$	$2.14 \pm 0.33$	$-2.40 \pm 0.27$	+ -	$3.7 \pm 1.1$	2.52	7.26	10.29	w ;
	02604627	274.2	-17.0	16.5			$3.48 \pm 0.19$	$5.68 \pm 0.31$	$-3.10 \pm 0.14$	+ -	$7.11 \pm 0.65$	11.02	4.49	16.03	Z;
_	02798228	285.4	-21.8	4.		$3.15 \pm 0.10$	$3.30 \pm 0.13$	$6.74 \pm 0.27$	$-3.09 \pm 0.10$	$4.14 \pm 0.28$	$8.46 \pm 0.58$	2.34	14.02	16.86	Z;
	03096580	267.9	-28.5	3.1			$2.94 \pm 0.15$	$4.80 \pm 0.25$	$-3.34 \pm 0.14$	₩ -	$5.67 \pm 0.47$	13.21	6.34	21.69	M o
	03181438	264.1	-21.1	8. G			$2.28 \pm 0.15$	$9.30 \pm 0.64$	$-3.34 \pm 0.19$	$2.54 \pm 0.28$	$10.3 \pm 1.1$	10.02	22.19	34.72	w o
	03354069	251.4	-27.0	11.3			$1.74 \pm 0.12$	$7.10 \pm 0.52$	$-3.33 \pm 0.19$	$2.09 \pm 0.23$	$8.53 \pm 0.95$	14.32	12.53	31.14	w o
	03360324	256.1	-45.3	15.0		$3.75 \pm 0.27$	$2.39 \pm 0.20$	$2.93 \pm 0.25$	$-2.80 \pm 0.19$	₩.	$4.08 \pm 0.63$	2.47	2.80	10.42	so i
	03508786	281.4	-56.1	8.0			$2.10 \pm 0.19$	$2.57 \pm 0.23$	$-3.05 \pm 0.25$		$3.23 \pm 0.52$	1.43	5.71	7.49	w i
	03629893	269.0	-16.2	27.1			$1.17 \pm 0.13$	$2.39 \pm 0.27$	$-3.57 \pm 0.36$	#	$2.48 \pm 0.45$	2.33	2.75	66.9	S
	04160809	92.6	80 10	6.2	4 U0614 + 09		$4.78 \pm 0.14$	$7.81 \pm 0.24$		+	$10.01 \pm 0.53$	1.63	9.12	11.04	w
	04246220	141.3	-55.5	2.7	$2S_{-}0918-549$		$2.690 \pm 0.063$	$21.97 \pm 0.51$	# -	+1 -	$28.5 \pm 1.2$	38.92	25.87	66.49	Ω.
	04336902	269.0	-41.8	21.7		$4.00 \pm 0.43$	$1.76 \pm 0.22$	+ -	$-2.52 \pm 0.26$	₩.	$2.46 \pm 0.68$	1.30	7.81	9.72	w ¦
	04376509	282.9	-24.5	10.7			+1	+1 -	# -	# -	₩.	13.54	14.71	28.84	M
	04471965	264.6	-19.5	23.9	i i	$3.79 \pm 0.34$		$2.74 \pm 0.30$	$-2.76 \pm 0.26$	H -	$3.52 \pm 0.80$	5.14	6.77	12.72	n ;
0.000027750001	04483923	271.3	2.3	4.0	Ser_A-1 (1.4) S: 4 1185003 2 005637 (	3.59 ⊞ U.15	5.64 ± 0.19	5.93 ± 0.31	$-2.93 \pm 0.13$	$4.54 \pm 0.43$	7.40 ± 0.71	10.31	9.49	21:02	IVI
					SAX_J1818.7+1424 (2.0)										
10050418435	04597637	271.0	-56.0	9.7		$4.36 \pm 0.34$	$3.19 \pm 0.29$	$2.60 \pm 0.24$	2.43	$5.3 \pm 1.1$	$4.34 \pm 0.90$	2.90	7.21	10.71	S
10050454634	04633836	283.3		7.1		$3.19 \pm 0.19$		$4.90 \pm 0.36$	$-3.07 \pm 0.18$		$6.08 \pm 0.76$	11.25	9.40	21.79	M
	04674420	269.2		16.2		$3.26\pm0.42$	$1.40 \pm 0.21$	$2.29 \pm 0.35$	+	$2.04 \pm 0.52$	$3.32 \pm 0.85$	3.70	7.03	11.38	S
10050521228	04686830	248.8	10.5	3.9		$2.570 \pm 0.090$	$8.68 \pm 0.34$	$10.62 \pm 0.42$	$-3.79 \pm 0.14$	$8.02 \pm 0.42$	$13.08 \pm 0.69$	1.92	5.36	15.89	M
10050602786	04754781	278.0	-11.5	12.5			+	$2.41 \pm 0.19$	+	+	$3.08 \pm 0.41$	3.16	5.43	9.03	Ø
10050632732	04784739	8.6		17.0		$2.14 \pm 0.18$	-	$7.17 \pm 0.76$	$-4.42 \pm 0.34$	+	$7.18 \pm 0.89$	6.42	16.38	22.91	w
	05310139	277.2		8.6		$2.97 \pm 0.16$	+	$6.15 \pm 0.42$	#	+	$7.02 \pm 0.78$	12.80	4.29	17.68	M
	05499173	241.0	-15.8	8.7			$3.74 \pm 0.15$	$21.40 \pm 0.87$	$-3.40 \pm 0.12$	+ +	$24.8 \pm 1.4$	155.20	118.79	299.10	M
	05828220	92.0	4.4	8.9	$4 \mathrm{U}0614 + 09$		₩.	$5.12 \pm 0.25$	$-3.17 \pm 0.14$	+1 -	$5.86 \pm 0.40$	1.35	10.59	14.12	w i
	06233262	289.3	-24.0	7.1			$2.42 \pm 0.16$	$5.93 \pm 0.40$	$-3.17 \pm 0.18$	$2.79 \pm 0.32$	$6.84 \pm 0.78$	10.01	5.70	17.58	w ¦
	06401638	278.5	-6.7	11.3			$1.77 \pm 0.10$	$5.80 \pm 0.33$	$-3.18 \pm 0.16$	<b>H</b> -	$6.94 \pm 0.70$	13.22	8.23	22.27	Σ
	06404738	243.1	8.6	1.7			$2.117 \pm 0.062$	$50.1 \pm 1.4$	$-2.980 \pm 0.075$	H -	$64.2 \pm 3.2$	85.70	121.69	225.71	o o
10052518743	06412353	269.9	1.62-	10.6		$2.98 \pm 0.17$	$1.56 \pm 0.11$	$4.45 \pm 0.32$	$-3.25 \pm 0.19$	1.88 ± 0.20	5.37 ± 0.59	7. I3	10.70	20.42	n Ş
	06416527	140.0	1.01			0.24 H 0.10	1 ESG ± 0.073	02.0 H 2.3	ΗН	ΗН	00.0 H 4.3	27.80	110 05	024.00	I N
	06420060	104.9	1.0.1	1 -	411 0614 1 00		НН	09:0 H 1:7	H 700 C	НН	40.0 ± 3.0	1 22	7 500	0.01	Į,
	06704513	218.8	1.54	6.7	50-F	$4.69 \pm 0.31$	$0.915 \pm 0.073$	13.4 + 1.0	+ +	+	22.0 + 4.1	22.90	49.07	92.90	o w
	06928805	284.7	20.1	8.5	SAX_J1818.7+1424 (1.3)	$3.29 \pm 0.20$	+	$7.35\pm0.55$	+	+	$9.4\pm1.2$	13.08	39.91	52.34	M
					Ser_X-1 (1.9)						-				i
	06956268	277.6	9.0	ος 10:	1		+ +	+ -	н.	$2.20 \pm 0.18$	$9.01 \pm 0.73$	24.71	25.24	51.27	w :
10053151022	06963042	287.1	6.8	4 t 2 n	Ser_X-1	$3.09 \pm 0.14$	$1.83 \pm 0.10$	$8.22 \pm 0.47$	$-3.19 \pm 0.15$	$2.28 \pm 0.19$	$10.26 \pm 0.88$	39.99	10.12	52.65	w c
	07684911	0.42	0.07-	. o		0.00 H 0.10	Н-	Н -	0.12 0.00 0.00 0.00	Н -	0.00 H 0.00	10.01	14.20	16.05	υ <u>Σ</u>
	07896841	202.6	5.53	21.7		$3.030 \pm 0.037$	+ +	0.83 ± 0.23	2.00 H	+	3.00 ± 0.44 1.90 + 0.63	2.75	5.06	10.90 8.33	Z V
	07970037	280.0	-27.7	16.2		2.82 + 0.18	+ +	+	+ 68.8	+	$06.0 \pm 92.7$	2.06	21.62	23.88	2 ≥
	08235258	261.7	-32.4	9.1		$3.10 \pm 0.16$	$2.59 \pm 0.15$	1 #1	3.20 ±	1 +1	$5.17 \pm 0.49$	7.67	11.66	21.47	ß
_	08258245	280.2	-17.0	11.6			+	+	3.28 ±	+	$5.70 \pm 0.48$	13.11	3.78	17.48	M
10061611550	08305951	261.3	-31.0	6.2		$3.32 \pm 0.20$	$2.44 \pm 0.17$	+	$-3.01 \pm 0.18$	+	$3.87 \pm 0.49$	5.71	4.60	10.86	Ø
	08756809	251.6	-45.0	19.5		0	+	+	+	$^{H}$	$2.83 \pm 0.88$	3.93	1.85	6.57	S
	08923014	275.3	-30.5	24.0			$1.93 \pm 0.23$	$1.58 \pm 0.19$	+	+	$1.90 \pm 0.41$	1.33	4.31	5.88	Ω
	08931699	285.3	3.4	∞		$3.07 \pm 0.17$	$1.449 \pm 0.095$	$5.32 \pm 0.35$	+1 -	+1 -	$7.54 \pm 0.64$	20.29	33.76	79.56	w :
	08943185	284.6	-59.0	12.7			H -	$10.4 \pm 1.0$	$-3.41 \pm 0.33$	H -	$12.3 \pm 1.7$	26.45	10.89	52.51	N O
10062424054	08951870	276 F	16.5	10.1		2.74 ± 0.13 3.73 ± 0.95	1.78 ± 0.10 9 of ± 0.33	$5.10 \pm 0.31$	$-3.40 \pm 0.17$ $-2.96 \pm 0.20$	2.13 ± 0.19	$6.09 \pm 0.54$ $3.05 \pm 0.41$	12.98	17.54 5.93	31.49	מ מ
	030000000	5		1		7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7	Н	7 17 7	Н	Н	4 20.0	i i	34.	01.01	ù

Table 5:: GBM Type 1 Events continued from previous page

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ID	Peak	Ка	Dec	Error	Name(distance) (sigma)	$_{ m keV}$	$^{10-8} erg cm^{-2} s^{-1}$	$_{10^{-7}\mathrm{erg~cm}^{-2}}$	PL index	$^{ m PL~Flux}_{10^{-8}{ m erg~cm}^{-2}~{ m s}^{-1}}$	$^{1}$ PL Flnc $^{10-7}$ erg cm $^{-2}$	Rise	Fall	Duration Sec	Structure
10062541316	09113317	256.6	-23.1	6.6		$3.53 \pm 0.24$	$2.97 \pm 0.23$	$2.42 \pm 0.19$	$-2.89 \pm 0.18$	$4.08 \pm 0.54$	3.33 ± 0.44	4.37	8.18	13.74	ω
			-38.8	9.6		$3.27 \pm 0.21$	$0.395 \pm 0.080$	+	$-3.03 \pm 0.19$	+	$1.90 \pm 0.35$	4.32	6.57	13.14	S
10062844993	09376201		-46.2	7.4		$4.13 \pm 0.24$	$3.26 \pm 0.23$	$3.99 \pm 0.28$	$-2.51\pm0.14$	$5.50 \pm 0.80$	$6.72 \pm 0.97$	5.41	89.9	12.59	W
10062927638	09445249	261.7	-28.9	6.6		$3.45 \pm 0.27$	$3.00 \pm 0.28$	$2.45 \pm 0.23$	$-3.08 \pm 0.25$	$3.56 \pm 0.57$	$2.90 \pm 0.46$	2.56	5.12	8.34	S
			-12.5	18.2		$4.13 \pm 0.28$	$3.31 \pm 0.26$	$2.70 \pm 0.21$	$-2.78\pm0.19$	+	$3.55\pm0.53$	2.86	3.99	7.45	S
			-25.7	10.3			$2.91 \pm 0.23$	$2.38 \pm 0.19$	$-2.78 \pm 0.19$	₩.	$3.40 \pm 0.53$	1.24	6.43	7.95	SO :
			.38.3	23.2			$4.96 \pm 0.40$	$6.07 \pm 0.49$	$-3.81 \pm 0.25$	+ -	$6.50 \pm 0.63$	3.69	5.24	10.80	w t
		264.5	5.5	7.5	$SAX_J1818.7 + 1424$		₩ -	$10.75 \pm 0.94$	$-2.64 \pm 0.18$	H -	$13.6 \pm 1.4$	62.06	19.82	90.96	o o
		265.2	8.1.	10.7			$1.70 \pm 0.14$	$3.47 \pm 0.30$	$-2.85 \pm 0.21$	$2.24 \pm 0.35$	$4.58 \pm 0.72$	4.63	5.18	10.53	o c
		250.6	-23.5	11.0		$3.52 \pm 0.19$	$3.54 \pm 0.22$	2.89 ± 0.18	$-3.10 \pm 0.17$		$3.43 \pm 0.35$	4.53	5.21	10.47	n ;
			-10.0	11.3			$3.27 \pm 0.17$	5.35 ± 0.28	$-2.92 \pm 0.13$	H -	$7.22 \pm 0.71$	2.78 2.78	13.13	14.70	M
		_	1.02-	0.0	00 - 4 100 114		Z.11 ± 0.15	$4.32 \pm 0.32$	$-3.21 \pm 0.21$	H -	5.39 ± 0.60	00.00 10.00	4.00	13.20	Ωα
		4.78	2.5	0.0	40-0614+09		$4.48 \pm 0.14$	$9.15 \pm 0.29$	$-3.222 \pm 0.090$	H -	$11.05 \pm 0.58$	2.59	9.73	12.76	n o
10070910049	10001238	200.0	0.0	0.0		4.21 H 0.39	2.40 H 0.24	1.90 H 0.20	-2.01 H 0.23	3.04 H 0.73	2.90 H 0.01	1. / S	99.91	10.7	0 0
		0.007	8.4°-	o. 4		$3.12 \pm 0.20$ $4.07 \pm 0.35$	1.40 ± 0.11	8.00 H 0.73	$-3.11 \pm 0.20$ $-2.48 \pm 0.20$	1.80 ± 0.23	13.6 + 2.9	93.33	30.79	58 73	מ מ
10070842735			-21.0	12.9			0.99 + 0.25	$0.81 \pm 0.20$	+ +	+	1.18 + 0.30	1.45	89.9	2 8 8	o v
10070882227			-28.9	12.8			$1.58 \pm 0.17$	$2.57 \pm 0.29$	$-3.03 \pm 0.26$	+	$3.66 \pm 0.64$	1.42	11.55	15.23	o vo
		278.5	3.5	8.9		0	$1.43 \pm 0.10$	$5.25 \pm 0.37$	$-3.50 \pm 0.20$	1	$6.24 \pm 0.62$	39.60	58.69	102.53	S
10071070739	10438736		-18.8	14.0			$1.80 \pm 0.18$	$2.20 \pm 0.22$	$-2.83 \pm 0.23$	+	$3.15 \pm 0.54$	2.64	8.76	12.04	Ø
10071082848	10450865		36.4	6.4			$0.886 \pm 0.084$	$13.0 \pm 1.2$	$-3.32 \pm 0.26$	+	$16.0 \pm 2.4$	16.30	37.02	59.53	W
		262.4	14.9	9.1	SAX_J1818.7+1424		$2.16 \pm 0.12$	$10.61 \pm 0.62$	-11	+	$9.29 \pm 0.81$	22.63	15.86	41.97	S
		289.6	8.3	6.9			$2.17 \pm 0.14$	$5.31 \pm 0.35$	$-3.36 \pm 0.18$	$2.56 \pm 0.26$	$6.29 \pm 0.64$	7.60	13.65	29.62	S
			13.8	8.9			$2.63 \pm 0.12$	$4.30 \pm 0.20$	$-3.13 \pm 0.13$	#	$5.52 \pm 0.44$	7.00	4.18	11.81	ß
			-10.5	11.1			$1.96 \pm 0.16$	$4.00 \pm 0.33$	$-3.20 \pm 0.22$	#	$4.79 \pm 0.64$	12.92	12.80	26.19	M
10071316905		303.2	10.2	10.5		0	$1.700 \pm 0.091$	$9.71 \pm 0.52$	$-3.59 \pm 0.16$	# -	$11.27 \pm 0.87$	72.23	14.03	87.66	M
		254.9	-22.7	11.3			$0.55 \pm 0.12$	$0.45 \pm 0.10$	$-2.91 \pm 0.22$	+	$0.95 \pm 0.21$	2.51	4.00	96.9	w
		291.6	15.1	13.1		o i	₩.	$3.22 \pm 0.26$	₩.	# -	$4.04 \pm 0.61$	2.93	5.76	10.79	w :
		253.6	-47.7	12.5		$3.05 \pm 0.23$	$1.90 \pm 0.17$	$2.33 \pm 0.21$		₩.	$2.62 \pm 0.38$	3.19	2.84	7.23	w ¦
			34.6	2.0		0 0	₩ -	$71.3 \pm 2.6$	$-2.869 \pm 0.093$	₩ -	$92.7 \pm 6.5$	263.74	61.87	336.22	Σ;
			-38.0	19.6				H -	-3.29 ± 0.16	H -	7.23 ± 0.68	12.32	11.23	26.51	M
			1.00	13.0 V		$3.27 \pm 0.22$	2.39 ± 0.21	7 8	$-3.03 \pm 0.20$	Н-	77.0 = 0.7	7.7	7.73	10.31	Ωα
		202.2	-33.0	4. 0	11 0100 - 47 (0.0)		H -	Н-	-3.08 ± 0.20		9.10 ± 0.00	12.01	16.40	28.03	Ωα
10071845074	11104273	317.4	43.5	D. O	40.2129 + 47 (0.8) Cvg X-2 (1.2)	4.58 ± 0.44	$2.42 \pm 0.28$	9.	Н	4.5 ± 1.1	5.0 ± 1.3	3.05	18.42	22.05	Ω
10071919633	11165229	216.7	-53.8	7.5	()	$3.35\pm0.25$	$3.79 \pm 0.32$	+	$-3.07 \pm 0.25$	$4.32 \pm 0.68$	$3.52 \pm 0.55$	1.45	4.03	69.2	w
10071983763	11229370	264.1	-73.7	7.4	IGR_J17062-6143	$4.07 \pm 0.42$	$1.71 \pm 0.20$	+	$-2.67 \pm 0.27$	$2.45 \pm 0.61$	+	6.91	8.30	16.02	M
			5.6	11.6		$3.05 \pm 0.21$	+	$5.76 \pm 0.47$	+	+	+	2.61	3.04	7.82	S
			-18.2	13.1		$3.28 \pm 0.20$	+1 -	+1 -	⊕ .0	+1 -		11.88	17.63	31.32	M
10072103225	11321629	334.5	11.1	11.8	4U_2129+12 (1.0)	$1.98 \pm 0.12$	$1.55 \pm 0.11$	$7.62 \pm 0.58$	$-4.52\pm0.27$	$1.60 \pm 0.15$	$7.85 \pm 0.77$	11.28	20.96	35.32	M
10072175731	11394122	267.3	-37.8	11.1	(0:1) 000-07176-1110	$3.39 \pm 0.21$	$2.69 \pm 0.19$	$2.19 \pm 0.15$	$-3.21 \pm 0.20$	$3.20 \pm 0.35$	$2.61 \pm 0.29$	2.56	7.07	10.22	Ø
			-21.0	8.4		O	1	1			1	2.88	80.9	10.10	S.
10072282320	11487113	254.2	-20.8	12.4		$3.59 \pm 0.13$	$4.47 \pm 0.19$	+	$-2.95 \pm 0.10$	$5.68 \pm 0.42$	$11.59 \pm 0.86$	3.84	13.74	18.26	M
		253.8	-25.7	11.3		$3.67 \pm 0.25$	-	$3.35 \pm 0.25$	$-2.85 \pm 0.19$		$4.55 \pm 0.65$	4.55	6.18	11.22	S
			-22.0	8.5			$2.56 \pm 0.12$	$5.23 \pm 0.26$	$-3.25 \pm 0.14$	+1 -	$6.26 \pm 0.50$	17.91	7.60	26.54	w i
			-30.1	20.3		$4.22 \pm 0.48$	$2.37 \pm 0.31$	$1.93 \pm 0.25$	$-2.71 \pm 0.30$	₩.	$2.60 \pm 0.68$	5.50	4.53	11.32	so i
			-32.4	13.1			$2.57 \pm 0.19$	$2.09 \pm 0.15$	3.18	₩ -	$2.56 \pm 0.32$	2.35	5.07	9.71	so o
			-34.2	7.5		$3.80 \pm 0.34$	$2.28 \pm 0.23$	$1.86 \pm 0.19$	$-2.90 \pm 0.27$	<b>H</b> -	$2.04 \pm 0.28$	1.89	4.02	6.25	w c
	_	900	19.7			<u> </u>	н-	16.98 # 0.53	$-2.440 \pm 0.060$	Н-	22.19 ± 0.77	30.84	25.77	59.37	n ;
10072670238	11820631	282.9	9.9	- x		$3.22 \pm 0.14$	$1.392 \pm 0.078$	18.1 ± 1.0	$-3.02 \pm 0.13$	$6.70 \pm 0.10$	$19.2 \pm 1.4$ 8 2 $\pm$ 1 1	12.75	34.34	86.74	WI O
			2.0-	o ro d t		2.01 H 0.29	Н Н	0.1 ± 90.0	-5.75 H 0.36	H +	0.0 ± 1.1	1 - 1	2 1 2	10.20	ט מ
	_		-39.3	11.4		3.35 ± 0.22	1 +1	1 +	1 +	1 +	$2.74 \pm 0.34$	2.71	4.35	7.92	ω
			6.1	9.4	$4U_{-}0614+09$			$3.62 \pm 0.15$	3.09	+	$4.77 \pm 0.34$	1.44	9.42	11.20	w
10073104971	12187376	276.1	-49.1	19.2		$4.13 \pm 0.31$	$4.74 \pm 0.42$	$1.93 \pm 0.17$	$-2.69 \pm 0.19$	+	+	1.71	11.76	13.87	Ø
10080339189	12480794	279.5	-17.4	9.6		$2.60 \pm 0.10$	$2.57 \pm 0.12$	$6.31 \pm 0.29$	$-3.71\pm0.14$	$2.80 \pm 0.18$	$6.87 \pm 0.45$	12.82	12.26	26.46	W

Table 5:: GBM Type 1 Events continued from previous page

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3	reak	гa	Dec	Error	Name(distance) (sigma)	bb temp keV	$10-8 \text{ erg cm}^{-2} \text{ s}^{-1}$	$10^{-7} \mathrm{erg} \; \mathrm{cm}^{-2}$	FL maex	$10^{-8} {\rm erg \ cm^{-2} \ s^{-1}}$	$^{1}$ $^{10^{-7}}$ erg cm $^{-2}$	Rise	sec	Duration	Structure
10080513547 1	12627941	301.4	-4.9	8.9	XB_1940-04 (0.8) 4U_1915-05 (1.7)	$3.71 \pm 0.31$	$1.41 \pm 0.14$	$4.04 \pm 0.41$	$-2.61 \pm 0.21$	$2.22 \pm 0.46$	$6.3 \pm 1.3$	1.58	18.82	20.64	M
			8.6	2.5		$3.312 \pm 0.048$	+ +	$27.11 \pm 0.48$	+1 -	+ -	$34.3 \pm 1.0$	5.96	23.28	30.67	ω i
10081304857 1	13310455	144.0	-53.2	2.5	$2S_{-}0918-549$	$3.09 \pm 0.33$ $2.27 \pm 0.10$	$1.87 \pm 0.66$ $1.81 \pm 0.10$	$2.29 \pm 0.80$ 4 43 + 0.24	$-3.34 \pm 0.34$ $-3.95 \pm 0.18$	$4.2 \pm 1.3$ $1.99 \pm 0.15$	$5.1 \pm 1.6$ $4.88 \pm 0.38$	2.50	7.74	12.62 20.94	w w
			-21.6	10.5				3.89 ± 0.59	1 +1	1 +1	$5.6 \pm 1.1$	2.34	3.18	6.02	o w
10081545957 1	13524362	285.0	3.2	8.9		$2.507 \pm 0.095$	+	$7.55 \pm 0.36$	$-3.67\pm0.13$	+	$8.27 \pm 0.55$	1.87	9.19	23.82	M
		91.7	16.0	12.5	4U0614+09		+	$2.85 \pm 0.26$	+	+	$3.73 \pm 0.60$	2.10	14.71	17.25	w
			8.8	3.8			₩-	$25.43 \pm 0.83$	+ -	+1 -	$29.0 \pm 1.4$	62.18	114.13	179.18	ω <u>;</u>
			-26.9	12.4			H -	$2.69 \pm 0.29$	# -	# -	$3.69 \pm 0.75$	1.49	12.49	14.32	M
	13832150		-11.4	11.0		$3.14 \pm 0.15$	2.61 ± 0.15	$5.32 \pm 0.31$	-3.16 ± 0.16	3.18 ± 0.30	6.49 ± 0.63	13.31	5.00	18.99	M N
10082073275 1		202.8 242.8	-29.0 25.7	11.9		3.16 ± 0.22	2.17 ± 0.18	4.44 ± 0.37	ΗН	2.72 ± 0.37	5.55 ± 0.76	11.04	11.04	23.90	M O
			-55.7	10.1			2.01 ± 0.13 3.36 ± 0.17	6 87 ± 0.88	-2.80 ± 0.13	H +	7.48 ± 0.58	12.06	6.07	20 70	o >
			-32.2	15.3			+ +	$1.10 \pm 0.26$	4 +	+	$2.63 \pm 0.56$	1.55	4.10	6.03	z v
			-55.0	2.1			$2.33 \pm 0.16$	$22.8 \pm 1.5$	1 #1	1 +1	$27.5 \pm 2.5$	27.82	86.06	137.35	ı w
			-23.0	18.5		0	+	$1.48\pm0.19$	$-2.53\pm0.26$	+	$2.66 \pm 0.69$	1.24	86.9	8.58	Ø
			-7.1	17.7			$1.96 \pm 0.16$	$5.61 \pm 0.46$	$-2.79\pm0.18$	+ +	$7.6\pm1.2$	13.59	9.64	23.84	M
			-56.0	24.1		0	₩.	$1.46 \pm 0.21$	₩.	+1 -	$2.96 \pm 0.90$	1.68	8.42	12.07	M
		281.0	8:1	12.3		$2.08 \pm 0.26$		$5.22 \pm 0.77$	$-4.61 \pm 0.53$	<b>H</b> -	$5.64 \pm 0.88$	1.87	3.80	6.96	m;
		259.0	9.7-	7.7			Н-	1.76 ± 0.38	$-3.36 \pm 0.30$	H -	$9.2 \pm 1.1$	7.16	3.27	17.17	M
	15521084	293.9	13.8	1.0		$2.67 \pm 0.12$	$1.481 \pm 0.087$	$7.25 \pm 0.42$	-3.33 ± 0.16	$1.77 \pm 0.14$	8.68 # 0.68	23.40	21.27	48.62	n c
100907060698 1			2.1.7	10.1			$1.222 \pm 0.090$ 1 98 $\pm$ 0 21	$11.91 \pm 0.94$ $169 \pm 0.17$	-3.30 ± 0.22 -2.99 ± 0.28	1.40 ± 0.18	14.3 H 1.8	95.69	0.4.55	19 90	מ מ
			. x	16.9			$1.53 \pm 0.21$	$2.50 \pm 0.17$	$-2.30 \pm 0.23$ $-2.70 \pm 0.23$	+ +	$4.02 \pm 0.38$	2 10 10 10 10 10 10 10 10 10 10 10 10 10	0.00	14 94	o v
		157.9	α. 2. 2. 3. 3. 3. 3. 3. 3. 3. 3. 3. 3. 3. 3. 3.	3.1			2.994 + 0.090	51.3 + 1.5	$-2.579 \pm 0.23$	+	78.9 + 5.4	233.69	106.06	343.29	⊇
			-25.1	1.0		$3.21 \pm 0.17$	$2.82 \pm 0.17$	$3.45 \pm 0.21$	$-2.97 \pm 0.15$	1 +1	$4.68 \pm 0.51$	9.04	5.62	15.49	w
			15.2	4.5	4U_0614+09		+	$16.09 \pm 0.41$	+	+	$19.06 \pm 0.76$	2.73	24.53	28.32	S
10091109490 1	15820691	243.9	-42.7	11.4		$3.14 \pm 0.23$	$1.79 \pm 0.15$	$2.20 \pm 0.18$	$-3.12 \pm 0.24$	$2.33 \pm 0.33$	$2.86 \pm 0.40$	1.72	9.73	12.02	Q
			-26.5	10.7		0	+	$0.93 \pm 0.28$	#	$\mathbb{H}$	$1.90 \pm 0.56$	1.46	5.07	10.73	S
			-57.6	12.2		$3.16 \pm 0.22$	₩.	$2.39 \pm 0.19$	$-3.18 \pm 0.22$	+ -	$2.87 \pm 0.39$	3.38	5.87	10.94	w ¦
			-25.9	7.0			+ -	$4.31 \pm 0.37$	+ -	+1 -	$5.36 \pm 0.76$	1.61	15.44	17.52	M o
			-12.8	6.1			₩ -	H -	H -	H -	$9.00 \pm 0.48$	3.47	15.54	19.79	N (
10091509603 1	16166406	256.3	-27.6	10.9		$2.96 \pm 0.16$	$1.80 \pm 0.11$	$4.41 \pm 0.29$	$-3.34 \pm 0.18$	$2.12 \pm 0.21$ 2 68 $\pm$ 0 33	$5.19 \pm 0.51$	5.93	20.27	10.03	w Ş
			-20.1	χ <u>τ</u>			H +	2.55 ± 0.13	H +	H +	3.31 ± 0.40	0.77	6.52	11.89	Z N
			-14.6	. x			+	$4.08 \pm 0.20$	1 +	1 +	$5.91 \pm 0.35$ $5.01 \pm 0.39$	86.8	99.6	19.32	, so
			-18.9	8.6				$2.50 \pm 0.23$	1	1	$3.18 \pm 0.58$	3.21	7.84	11.59	S
			-20.5	11.6		0	+	$4.80 \pm 0.44$	$-2.63 \pm 0.20$	+	$6.8 \pm 1.3$	8.10	5.92	14.64	Ø
			-33.6	15.1		$3.70 \pm 0.34$	$1.97 \pm 0.20$	$1.60 \pm 0.16$	$-2.94 \pm 0.27$	$2.52 \pm 0.46$	$2.05 \pm 0.38$	2.79	4.97	8.49	w
		262.0	7.5	13.1		0	+	# -	$-2.57 \pm 0.13$	$3.08 \pm 0.40$	$71.6 \pm 9.5$	177.30	213.28	396.47	SO :
10092223972 1	16785400	243.5	15.0	17.4 S	UW_Crb (0.6) SAX J1818.7+1424 (1.7)	$3.25 \pm 0.13$	$1.915 \pm 0.093$	$78.1 \pm 3.8$	$-2.95\pm0.12$	$2.46 \pm 0.23$	$100.4 \pm 9.5$	123.74	300.16	440.16	SO.
10092524811 1	17045618	266.2	-14.9	7.6		$3.00\pm0.19$	$2.30 \pm 0.18$	$4.69 \pm 0.36$	$-3.20 \pm 0.20$	$2.87 \pm 0.35$	$5.87 \pm 0.71$	1.15	3.97	17.55	M
			-34.4	6.6		$2.96 \pm 0.23$	+	$3.10 \pm 0.29$	$-3.21 \pm 0.24$	+	$3.84 \pm 0.54$	3.04	3.91	11.50	S
			-20.7	6.9			+ -	$3.25 \pm 0.32$	+ -		$4.10 \pm 0.71$	7.52	8.57	16.62	w c
			-13.6	19.8		$3.79 \pm 0.42$	$1.81 \pm 0.24$	$2.22 \pm 0.29$	Н-	H -	$2.52 \pm 0.45$	4.22	5.20	10.07	n ;
		289.6	15.9	9.5		$3.03 \pm 0.10$		$7.02 \pm 0.30$	$-3.19 \pm 0.11$	H -	$8.49 \pm 0.61$	9.54	8.03	28.11	M
101001/5234 1	17629594	290.1 291.5	22.4	13.5		$3.51 \pm 0.20$ $3.19 \pm 0.24$	2.13 ± 0.13	$8.91 \pm 0.61$ 2 60 + 0.23	$-2.87 \pm 0.16$ $-2.90 \pm 0.21$	$2.34 \pm 0.22$ $2.79 \pm 0.32$	10.53 ± 0.98	8.55 8.55	5.21	40.15 14.65	ΩŒ
		281.6	24.3		SAX 11818 7+1424	$3.27 \pm 0.25$	1.66 + 0.15	3.40 + 0.31	200	+ +	$4.20 \pm 0.39$	13.52	80.9	20.10	o v
	_	286.0	-2.2	_		$3.71 \pm 0.27$	1 +	$3.08 \pm 0.25$	1 +	1 +	$3.78 \pm 0.42$	2.92	7.44	11.11	o vo
			-28.8	6.3		$3.35 \pm 0.10$	$1.806 \pm 0.066$	$12.53 \pm 0.46$	$-2.960 \pm 0.094$	1	$15.44 \pm 0.75$	25.52	51.68	85.05	w
			-30.7	9.2		$2.83 \pm 0.23$		$3.35 \pm 0.31$	# -	+1 -	$3.91 \pm 0.43$	8.81	28.80	39.30	Ø.
			-14.4	1.4		$3.012 \pm 0.098$	+ -	$17.19 \pm 0.68$	+1 -	# -	$19.5 \pm 1.1$	35.58	21.78	59.20	w i
10100655628 1	18026828	260.9	-39.7	16.7		$3.52 \pm 0.26$	$3.44 \pm 0.30$	$2.81 \pm 0.24$	$-2.90 \pm 0.20$	$4.05 \pm 0.45$	$3.31 \pm 0.37$	5.18	6.16	11.78	w w
			9.	11.1		$\overline{}$	Н	17.0 日 07.7	Н	Н	2.02 H 0.00	1.00	5.0	0.10	Q

Table 5:: GBM Type 1 Events continued from previous page

<u> </u>	Peak	Ra	Dec	Error	Name(distance)	$_{ m keV}^{ m keV}$	BB flux 10-8 erg cm <sup>-2</sup> s <sup>-1</sup>	$^{10-7}_{ m erg~cm}^{-2}$	PL index	$^{ m PL~Flux}_{10^{-8}{ m erg}~{ m cm}^{-2}~{ m s}^{-1}}$	PL Flnc $10^{-7}  \mathrm{erg  cm^{-2}}$	Rise	Fall I	Duration S	Structure
10100818955	18162967	228 1	8 95	10.1	(2000)	3 37 + 0 21	2.57	3 14 + 0 24	$-2.96 \pm 0.18$	3.02 + 0.29	3 70 + 0 36	1 88	4 22	10.32	V.
	18236081	259.4	-27.0	19.9		$3.64 \pm 0.30$	$2.71 \pm 0.26$	1 #1	1 44	1 +1	$3.81 \pm 0.47$	1.67	5.11	7.18	o oo
10101200985	18490583	262.7		11.3		$3.23 \pm 0.19$		$3.51 \pm 0.24$	$-3.25 \pm 0.19$	$3.08 \pm 0.28$	$3.77 \pm 0.34$	4.22	8.62	13.18	W
	18503686	94.9	13.7	7.6	$4U_0614+09$		$1.545 \pm 0.093$	$+\!\!\!+\!\!\!\!+$	$-3.47 \pm 0.20$	$1.81 \pm 0.16$	$7.40 \pm 0.68$	2.59	14.71	19.77	S
	18742345	264.4		5.7		$3.51 \pm 0.19$	$1.75\pm0.11$	+	$-2.97 \pm 0.15$	+	$11.3 \pm 1.1$	15.46	5.02	21.29	M
	18661444	197.8		11.1			# -	$7.13 \pm 0.85$	44 -	# -	$8.1 \pm 1.1$	1.52	7.52	13.32	w t
10101546522	18708924	234.6	-73.4	12.4		$3.15 \pm 0.28$	$3.48 \pm 0.36$	5.68 \(\pi\) 0.58	$-3.14 \pm 0.27$	$4.21 \pm 0.52$	6.88 # 0.85	5.61	3.74	11.12	w o
	10033021	200.9		4.1.		2.03 H U.14	2.04 H 0.16	4.31 H 0.26	-3.60 H 0.19	2.00 H 0.24	4.71 H 0.40	9.78	10.97	14.00	ΩŅ
	18930404	235.2	-49.3	0.0		3.50 ± 0.18	2.99 ± 0.21 2.86 ± 0.26	H +	-2.85 + 0.22	H +	2 83 + 0.33	1.70	- 8 0 6 73	14.09 8 93	N V
	18972485	250.5	-59.7	16.6			$2.69 \pm 0.24$	$3.29 \pm 0.29$	$-2.83 \pm 0.21$	$3.54 \pm 0.59$	$4.33 \pm 0.72$	4.67	4.01	9.57	ω
	18948275	244.3	-71.0	13.1		$2.98 \pm 0.23$	$6.77 \pm 0.60$	1	$-3.46 \pm 0.26$	1	$9.36 \pm 0.99$	4.79	5.31	11.05	S.
10101927930	19122334	296.1	-17.9	29.1			$0.61 \pm 0.14$	+	$-2.91 \pm 0.24$	$1.01 \pm 0.24$	$1.66 \pm 0.40$	4.22	6.73	11.61	Ø
10101974133	19168534	256.3	-51.4	11.2			$3.64 \pm 0.34$	$1.48 \pm 0.14$	$-2.92 \pm 0.23$	$4.97 \pm 0.78$	$2.03 \pm 0.31$	4.62	3.08	9.35	w
	19301624	93.4	13.3	5.7	$4U_0614+09$		+	$7.96 \pm 0.26$	$-3.416 \pm 0.096$	+	$9.31 \pm 0.46$	2.66	11.28	14.45	S
_	19373390	230.5	-54.3	15.2			$2.04 \pm 0.18$	+	$-2.90 \pm 0.21$	+	$4.33 \pm 0.69$	2.48	6.03	9.46	w
	19395912	253.1	-40.5	15.9		0	$3.09 \pm 0.26$	# -	# -	# -	$3.23 \pm 0.45$	2.38	1.90	6.18	w c
	19448171	248.6	-51.9	11.0			$1.88 \pm 0.15$	$3.07 \pm 0.25$	$-2.98 \pm 0.21$	<b>H</b> -	$3.53 \pm 0.37$	3.04	3.75	8.00	w c
	19742140	200.4	-31.1	T 1.1		<u> </u>	H -	Н-	$-2.93 \pm 0.32$	Н-	2.95 ± 0.67	1.05	10.20	12.16	Ωō
10102823100	19868303	272.9	4.84 4.04	15.5 7.3		$3.44 \pm 0.30$	$1.85 \pm 0.18$ $3.64 \pm 0.48$	$1.51 \pm 0.15$ $1.48 \pm 0.19$	$-2.89 \pm 0.25$ $-2.01 \pm 0.32$	2.59 ± 0.47	$2.12 \pm 0.38$	2.88	3.08	5.27	ט ט
	10040710	0.4.0	10.0	0.7 C.71			3.04 ± 0.48	2 07 ± 0.34	9.67 ± 0.35	НН	2 55 ± 0.43	0 1.7	00.0	19.70	ס מ
	20132143	970.9	6.01-	11.1			3.01 ± 0.14	2.01 ± 0.24 2.46 ± 0.17	4 +	+	3.06 ± 0.75	0 7.7 7.7	7.00	8 77	מ מ
	20102143	267.2	15.7	1.1.			2.96 + 0.35	$1.21 \pm 0.17$	$-2.71 \pm 0.26$	+ +	1.98 + 0.43	1.36	5.40	7.01	o v
	20192402	285.0	-18.7	8.2			$2.49 \pm 0.17$	$3.05 \pm 0.21$	$-3.24 \pm 0.20$	+	$3.48 \pm 0.31$	6.37	4.18	11.71	ı va
	20284403	257.1	-47.9	14.5			$1.52 \pm 0.13$	$3.10 \pm 0.27$	$-2.99 \pm 0.22$	1 +	$3.99 \pm 0.62$	8.57	6.03	15.33	Σ
	20330498	270.6	15.6	15.0			$8.9 \pm 1.0$	$10.9 \pm 1.2$	$-4.74 \pm 0.40$	1	$11.5 \pm 1.2$	2.06	5.36	7.82	Ø
	20342173	275.1	-40.9	12.9			+	$2.60 \pm 0.19$	+	#	$3.13 \pm 0.37$	1.27	5.08	12.78	M
	20674497	243.1		8.9			$0.79 \pm 0.13$	$2.25 \pm 0.38$	$-2.49 \pm 0.34$	+	$3.8 \pm 1.3$	3.51	88.9	10.87	w
	20721266	294.0		12.5		0	+	$3.71 \pm 0.33$	+	+	$4.42 \pm 0.52$	8.88	5.64	15.11	M
	20751547	267.8	-32.1	13.7			$2.54 \pm 0.20$	$2.07 \pm 0.17$	$-2.74 \pm 0.18$	# -	$3.20 \pm 0.46$	1.58	5.96	7.90	ω;
	20825089	253.4	-45.9	6.5			$0.918 \pm 0.097$	$21.7 \pm 2.3$	$-3.62 \pm 0.29$	H -	23.9 ± 3.3	193.77	60.79	259.91	M c
~	20834608	276.0	-30.6	14.5			$2.81 \pm 0.22$	$2.29 \pm 0.17$	$-2.94 \pm 0.19$	H -	$3.03 \pm 0.42$	28.5	6.02	12.83	n ;
10110812871	20835277	273.3	-11.8	7.4		$3.15 \pm 0.24$ $2.75 \pm 0.20$	$2.15 \pm 0.19$ 1 83 $\pm$ 0 17	$3.51 \pm 0.32$ 5 22 $\pm$ 0 40	$-3.07 \pm 0.23$	$2.52 \pm 0.30$	$4.12 \pm 0.49$ 6 56 $\pm$ 0.78	8.99	7.57	17.55	Σv
	20859105	0.707		4.6			H +	6.39 ± 0.43	$-3.28 \pm 0.14$	3.08 ± 0.27	7.54 + 0.52	0.70	19.70	21.22	n >
	20880182	125.6		0.0	2S_0918-549			$6.35 \pm 0.40$	$-3.28 \pm 0.17$	1 +	$7.50 \pm 0.63$	7.14	12.21	21.22	E o
10110938413	20947217	255.2	-36.3	5.5			$2.11 \pm 0.17$	+	$-3.03 \pm 0.24$	$2.51 \pm 0.27$	$3.08 \pm 0.33$	15.06	9.85	25.65	M
	21432624	284.2	-28.1	10.8		0	+	+	#	$\mathbb{H}$	$2.19 \pm 0.35$	3.73	4.24	8.36	S
	21532012	291.0		7.9		$3.50 \pm 0.22$	+ +	$0.83 \pm 0.16$	$-2.92 \pm 0.19$	+1 -	$1.67 \pm 0.33$	13.93	3.50	18.57	M
10111932453	21805260	273.5	-10.6	8.9 8. u		$2.96 \pm 0.13$ 3.45 $\pm$ 0.23	$2.92 \pm 0.16$	$5.97 \pm 0.32$	$-3.30 \pm 0.15$ $-2.88 \pm 0.20$	$3.15 \pm 0.23$ $3.24 \pm 0.51$	$6.43 \pm 0.47$ 5.28 $\pm$ 0.84	3.03 8.03	2.71	16.00	Συ
	21937011	269.6	-30.5	8.6		$3.16 \pm 0.30$	+	+	$-3.18 \pm 0.30$	+	$2.86 \pm 0.51$	4.23	6.83	11.55	o vo
10112248938	22080937	300.9	-25.7	17.0				$1.90 \pm 0.18$	$-3.02 \pm 0.24$	$3.05\pm0.50$	$2.48 \pm 0.41$	1.78	7.15	9.36	S
	22207626	289.5	-27.3	12.5			#	$4.04 \pm 0.27$	$-3.30 \pm 0.19$	$\mathbb{H}$	$4.47\pm0.51$	9.44	7.29	17.48	w
	22492258	275.2	-15.5	8.0		$3.34 \pm 0.12$	$3.97 \pm 0.17$	$6.49 \pm 0.27$	$-3.13 \pm 0.11$	+ -	$7.38 \pm 0.41$	1.91	6.01	14.33	M
	22477847	91.2	17.1	5.3	40_0614+09	$3.651 \pm 0.076$	H -	$12.25 \pm 0.29$	9.00	H -	$16.57 \pm 0.75$	2.40	13.62	18.16	n ;
10113068478	22791695	281.7	13.6		SAX_J1818.7+1424 (1.0) Ser_X-1 (1.3)	$2.57 \pm 0.21$	$1.55 \pm 0.16$	$5.71 \pm 0.62$	$-3.67 \pm 0.30$	$1.68 \pm 0.23$	$6.16 \pm 0.87$	12.23	10.67	23.19	M
10119084769	25907061	111	000	Sw 7 1	Swift_J185003.2-005627 (	4 90 6	900 + 016		4 00 6		4 88 6	2	1 1	Б	ŭ
	24563207	33.0	0.02	7. T		3.00 H 0.13	H +	3.78 H 0.19	0 4 H X H +	H +	2 22 + 0 36	1.01	4.57	6.10	o o
	24654980	312.7	-51-9	13.0		3.07 + 0.20	+ +	+ +	3.31	+	3.38 + 0.33	2.35	986	12.72	o o
	24839171	238.1	-23.6	9.5		$3.07 \pm 0.25$	1 +1		$-3.19 \pm 0.27$	1 +1	$6.84 \pm 0.93$	2.82	5.94	9.31	N W
10123019999	25248806	108.4	-53.6	4.0		$3.26 \pm 0.15$	$1.78 \pm 0.10$	+	$-2.99\pm0.14$	$2.37 \pm 0.23$	$10.6 \pm 1.0$	41.50	46.14	93.73	W
	25560667	261.8	-40.9	8.0		$3.60 \pm 0.22$	$2.03 \pm 0.14$	$3.32 \pm 0.23$	$-2.97\pm0.18$	$2.62 \pm 0.32$	$4.28 \pm 0.52$	4.59	8.19	13.58	S
11010204290	25492292	248.2	-42.7	5.2		$2.97 \pm 0.20$	$2.09 \pm 0.17$	$4.27 \pm 0.35$	$-3.20 \pm 0.22$	$2.46 \pm 0.26$	$5.03 \pm 0.53$	2.00	4.46	11.90	ω

Table 5:: GBM Type 1 Events continued from previous page

E	-	ė		F	( ); ( ); ( )	9	ממ	ם מם	10	Id	10		-		
j	S	1.00	3		(sigma)	keV		$10^{-7} {\rm erg \ cm^{-2}}$	7	$10^{-8} \text{erg cm}^{-2} \text{s}^{-1}$	$10^{-7} \text{ erg cm}^{-2}$	sec		sec	o in a contra
11010250886	25538890	219.2	-7.0	3.9		$3.09 \pm 0.30$	$1.34 \pm 0.16$	$3.30 \pm 0.40$	$-3.03 \pm 0.29$	$1.77 \pm 0.35$	$4.34 \pm 0.87$	10.18	66.9	18.54	S
11010325049	25685847	118.9	-16.4	6.9		$2.56 \pm 0.11$	$3.66 \pm 0.19$	$^{\rm H}$	#	$\mathbb{H}$	$5.14 \pm 0.38$	2.15	5.39	9.83	w
11010413969	25761180	303.7	-62.3	8.6		$3.67 \pm 0.36$		+1 -	41 -	+ +	$3.54 \pm 0.81$	3.45	21.30	25.39	w i
11010448458	25795662	232.3	-41.1	10.2		$4.14 \pm 0.29$	$2.95 \pm 0.23$	$2.41 \pm 0.19$	Н-	H -	$4.10 \pm 0.66$	3.70	3.36	8.16	w ;
11010449579	25796791	163.8	-51.5	7.01	38 0018 570	$3.02 \pm 0.31$ $2.68 \pm 0.40$	$2.28 \pm 0.28$ 1 40 $\pm$ 0 21	5.60 ± 0.69	$-3.12 \pm 0.31$	$2.80 \pm 0.40$	6.86 ± 0.99	7.23	5.16 F 41	13.03	N v
11010582534	25916138	258.2	-59.0	7.5	01010107	$3.18 \pm 0.25$	$1.30 \pm 0.21$ $1.30 \pm 0.11$	$2.66 \pm 0.24$	$-3.03 \pm 0.43$ $-3.00 \pm 0.23$	1 +	$3.42 \pm 0.39$	1.65	9.22	11.33	o vo
11010620727	25940731	266.8	-13.2	8.9			$2.27 \pm 0.12$	1 +1	1	+	$6.02 \pm 0.51$	10.86	4.79	16.18	M
11010726558	26032961	258.7	-32.5	6.7		$3.25\pm0.30$	$1.65 \pm 0.18$	+	$-3.13 \pm 0.30$	$1.88 \pm 0.36$	$2.30 \pm 0.45$	2.67	3.89	8.82	w
11010813802	26106601	236.5	-62.1	16.3			$2.45 \pm 0.31$	+	$-3.22 \pm 0.33$	$3.29 \pm 0.58$	$5.37 \pm 0.95$	1.51	8.78	10.55	w
11010842360	26135177	236.1	-24.9	8.1			$3.59 \pm 0.46$	+1 -	$-2.53 \pm 0.26$	+1 -	$1.85 \pm 0.28$	2.39	4.11	7.29	w i
11010912247	26191447	259.1	-28.5	6.1			$2.58 \pm 0.23$	+ -	2.92	# -	$2.84 \pm 0.44$	1.81	5.18	9.07	w i
11011063322	26242531	248.8	-28.6	9.4		$4.35 \pm 0.85$	$0.97 \pm 0.21$	+1 -	$-2.31 \pm 0.40$	$1.37 \pm 0.32$	$2.23 \pm 0.53$	2.62	5.03	8.18	w o
11011173802	26425793	84.6	ος co	8.4	4U_0614+09		$4.20 \pm 0.15$	H -	.001 ±	<b>H</b> -	$10.97 \pm 0.70$	2.64	14.18	17.50	w c
11011455616	26666810	89.3	16.2	16.3	40-0614+09	$3.30 \pm 0.26$	$2.15 \pm 0.20$	$1.76 \pm 0.16$	3.06 #	H -	$2.16 \pm 0.36$	1.84	80.00 80.00	6.02	v c
11011/22238	20800233	243.4	-47.9	0.4			2.70 ± 0.19	4.40 ± 0.31	Н-	Н-	4.95 # 0.47	0.00	8.12 144 of	15.51	Ωα
11011727613	26811654	300.9	21.2	7 -		3.33 ± 0.12	$1.940 \pm 0.088$	$17.41 \pm 0.79$	$-2.82 \pm 0.10$	$2.87 \pm 0.24$	Z5.7 ± Z.1	34.05	10.69	220.61	במ
11011/45692	27012303	0.412.0	0.12-	1. J		2.70 H 0.13	2.39 ± 0.21	4.22 H 0.33	ΗН	3.10 ± 0.40	0.17 H 0.65	12.34	10.07 07 7	11 22	IVI O
11011882052	27038861	2.00	13.9	10.4	411 0614±09		1 +	1 +	1 +	1 +	6.76 + 0.56	9.01	10.97	15.44	o o
11012121527	27237535	274.4	-24.1	2.8	1	$3.16 \pm 0.19$	$1.92 \pm 0.13$	+	3.17 +	+	$4.68 \pm 0.54$	9.76	6.95	18,09	Σ
11012277383	27379783	232.2	-64.0	17.9		$4.13 \pm 0.35$	+	1 +1	2.79 土	54 H	3.85 ± 0.76	1.62	7.10	9.04	ß
11012370122	27458916	279.4	12.1		SAX_J1818.7+1424 (0.9)	$3.07 \pm 0.11$	$1.749 \pm 0.072$	$10.00\pm0.41$	$-3.09\pm0.11$	$2.13 \pm 0.11$	$12.18 \pm 0.68$	11.59	35.61	49.85	w
					Ser_X-1 (1.3)										
11012439618	27514823	86.1	9.2	0.9	4U0614+09	$3.71 \pm 0.13$	+ -	$7.33 \pm 0.32$	$-2.76 \pm 0.10$	+ -	$10.45 \pm 0.92$	2.65	10.97	14.26	ω;
11012700992	27735397	276.5	-28.5	×			Η.	Н.		₩ .	$4.73 \pm 0.49$	9.33	48.8	13.84	M
11012819227	27840026	7.88.7	18.3	7:5	$SAX_J1818.7+1424$	$3.26 \pm 0.21$	₩-	$7.56 \pm 0.58$	H -	H -	$9.14 \pm 0.89$	29.35	42.78	75.61	v c
11012875313	27896138	281.5	0.4 0 r	1.6 0.0			Н-	10.40 ± 0.68	Н-	Н-	$12.0 \pm 1.0$	24.80	17.07	71.42	מ מ
11012947891	27955100	260.0	رة ت ز	x 5		$3.17 \pm 0.14$	$3.06 \pm 0.16$	H -	Н-	Н-	6.6U ± 0.65	12.62	0.93	19.81	M Z
11012962750	27005054	0.077	9000	0.12		3.45 H 0.50	2.12 H 0.28	7 04 ± 0.23	-3.23 H 0.28	3.00 ± 0.52	2.45 H 0.42	07.0	7.95	10.21	M
11013001338	27896988	2.4.7	36.0	15.0			H +	H +	0.04 76 H	H +	9.9 H 1.3	2.13	00.7	7.7.0	īνī
11013143903	28123897	265.5	-31.0	14.0		$3.20 \pm 0.27$	+ +	$2.29 \pm 0.23$	3.03 +	+ +		3.12	4.29	7.94	o vo
11013161975	28141978	270.4	-28.3	12.7				1 +1	3.16 ±	1 +1	1 +1	2.52	4.12	8.06	o co
11020266346	28319148	241.4	-38.1	14.1			+	+	+	+	+	2.95	6.10	89.6	w
11020334488	28373687	246.6	-8.4	12.4		$3.13 \pm 0.25$	$2.16 \pm 0.20$	+	#	+	+	2.02	4.02	6.41	w
11020404516	28430120	258.4	-9.1	10.8			+ +	# -	2.86 ±	+ +	₩.	2.58	3.55	7.00	w ¦
11020451394	28476996	282.4	-31.1	11.3		$3.20 \pm 0.15$	$2.32 \pm 0.12$	$3.79 \pm 0.20$	$-3.13 \pm 0.15$	$2.90 \pm 0.27$	$4.75 \pm 0.44$	1.52	3.52	11.52	Σū
11020347877	28588233	7.011	782	19.5		0.0 H 1.0 4 47 + 0 50	$1.13 \pm 0.21$ $1.75 \pm 0.22$	H +	2.24 2.37 H	H +	H +	20.0	0.00 00.00	2 8	מ מי
11020579446	28591447	49.8	-57.0	11.4			1 +1	1 +	2.80 #	+	1 +1	6.20	25.27	32.66	υw
11020683414	28681805	289.8	-16.0	8.2		$2.64 \pm 0.10$	$2.68 \pm 0.13$	+	+	+	+	1.23	14.35	15.75	M
11020848578	28819770	321.2	-37.1	7.3		$1.714 \pm 0.060$	+	+	+	+	+	4.41	13.32	18.11	S
11020913639	28871238	259.1	-11.6	10.2		$3.23 \pm 0.13$	+1 -	+	+	+1	+	2.15	18.65	21.08	M
11020913936	28871541	270.1	-38.5	17.2		$3.37 \pm 0.21$	₩ -	+1 -	# -	# -	+ -	3.28	9.77	19.42	Σ
11021048257	28992261	239.5	-19.9	12.6			H -	H -	H -	H -	H -	1.32	5.85	7.43	ΩŒ
11021076838	29020844	200.5	4.7-	10.0 7.0		$3.71 \pm 0.42$ $2.14 \pm 0.15$	1.80 ± 0.25	$2.27 \pm 0.30$	$-2.61 \pm 0.28$	$2.90 \pm 0.77$	3.55 ± 0.94	4.20	9.20	14.20	υŽ
11021270647	29187444	275.7	-15.6	5.7		0	+	++	++	+	+	1.56	11.95	14.17	Z
11021373353	29276556	265.6	-16.5	15.1		4.11 + 0.33	+	+	+	+	+	1.68	11.34	13.25	O.
11021511600	29387612	306.6	13.9	7.6			#	1	1	1	1	19.04	39.84	83.11	ı w
11021529971	29405970	254.6	-47.5	8.1		$4.34 \pm 0.46$		$1.55 \pm 0.18$	$-2.59 \pm 0.28$	$2.90 \pm 0.79$	+	1.53	5.60	7.45	ß
11021674170	29536576	272.6	-24.0	15.1			+	$^{\rm H}$	$-2.98 \pm 0.21$	$^{\rm H}$	$4.49 \pm 0.66$	6.45	3.49	10.68	SO
11021745757	29594558	256.0		5.2		$2.94 \pm 0.16$	$4.41 \pm 0.31$	+	$-3.35 \pm 0.19$	+	$6.59 \pm 0.63$	1.16	12.50	14.04	M
11021808838	29644046	246.9		13.0			$2.39 \pm 0.22$	+ -	$-2.77 \pm 0.22$	+ -	+1 -	2.66	4.66	7.94	ω;
11021909951	29731560	303.3	39.9	9.01	40.2129+47 (1.5) Cvg.X-2 (1.7)	$3.26 \pm 0.32$	$1.88 \pm 0.22$	$3.83 \pm 0.44$	$-2.88 \pm 0.26$	$2.43 \pm 0.35$	$4.96 \pm 0.72$	1.99	6.46	16.98	M

Table 5:: GBM Type 1 Events continued from previous page

	Ка	Dec	Error	Name(distance)	+ 0000	אויי			1	DI H	Sign	F.a.		
				(sigma)	bb temp keV	$10-8 \text{ erg cm}^{-2} \text{ s}^{-1}$	$10^{-7}$ erg cm $^{-2}$	FL index	$10^{-8} {\rm erg \ cm^{-2} \ s^{-1}}$	$^{1}$ $^{10-7}$ $^{\rm erg~cm}^{-2}$	sec		Duration 5	Structure
	9 267.7	-31.5	11.2	(8)	$3.48 \pm 0.26$	2.57	$2.10 \pm 0.18$	-2.91 + 0.22	3.33 + 0.53	$2.72 \pm 0.43$	2.16	6.93	9.43	S
		-17.9	3.5		$4.73 \pm 0.26$		$29.7 \pm 1.8$	$-2.55 \pm 0.14$	1	$44.9 \pm 6.2$	5.03	42.38	49.57	M
			7.4		$3.08 \pm 0.17$		$2.90 \pm 0.19$	$-3.30 \pm 0.19$	$2.84 \pm 0.29$	$3.48 \pm 0.35$	1.73	8.74	10.80	ß
		-36.8	17.9		$3.21 \pm 0.25$	$1.61 \pm 0.15$	$2.63 \pm 0.24$	$-2.99 \pm 0.24$	$2.07 \pm 0.34$	$3.38 \pm 0.56$	2.67	6.57	9.74	S
			18.8		$3.60 \pm 0.32$	$1.92 \pm 0.19$	$2.35 \pm 0.24$	$-2.98 \pm 0.26$	+	$2.87 \pm 0.52$	1.56	3.71	5.65	S
		-29.2	9.4			₩.	$1.53 \pm 0.14$		# -	$1.83 \pm 0.32$	1.70	4.22	6.17	w :
		0.1	£.4 6.6		$3.043 \pm 0.071$	$7.01 \pm 0.20$	$11.45 \pm 0.32$	$-3.251 \pm 0.077$	H -	$13.55 \pm 0.62$	2.69	9.16	16.37	ΩŞ
		4.T	12.8			H -	H -	H -	H -	5.97 H 0.70	11.21	10.7	19.47	M
		-13.5	15.5		$2.83 \pm 0.19$	$2.27 \pm 0.19$	2.78 ± 0.23	$-3.37 \pm 0.23$	H +	$3.28 \pm 0.42$	1.II	10.50	12.03	Ωō
11022312308 30232324		10.07	16.0			2.00 H 0.14	5.57 H 0.25	-5.54 H 0.19	Н -	4.05 H 0.41	0.00	0.40	13.49	n u
	261.0	-13.4	10.8			H +	$1.88 \pm 0.22$	$-2.87 \pm 0.29$	$2.05 \pm 0.45$	$2.51 \pm 0.56$ $4.00 \pm 0.57$	7. 1. 4. 2. 4.	1 0.0 1 0.0 1 0.0	3.75	ΩŞ
		0.72-	n 0			1.90 ± 0.15	Н-	$-3.07 \pm 0.21$	Н-	4.09 \(\pi\) 0.57	00.7	7.08	17.10	M
		4.0-	1 m		2.23 ± 0.10	5.08 H 0.18	5.05 ± 0.29	н -	Н-	0.02 H 0.45	1.93	10.47	10.07	M
11030464436 30929224	0.107 #	100.0	- 1 i c		3.00 H 0.17	2.17 H 0.14	4.44 H 0.29	-5.20 H 0.17	2.71 H 0.27	0.00 H 0.00	90.00	0.20	17.75	M
			- 1 9 0			2.02 ± 0.14 1.86 ± 0.19	H +	-3.33 H 0.13	2.94 ± 0.20 3.04 ± 0.41	0.40 H 0.30	3.00	16.75	27.72	Ţυ
			0.0			1.00 ± 0.12	4 +	4 4	4 4	0.4 H 1.0	10.40	10.70	4 1 . C . T . T . T . T . T . T . T . T . T	ט מ
		40.07	19.3			1 71 + 0 18	2.30 ± 0.27	-2.70 ± 0.21 -2.97 + 0.28	4 +	2 68 + 0 50	4.00	3.80	27.42 27.52	מ מ
		-27.8	19.0			+ +	+	+ +	+	3.11 + 0.51	1.84	3.10	7.01	o vo
		8	0 0	411 0614±09		264 + 0.19	3 24 + 0 24	$-3.41 \pm 0.21$	+	3 84 + 0 42	48	90.8	0.83	o o
		-52.3	12.4	-	· -	$2.67 \pm 0.26$	$2.17 \pm 0.21$	+ +	+	3.38 + 0.60	1 12	7.44	9.17	o vo
		-10.1	20.2			$2.62 \pm 0.26$	$2.14 \pm 0.21$	$-3.03 \pm 0.25$	+	$2.62 \pm 0.43$	1.08	9.33	10.73	S CO
		-0.8	7.0			$2.18 \pm 0.10$	$5.35 \pm 0.24$	4.46 ±	1	$5.80 \pm 0.32$	96.9	31.00	39.12	w
	CA		20.0		$4.17 \pm 0.44$	$2.55 \pm 0.31$	$2.07 \pm 0.25$	+	+	$2.74 \pm 0.73$	1.21	4.83	6.31	Ø
			4.5			$2.449 \pm 0.088$	$10.99 \pm 0.39$	1	1	$12.27 \pm 0.67$	24.59	8.84	34.53	M
		-31.3	21.0		$2.92 \pm 0.22$	$1.72 \pm 0.15$	$3.50 \pm 0.32$	3.51 ±	+	$3.77 \pm 0.52$	1.09	10.02	11.49	M
11032483901 32656707		-29.1	12.8		$3.15 \pm 0.25$	$2.63 \pm 0.25$	$2.15 \pm 0.20$		+	$2.45 \pm 0.37$	3.96	2.94	9.31	Ø
11032703749 32835766		-2.5	11.1			$3.11 \pm 0.12$	+	$-3.39 \pm 0.11$	+	$6.05 \pm 0.37$	12.77	3.98	17.74	M
11032715596 32847619	8.908 6	3.0	16.0		$3.05 \pm 0.41$	$0.89 \pm 0.14$	$2.56 \pm 0.41$	$-3.08 \pm 0.42$	$1.11 \pm 0.24$	$3.19 \pm 0.68$	11.99	4.00	17.57	M
11032760330 32892324	4 286.9	20.1	11.6		$2.79 \pm 0.20$	$1.64 \pm 0.14$	$3.35 \pm 0.29$	$-3.33 \pm 0.23$	$1.94 \pm 0.22$	$3.97 \pm 0.45$	11.41	2.09	15.28	S
11032765184 32897188		5.1	7.7		$2.43 \pm 0.10$	$2.31 \pm 0.12$	$5.65 \pm 0.30$	$-3.69 \pm 0.15$	$2.66 \pm 0.19$	$6.53 \pm 0.47$	21.28	4.84	30.17	w
11032902720 33007524	4 245.7	9.2	16.4	$UW_{-}Crb$ (1.0)	$3.73 \pm 0.76$	$1.36 \pm 0.31$	$1.67 \pm 0.38$	$-3.06 \pm 0.69$	$1.58 \pm 0.68$	$1.93 \pm 0.84$	0.91	8.07	9.12	S
				SAX_J1818.7+1424 (1.8)	-									i
			6.5			+	$3.82 \pm 0.34$	$-3.19 \pm 0.24$	+	$4.71 \pm 0.67$	5.98	2.99	10.59	w
			4.7				$6.16 \pm 0.27$	$-3.12 \pm 0.11$	₩.	$7.02 \pm 0.39$	1.59	15.99	17.96	M
			5.5 5.0		$4.64 \pm 0.15$		$13.50 \pm 0.51$	H -	<b>H</b> -	$22.2 \pm 1.9$	34.15	65.98	104.93	so c
			o.x.			H -	2.47 ± 0.18	Н-		3.15 ± 0.39	5.40	5.I.	12.20	מ כ
11040323314 33460103	280.0	v. 0.	10.3		2.93 ± 0.18	1.99 ± 0.14	4.87 ± 0.36	$-3.34 \pm 0.21$	2.30 ± 0.20	5.78 H 0.65	1.38	15.91	16.47	M
		0.12	5.1.5 8.0	28 0918-549	3.13 ± 0.16		$3.01 \pm 0.22$ $17.77 \pm 0.49$	H +	H +	4.40 H 0.40	36.77	33.94	10.01	Z V
		-40.4	0.81			1 +	2 39 + 0 20	1 +	+	3 18 + 0 48	3 90	12.72	9.43	o o
		-28.2	13.3			1 +1	$3.27 \pm 0.30$	$-3.09 \pm 0.23$	1 +1	$4.18 \pm 0.62$	2.03	7.82	10.37	M
	4 238.0	-43.6	10.2		0	$1.91 \pm 0.18$	+	+	+	$4.85 \pm 0.84$	3.82	7.94	12.39	S
11040940928 33996128	8 263.7	-27.8	5.4			$3.61 \pm 0.24$	+	#	+	+	2.12	10.85	14.41	S
		-30.8	17.0			$3.85 \pm 0.28$	#	$-3.56 \pm 0.22$	+	$3.53 \pm 0.35$	3.14	4.85	10.54	S
		-22.1	8.3		$3.11 \pm 0.21$	$1.79 \pm 0.14$	$3.66 \pm 0.30$	$-3.36 \pm 0.23$	+ -	$4.08 \pm 0.42$	8.51	15.50	26.44	so o
		-33.1	20.2			$1.37 \pm 0.26$	Н-	$-3.06 \pm 0.49$	H -	3.08 ± 0.87	2.45	2.57	5.99	Ωō
		1.58.1	5.7			H -	H -	$-2.65 \pm 0.14$	$4.01 \pm 0.53$	$4.92 \pm 0.65$	4.20	9.46	16.33	ΩŞ
11041129388 34157398	3 359.0	25.4	20.1		$3.53 \pm 0.25$ $2.61 \pm 0.26$	2.23 ± 0.18	5.48 ± 0.45	$-3.06 \pm 0.21$	2.57 ± 0.27	0.29 ± 0.67	1.60 2.16	14.88	10.94	Z v
		4.00	2.0		3.0 H 0.20	ΗН	ΗН	-2.01 H 0.21	ΗН	9.03 ± 0.04	1.10	40.7	0.91	מ מ
		1 1 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	1.01 1.12		· -	+	+ +	4 +	+	3.03 H 0.40	1.90 7.85	4.04	19.78	n o
	_	33.8	. o		$3.17 \pm 0.23$	+	2.26 ± 0.19	$-3.24 \pm 0.24$	+	$2.57 \pm 0.35$		4 25	6.04	o v
		-23.4	13.1		0	1 +1	1 +1	$-3.16 \pm 0.20$	1 +1	1 +1	1.95	4.36	7.44	ı w
11042255142 35133547	7 237.2	-47.5	17.6		$3.38 \pm 0.23$	$3.97 \pm 0.31$	$3.24 \pm 0.25$	$-3.16 \pm 0.21$	$4.79 \pm 0.60$	$3.91 \pm 0.49$	2.66	4.84	8.05	w
		-25.0	15.4		0	+	$1.64 \pm 0.24$	+	$\mathbb{H}$	$2.88 \pm 0.87$	1.68	4.55	6.64	SO
11042452346 35303536	3 93.2	14.3	3.4	4U_0614+09	$3.057 \pm 0.064$	$4.95\pm0.12$	$14.15\pm0.35$	$-3.271 \pm 0.071$	$5.93 \pm 0.23$	$16.95 \pm 0.67$	6.01	13.31	20.05	w

Table 5:: GBM Type 1 Events continued from previous page

								i		į	i	i			
OI.	Feak I	Ka D	Dec	Error	Name(distance) (sigma)	$egin{array}{c}  ext{EB} &  ext{temp} \  ext{keV} \end{array}$	$10-8 \text{ erg cm}^{-2} \text{ s}^{-1}$	$_{10^{-7} m erg~cm^{-2}}$	PL index	$10^{-8} \text{erg cm}^{-2} \text{s}^{-1}$	$_{10^{-7}}^{\rm Flnc}$	Kise	Fall	Duration Sec	Structure
	35431531 34	347.4 -48		10.4		$3.39 \pm 0.25$		$6.80 \pm 0.57$	$-3.42 \pm 0.25$	$9.6 \pm 1.0$	$7.87 \pm 0.82$	2.95	3.51	6.99	S
				13.4		$2.65 \pm 0.15$	$2.05 \pm 0.15$	+	$-3.43 \pm 0.19$	$2.42 \pm 0.26$	$6.92 \pm 0.76$	12.91	80.9	19.53	M
				8.9		$4.22 \pm 0.29$		+	$-2.50\pm0.17$	$4.39 \pm 0.82$	$5.3 \pm 1.0$	6.85	3.75	11.47	S
		~		10.2			₩.	$2.12 \pm 0.26$	$-2.82 \pm 0.28$	₩.	$3.24 \pm 0.69$	1.05	12.83	14.33	M
11042827941 350	35624740 6			80 d 4. 4	0710 000011 0001	$2.37 \pm 0.13$	$1.46 \pm 0.10$	$5.98 \pm 0.44$	$-3.72 \pm 0.20$	$1.77 \pm 0.18$	$7.23 \pm 0.74$	9.91	8.15	19.53	w c
		255.9 -39	-39.9	4.7	1Gr.J.11062-0143	$2.22 \pm 0.21$ $2.95 \pm 0.23$	Н Н	$2.30 \pm 0.22$	$-2.91 \pm 0.19$ $-3.20 \pm 0.24$	$1.92 \pm 0.20$ $3.49 \pm 0.52$	$4.72 \pm 0.30$ $2.85 \pm 0.42$	4.45	5.58	13.52	o co
				5.7	$4U_0614+09$		1 +1	$10.64 \pm 0.32$	$-3.094 \pm 0.079$	+	$13.58 \pm 0.69$	5.08	7.97	14.32	o o
11043075354 358	35844951 27	278.2 -18	-18.8	5.3		$3.38 \pm 0.19$	$2.36 \pm 0.16$	$5.02 \pm 0.34$	$-3.10\pm0.18$	$2.71 \pm 0.32$	$5.75 \pm 0.68$	7.16	4.08	11.80	M
				1.0			+	$24.65 \pm 0.59$	$-2.952 \pm 0.060$	$\mathbb{H}$	$28.50 \pm 0.90$	8.56	20.51	42.90	w
				9.4			+	$7.28 \pm 0.86$	+	+	$13.8 \pm 3.5$	2.49	8.90	17.95	M
			•	5.7			$1.80 \pm 0.16$	$4.41 \pm 0.41$	$-2.60 \pm 0.19$	+	$7.3 \pm 1.3$	1.68	7.61	17.49	M
				8.0	4U0614+09		$3.79 \pm 0.18$	$4.64 \pm 0.22$	$-3.20 \pm 0.13$	+	$5.63 \pm 0.45$	1.42	5.98	7.60	W
				8.6			$3.21 \pm 0.21$	$3.93 \pm 0.25$	$-3.27 \pm 0.17$	+1 -	$4.64 \pm 0.47$	8.29	5.94	15.16	M
				9.5		o o	H -	+ -	$-2.38 \pm 0.17$	₩-	$5.40 \pm 0.58$	22.63	22.98	54.80	w t
				4.2			$1.91 \pm 0.12$	$9.39 \pm 0.59$	$-2.86 \pm 0.15$	<b>H</b> -	$11.17 \pm 0.88$	20.06	65.06	80.38	w c
				13.3		o 0	₩-	$3.73 \pm 0.29$	H -	H -	$4.60 \pm 0.44$	8.56	80.00	13.27	n c
11051230231 303	30830025 ZE	209.9 - 5	-30.0 1	10.1		$2.92 \pm 0.20$	1.90 ± 0.15	3.88 H 0.31	-3.30 ± 0.22	2.29 # 0.27	4.08 H 0.50	14.75 8 EO	11 53	20.19	מ מ
				11:2			2:10 ± 0:14 3:57 ± 0:14	4.30 + 0.29	-3.38 ± 0.19	4 +	10 68 ± 0 66	0.00	15.00	10.07	ס מ
			_	. o		0	+ +	6.18 + 0.24	$-3.42 \pm 0.09$	+	7.17 + 0.40	11.40	8.64	21.25	Σ
				2 00	2S 0918-549		$2.04 \pm 0.16$	10.00 + 0.81	$-2.87 \pm 0.20$	+	13.7 + 2.1	15.13	25.90	44.15	v.
				5.4			$3.29 \pm 0.21$	$4.03 \pm 0.26$	$-3.66 \pm 0.18$	+	$4.54 \pm 0.40$	1.61	9.04	12.52	o vo
				12.2			$1.82 \pm 0.18$	$3.72 \pm 0.38$	$-2.55 \pm 0.19$	+	$5.07 \pm 0.60$	15.55	7.66	24.99	S
			٠.	20.1			$3.01 \pm 0.23$	$2.45 \pm 0.19$	$-3.24 \pm 0.21$	+	$2.89 \pm 0.35$	2.06	3.11	7.57	w
_				16.7			$2.13 \pm 0.19$	$2.61 \pm 0.23$	$-2.81\pm0.21$	+	$3.57 \pm 0.59$	2.46	3.67	6.97	Ø
11052531507 379	37961110 28	253.1 -3'	-37.5 1	11.9		$3.50 \pm 0.30$	$2.36 \pm 0.24$	$2.89 \pm 0.30$	$-2.76 \pm 0.23$	$3.39 \pm 0.67$	$4.15 \pm 0.82$	1.44	13.44	15.32	M
				10.8		0	+	$2.32 \pm 0.22$	+	+	$2.82 \pm 0.44$	1.71	9.76	11.84	w
				6.8			$2.24 \pm 0.16$	$3.66 \pm 0.26$	$-3.03 \pm 0.18$	+	$4.60 \pm 0.56$	2.80	8.40	12.42	w
				8.3		o.	+	$36.0 \pm 1.2$	+1 -	+	$52.4 \pm 3.8$	50.41	72.97	142.49	w.
				20.5				$4.98 \pm 0.37$	41 -	₩ -	$5.64 \pm 0.70$	4.13	12.08	17.05	w i
				7.1		$3.73 \pm 0.31$	$4.09 \pm 0.40$	$1.67 \pm 0.16$	$-2.86 \pm 0.22$	H -	$2.28 \pm 0.37$	2.11	5.09	8.50	w ;
				9.7.		$3.33 \pm 0.17$	H -	H -	$-2.96 \pm 0.15$	H -	$5.18 \pm 0.40$	3.32	14.06	23.77	N N
11060326889 38	38734086 24	248.5 -50	-56.2	11.6		$2.99 \pm 0.14$ $2.60 \pm 0.13$	$1.90 \pm 0.10$	4.66 ± 0.26	-3.18 # 0.15	$2.35 \pm 0.23$	5.76 ± 0.57	2.93	48.7	11.31	n u
				9.I		2.00 H 0.12	H +	H +	H +	H +	4.28 H 0.33	2.59 2.30	6.40	9.37	מ מ
				10.0		Ö	+ +	++	-3.11 + 0.34	++	$3.79 \pm 0.75$	5.03	68.9	12.58	o o
				16.3				1	$-2.89 \pm 0.36$	1	$2.34 \pm 0.60$	2.30	7.40	10.03	w
				4.0		+ 0	+	$^{\rm H}$	$-2.95\pm0.16$	$^{\rm H}$	$6.20 \pm 0.73$	10.62	3.55	15.19	M
				7.4			+	+	+1	#	$9.0 \pm 1.3$	34.18	7.25	42.21	M
				28.5		$3.44 \pm 0.33$	$5.13 \pm 0.54$	₩-	$-3.00 \pm 0.28$	H -	$11.1 \pm 1.7$	1.91	7.19	13.30	M
11060767901 39	39044159 28	257.3 -4.	242.1	0.21		2.94 ± 0.11	3.50 ± 0.16	$4.29 \pm 0.20$	$-3.24 \pm 0.12$ $-3.99 \pm 0.18$	4.27 ± 0.32 3.08 ± 0.38	9.23 H 0.40	9.53	9.79	14.63	ก บ
	_			7 c			+ +	++	+ +	++	4.06 + 0.42	2.5	3.72	9 41	2 ≥
				25.2		0	1 +1	+	1 +1	+	$0.93 \pm 0.21$	2.60	5.39	8.79	w
				11.0			+	+	+	+	$3.38 \pm 0.29$	1.71	7.13	9.23	w
11061046887 393	39358889 26	263.8 -4	-4.0 1	11.4		$3.31 \pm 0.14$	$3.69 \pm 0.19$	$^{\rm H}$	$-3.08 \pm 0.14$	$4.38 \pm 0.40$	$7.14 \pm 0.65$	8.64	7.90	17.29	Ø
				10.6		0	+	$^{\rm H}$		$^{\rm H}$	+	2.55	8.11	11.91	w
				8.9			+ -	+ +	$-3.64 \pm 0.21$	+ -	$2.93 \pm 0.30$	2.83	4.28	7.71	w i
				22.5		0 0	H -	₩ -	H -	H -	₩-	1.79	10.11	12.33	v c
11062250752 400	40399557 28	281.2 -20	-26.3	0.0		$3.46 \pm 0.19$	$2.97 \pm 0.19$	$3.64 \pm 0.24$	$-3.00 \pm 0.17$	$3.72 \pm 0.44$	$4.56 \pm 0.53$	9.54	5.12	15.00	n o
				g: -[			H +	4.48 ± 0.21 5.23 ± 0.36	H +	H +	5.54 H 0.40 6.30 + 0.74	1, 0 1, 0 1, 0	60.0	10.00	ט ט
				13.1		$3.13 \pm 0.19$	+	+	+	+	$4.29 \pm 0.52$	1.31	11.55	13.17	Σ
				7.8		$3.38 \pm 0.17$	+	+	$-3.07 \pm 0.18$	+	+	1.69	12.21	14.33	w
11062862784 409	40930001 28	281.3 -1'	-17.9	7.8		$3.32 \pm 0.19$	$2.12 \pm 0.15$	$4.34 \pm 0.30$	$-3.04 \pm 0.18$	$2.61 \pm 0.32$	$5.33 \pm 0.66$	9.72	4.68	14.82	S
11062870080 409	40937291 21	216.2 -4	-41.1	5.4		$3.07 \pm 0.13$	$3.99 \pm 0.20$	$4.88 \pm 0.24$	$-3.40\pm0.14$	$4.49 \pm 0.33$	$5.49 \pm 0.41$	2.42	11.36	15.04	W

Table 5:: GBM Type 1 Events continued from previous page

GI	Peak	Ka I	Dec	Error	Name(distance) (sigma)	$^{ m keV}$	$^{10-8} {\rm erg} {\rm cm}^{-2} {\rm s}^{-1}$	$^{1}$ BB Flnc $^{10-7}$ erg cm $^{-2}$	PL index	PL Flux 10 <sup>-8</sup> erg cm <sup>-2</sup> s <sup>-</sup>	$^{1}$ PL Flnc $^{1}$ $^{10-7}$ erg cm $^{-2}$	Rise	Fall	Duration S	Structure
11063045089 41	41085085 2	5- 0.966	-33.0	6.7	(mingra)	4 10 ± 0 33	- 2 2 2 3 3	3 80 + 0 34	$-284 \pm 0.16$	4 20	70	4 36	4 14	11.26	ŭ
				12.7		$2.84 \pm 0.14$	+ +	$2.94 \pm 0.18$	$-3.38 \pm 0.17$	+ +	$3.41 \pm 0.33$	4.44	11.38	16.39	o o
				12.4		$3.77 \pm 0.28$		$3.31 \pm 0.30$	1	1	$4.86 \pm 0.93$	2.72	10.11	13.48	S
11070350849 41	41350057 2		-39.6	17.8		$2.94 \pm 0.25$		$0.278 \pm 0.088$	$-3.35 \pm 0.28$	$0.86 \pm 0.24$	$0.70 \pm 0.20$	2.21	1.90	5.16	Ø
				1.0		$3.70 \pm 0.11$	+	$22.41 \pm 0.81$	.837 ±	+	$28.9 \pm 2.0$	26.53	36.06	64.66	SO
11070380162 41	41379373 20	263.4 -4	-45.0	11.6		$4.17 \pm 0.34$	$2.97 \pm 0.28$	$2.42 \pm 0.23$	$-2.72 \pm 0.24$		$3.04 \pm 0.64$	2.95	4.80	8.21	o o
				7.2		$2.97 \pm 0.29$	H +	H +	H +	2.48 + 0.18	H +	2 × 0	0 0 0	13.08	o oo
			-18.8	. 8		$3.08 \pm 0.11$	+ +	+	1 +1	+	1 +1	8.69	9.93	19.45	Ω M
11070924160 41		251.2 -		21.0		$2.74 \pm 0.20$		+	$-3.53 \pm 0.28$		+	5.54	7.72	14.01	w
11071012340 41	41916348 20	266.0 -2	-22.3	12.7		$3.51 \pm 0.27$		$3.26 \pm 0.30$	$-3.06 \pm 0.25$	+	$3.80 \pm 0.62$	6.62	3.31	12.79	S
				16.1		$3.61 \pm 0.30$	+	+	$-2.75\pm0.23$	+	+	4.12	11.98	16.61	S
11071102902 41		277.8 -1	-14.5	9.9		$3.95\pm0.31$	+	$4.37 \pm 0.38$	$-2.67\pm0.19$	$2.23 \pm 0.23$	+	1.30	13.66	15.32	M
				9.6			+	#	#	+	+	2.79	11.56	15.11	M
				15.0		$3.25 \pm 0.20$	₩.	₩.	₩.	₩.	+ -	2.62	5.55	8.60	w i
				5.0		$3.20 \pm 0.17$	₩ -	H -	H -	<b>H</b> -	₩ -	10.67	9.25	20.51	w c
11071264792 42	42141593 23	255.5 -	-4.7 5.5 E	14.6		$2.73 \pm 0.17$	$3.32 \pm 0.24$	$5.43 \pm 0.39$	$-3.81 \pm 0.23$	$3.51 \pm 0.33$	$5.73 \pm 0.54$	3.18 1.28	9.22	14.49	νo
			55.5	2.0		3.27 + 0.20	+	+	+ +	2.89 ± 0.40	+	2.93	2.2	11.79	2 CC
				10.2			+	+	$-2.98 \pm 0.18$	+	+	2.47	5.44	10.36	ı va
				6.5		Ö	1	1	1 +1	1	1	3.24	16.84	20.70	o so
11071529279 42	42365304 29	295.7 -1	-11.7	17.0		$4.03 \pm 0.37$		$2.94 \pm 0.32$	$-2.67 \pm 0.23$	$3.28 \pm 0.69$	$4.01 \pm 0.84$	2.65	6.84	9.87	Ø
				20.3			+	+	$-3.01 \pm 0.26$	+	+	1.96	5.79	86.8	w
			-12.7	14.8			+	+	+	$2.40 \pm 0.27$	$3.92 \pm 0.44$	1.54	15.12	16.95	M
				10.0				+	+	+	$4.30 \pm 0.67$	10.79	7.13	18.29	W
				6.1			+	+1 -	+	+1 -	+1 -	1.57	5.42	9.26	w.
				21.3			H -	<b>H</b> -	<b>H</b> -	₩-	$3.70 \pm 0.72$	2.63	7.48	10.61	w o
			-37.6	10.6		$2.71 \pm 0.16$	H -	H -	$-3.30 \pm 0.20$	H -	H -	2.23	15.42	19.62	n c
110/183203/ 42	42627840 2.	228.7 -6 257.1 -6		13.0		3.52 H 0.35	1.24 ± 0.15	3.30 H 0.43	$-2.74 \pm 0.25$	1.78 ± 0.40	0.1 H 1.1	0.41 14.0	20.05	10.67	מ מ
				13.1			H +	H +	H +	H +	3.24 + 0.38	1.00	20.02	8 74	מ מ
				. s. s		0	+	1 +1	1 +1	1 +1	1 +1	1.54	3.73	5.62	ω
				21.7				+	$-3.17 \pm 0.33$	+	$3.37 \pm 0.71$	4.19	26.71	32.59	S
	_			18.9			+	.77 ±	+	$^{+}$	+	8.60	4.22	13.81	S
				9.7			₩ -	₩.	+1 -	+ -	₩.	1.60	9.67	21.63	Μ;
				11.9			H -	$4.56 \pm 0.35$	H -	H -	$6.61 \pm 0.99$	1.46	16.38	18.19	Μū
11072412187 43	43125788 20	209.0 -	-33.5	19.6		3.71 ± 0.31	$0.45 \pm 0.10$	0.73 ± 0.16	$-2.80 \pm 0.22$	$0.97 \pm 0.20$	1.59 ± 0.32 8 00 ± 0 07	1.70	9.54 19.7E	13.00	ΩŞ
				17.1		2.94 + 0.19	+ +	+ +	$-3.25 \pm 0.22$	+	++	2.13	4.66	8.07	Z v
				8.6			1 +1	1	1 +1	+	+	1.65	13.70	15.72	o oo
				9.6		o.	+	+	+	+	+	4.03	2.87	7.73	S
			_	14.9			+ -	₩.	+ -	+ -	+ +	7.38	12.87	22.23	M
11073009704 43	43641712 28	280.7	0.0 " "	7.7		$3.74 \pm 0.41$	$0.94 \pm 0.12$	$3.09 \pm 0.40$	$-2.59 \pm 0.27$	$1.21 \pm 0.19$	$3.97 \pm 0.64$	2.47	14.24	17.38	s >
				14.0		$3.23 \pm 0.14$	H +	H +	H +	$4.69 \pm 0.40$ $3.12 \pm 0.40$	H +	1.67	3.96	5.94	Z v
				5.6				1	1	1	+	1.49	6.18	13.96	M
11080109379 43	43814180 20	260.8 -4	-40.7	9.2			$2.98 \pm 0.21$	$2.43 \pm 0.17$	$-3.01 \pm 0.18$	$3.73 \pm 0.44$	$3.04 \pm 0.36$	12.43	7.32	19.85	M
				0.6			+	+	+	+	+	8.77	88.9	16.34	M
				12.4	0		+ -	# -	+1 -	+1 -	# -	2.14	3.77	6.95	w o
				8.0	28_0918-549	<u> </u>	H -	H -	H -	H -	H -	8.25	15.58	24.47	n o
11080273568 43	43964759 20	201.02	-32.0	10.4		3.13 ± 0.28	$1.87 \pm 0.20$	$2.29 \pm 0.24$ 2 80 $\pm$ 0 36	$-3.28 \pm 0.30$	$2.11 \pm 0.36$ 1 64 $\pm$ 0.33	$2.59 \pm 0.44$ $2.25 \pm 0.48$	1.57	7.13	9.04	ΩŞ
				1 0			+ +	+	+	+	+	7.16	266	17.79	ď
				16.3		0	+	1	$-3.23 \pm 0.16$	1	+	1.24	6.87	14.58	M
11080460580 44		298.5 -		18.2			$4.83 \pm 0.39$	$1.97 \pm 0.15$	$-3.15\pm0.21$	$5.88 \pm 0.73$	$2.40 \pm 0.30$	1.21	00.9	7.71	w
				10.1		0	+ -	+ +		+ -	+ +	1.41	4.99	6.64	Ω i
11080521263 44	44171667 2′	275.9 -	-55 8.8 8.0	16.1		$2.67 \pm 0.23$	$1.18 \pm 0.13$	$1.92 \pm 0.21$	$-3.42\pm0.31$	$1.40 \pm 0.24$	$2.28 \pm 0.39$	3.16	16.38	20.06	w

Table 5:: GBM Type 1 Events continued from previous page

E	-	Ė			( ); ()	ני	9 00	10.00		ī	IG IG				
3	reak	Ρά	Dec	Error	(sigma)	рь сешр keV	$10-8 \text{ erg cm}^{-2} \text{ s}^{-1}$	$10^{-7} \mathrm{erg~cm}^{-2}$	r L maex	10 <sup>-8</sup> erg cm <sup>-2</sup> s <sup>-1</sup>	$10^{-7} \text{ erg cm}^{-2}$	sec	sec	Duration 5	Structure
11080870913 4	44480512	283.2	-19.5	9.6		$2.96 \pm 0.14$	$2.53 \pm 0.14$	$4.13 \pm 0.23$	$-3.36 \pm 0.16$	$2.87 \pm 0.25$	$4.69 \pm 0.42$	6.95	6.77	14.25	M
			-28.5	7.1		$3.00 \pm 0.15$	$2.33 \pm 0.14$	$3.81 \pm 0.23$	$-3.29 \pm 0.17$	$2.77 \pm 0.26$	$4.52 \pm 0.43$	2.97	12.84	16.24	w
11081446161 44	44974171	182.1	-24.8	12.6		$2.93 \pm 0.34$	$1.10 \pm 0.14$	+	$-3.32 \pm 0.39$	$1.29 \pm 0.22$	$4.24 \pm 0.72$	2.36	7.99	11.31	w
11081572794 45	45087184	247.7	-32.2	11.7		$3.45 \pm 0.30$	$2.28 \pm 0.22$	$1.86 \pm 0.18$	2.99	$2.96 \pm 0.49$	$2.41 \pm 0.40$	1.44	4.07	7.55	S
	45151795		-13.9	12.3		$3.40\pm0.25$	$2.79 \pm 0.25$	+	2.71	+	$5.09 \pm 0.90$	7.03	6.35	13.92	Q
			-15.1	8.0		$2.86 \pm 0.12$	$3.08 \pm 0.17$	$6.29 \pm 0.35$	3.30 ±	+	$7.39 \pm 0.66$	1.72	6.29	18.30	M
_			-16.1	8. 8.		$3.69 \pm 0.21$	$2.95 \pm 0.20$	+1	2.82	+	$4.86 \pm 0.69$	7.20	5.03	13.25	w.
			-51.6	6.2	$2S_{-}0918-549$	$2.863 \pm 0.088$	$2.461 \pm 0.092$	+	+	+	$11.46 \pm 0.71$	10.03	26.59	38.75	w
			-52.3	3.7		$2.69 \pm 0.14$	$2.88 \pm 0.19$	+	$-3.46 \pm 0.19$	$3.29 \pm 0.32$	$4.03 \pm 0.40$	6.81	11.62	20.06	W
			-31.1	9.2		$3.62 \pm 0.27$	$2.70 \pm 0.24$	$3.30 \pm 0.30$	$-2.80 \pm 0.21$	$3.68 \pm 0.66$	$4.50 \pm 0.81$	7.70	89.9	15.15	M
11082307171 48	45712763	287.2	10.3	6.4	Ser_X-1 (1.4)	$2.82 \pm 0.14$	$1.98 \pm 0.12$	$5.67 \pm 0.35$	$-3.38 \pm 0.18$	$2.25 \pm 0.18$	$6.45 \pm 0.53$	24.00	17.82	48.28	ω
				Š	MXB_1906+00 (1.6) Swift_J185003.2-005627 (										
11082319464 48	45725063	284.6	11.9	8.0		$2.61 \pm 0.11$	$2.05 \pm 0.10$	$7.54 \pm 0.39$	$-3.48 \pm 0.14$	$2.41 \pm 0.16$	$8.85 \pm 0.59$	28.81	3.45	33.59	Μ
			-17.7	6.2		$3.81 \pm 0.28$	$2.70 \pm 0.24$		$-2.81 \pm 0.21$	$3.07 \pm 0.38$	+	5.74	4.81	12.00	S
11082424294 45	45816297		-17.0	9.5		$3.53 \pm 0.25$	$1.80 \pm 0.15$	+	$-3.05 \pm 0.22$	+		6.33	5.57	14.58	w
11082505691 45	45884088		-21.5	15.1		$4.16 \pm 0.36$			$-2.80 \pm 0.24$	$2.16 \pm 0.41$	+	2.84	13.15	16.41	W
			-42.4	8.7		$3.43 \pm 0.26$	+	+	$2.96 \pm$	+	$5.39 \pm 0.87$	3.29	9.42	13.43	w
			-21.4	11.1		Ö	+	+	41	+	+	6.54	8.20	14.95	w
			-11.7	7.5			₩.	+ -	$-3.21 \pm 0.13$	+ +	$7.12 \pm 0.56$	89.6	92.9	17.20	M
			-17.2	11.7			₩ -	+ -	41 -	+ -	$4.17 \pm 0.65$	6.73	4.73	12.33	w o
			19.3	8.4			₩ -	₩ -	$-3.36 \pm 0.11$	Н-	$7.05 \pm 0.42$	2.83	12.16	15.55	n ;
			-14.7	11.3			$2.19 \pm 0.17$	$3.59 \pm 0.28$	$-3.42 \pm 0.22$	H -	$4.36 \pm 0.50$	1.21	4.69	14.33	M ;
			-18.0	5.7		2.88 ± 0.15	Н-	Н-	$-3.22 \pm 0.17$	Н-	8.92 ± 0.78	5.14	15.49	28.00	M ;
11090121122 46		27.7.3	-15.0	16.2		$2.92 \pm 0.16$	$2.21 \pm 0.14$	$3.61 \pm 0.24$	$-3.32 \pm 0.18$	$2.60 \pm 0.27$	$4.25 \pm 0.45$	79.0	9.09	18.60	M C
			14.7	0.7			2.08 \pm 0.12	Н-	Н-	Н-	5.30 H 0.48		12.30	18.07	o c
11090226964 40	46536555	240.4 25.4	0.7.0 0.7.0	14.1		3.00 ± 0.19	2.92 ± 0.23	H +	-3.10 H 0.19	3.05 H 0.45	3.00 H 0.39		10 5k	9.91 28.80	ΩU
			-31.4	12.8			2.23 + 0.20	+	$-3.65 \pm 0.19$ $-2.65 \pm 0.20$	+	5.15 + 0.99		7.79	15.48	n ≥
			-32.3	4.2		0	+	+	1 +	+	$10.53 \pm 0.73$	_	60.9	17.33	M
11090328160 46	46684167	278.8	-8.9	9.7		$2.95 \pm 0.10$		$6.06 \pm 0.26$	$-3.32\pm0.12$	+	$7.13 \pm 0.50$	11.55	7.50	19.98	M
	46726953	283.3	-29.7	14.1		Ö	+	$3.01 \pm 0.23$	$-3.14 \pm 0.20$	+	$3.75 \pm 0.46$	6.82	4.96	12.41	M
			-74.8	14.1			+	$1.91 \pm 0.24$	#	+	$2.46 \pm 0.40$	4.27	4.43	9.56	w
_			68.2	16.8		$4.01 \pm 0.54$	₩.	$1.40 \pm 0.21$	$-2.74 \pm 0.36$	$2.48 \pm 0.75$	$2.02 \pm 0.61$	3.82	5.68	9.97	w ¦
			-32.6	9.5			₩-	$3.34 \pm 0.37$	# -	$2.25 \pm 0.43$	$4.58 \pm 0.88$	8.98	7.01	17.02	M o
11090609290 46	46924491	237.8	- 35.5 E E	× 5		$3.16 \pm 0.12$	2.16 ± 0.10	$13.27 \pm 0.65$	$-3.07 \pm 0.12$	$2.67 \pm 0.22$	$16.4 \pm 1.3$	21.63	23.48	45.98	n >
			280	24.3			+ +	3.70 + 0.38	4 +	280 + 085	4.00 ± 0.91 6.0 ± 1.3	1.61	19 97	15.00	E D
			-45.8	15.1		0	+	$2.27 \pm 0.20$	1 +	1 +	$2.92 \pm 0.42$	6.18	5.50	14.15	S
			-14.6	12.3			$2.32 \pm 0.25$	$2.84 \pm 0.31$	1	Н	$3.70 \pm 0.85$	5.32	6.45	12.56	M
			-16.8	11.2		0	+	$5.65 \pm 0.67$	#	$\mathbb{H}$	$7.6\pm1.8$	_	5.28	18.76	M
			2.2	5.4		$2.56 \pm 0.12$	$2.06 \pm 0.12$	$5.04 \pm 0.30$	+1	$2.57 \pm 0.23$	$6.30 \pm 0.57$		20.13	33.90	W
			-62.3	io i			$2.22 \pm 0.10$	$19.09 \pm 0.94$	41 -	+ -	$28.2 \pm 3.1$		44.52	53.07	M
			-39.6	5.7		$2.96 \pm 0.14$	$2.07 \pm 0.12$	$5.08 \pm 0.30$	$-3.21 \pm 0.15$	$2.49 \pm 0.24$	$6.11 \pm 0.58$	6.54	9.00	17.66	ωo
11091104629 44	47204405	249.4	0.7 0.0	11.4			1.21 H 0.14	3.47 H 0.42	0.20 H 0.20	1.09 H 0.40	4.0 H 1.1		0 u	19.50	מ מ
			-25.1	0.7		$3.78 \pm 0.22$	H +	4.98 + 0.48	$-2.84 \pm 0.24$	3.03 + 0.58	6.1 + 1.1		5.05	9.64	2 OC
			-28.1	4.3			$1.61 \pm 0.12$	$7.25 \pm 0.54$	$-3.37 \pm 0.21$	$1.92 \pm 0.22$	$8.65 \pm 0.99$		38.16	41.35	M
	47739234		-34.2	7.1		$3.29 \pm 0.20$	$2.10 \pm 0.15$	$3.43 \pm 0.24$	$-3.02 \pm 0.18$	$2.71 \pm 0.33$	$4.42 \pm 0.55$	8.47	7.22	16.32	M
11091616566 47	47795775	243.8	-42.2	8.1		0	$1.15 \pm 0.12$	$3.29 \pm 0.35$	$-2.71 \pm 0.24$	$1.73 \pm 0.35$	$4.9 \pm 1.0$	27.60	6.97	34.85	M
			-43.1	8.9		$3.17 \pm 0.14$	$3.03 \pm 0.17$	$3.72 \pm 0.20$	$-3.11\pm0.14$	$3.74\pm0.34$	$4.58 \pm 0.42$	4.21	9.04	14.35	w
			-56.0	12.5			$3.97 \pm 0.18$	$3.24 \pm 0.15$	$-3.12 \pm 0.12$	$4.75 \pm 0.37$	$3.88 \pm 0.30$	3.11	10.64	14.40	w
			-30.8	20.5			$1.92 \pm 0.20$	$1.57 \pm 0.16$		$2.61 \pm 0.49$	$2.13 \pm 0.40$	2.44	90.9	9.53	w.
			-69.3	24.5			$10.15 \pm 0.88$	$16.5 \pm 1.4$	$-4.64 \pm 0.33$	$11.17 \pm 0.94$	$18.2 \pm 1.5$	2.24	13.34	16.08	ω ;
			-20.0	2.1		$3.028 \pm 0.070$	4.15 ± 0.11	27.12 ± 0.75	.304 ±	$4.81 \pm 0.20$	$31.4 \pm 1.3$	13.09	50.32	74.04	M
11092081572 48	48206372	285.0	2.76-	10.4		3.02 ± 0.20 2.93 ± 0.17	3.57 ± 0.23	2.42 H 0.19	-3.28 ± 0.23	3.42 H 0.43 4 03 + 0 39	3.29 ± 0.35	1.7.1 9.65	3.69	9.00	מ מ
			)	1.		4	1	+	) 	+	+	i	5	1	2

Table 5:: GBM Type 1 Events continued from previous page

Ω	Peak	Ra D	Dec	Error	Name(distance)	${ m BB~temp}_{ m keV}$	BB flux 10-8 erg cm <sup>-2</sup> s <sup>-1</sup>	$^{10-7}_{ m erg~cm}^{-2}$	PL index	PL Flux 10 <sup>-8</sup> erg cm <sup>-2</sup> s <sup>-1</sup>	PL Flnc $10^{-7}$ erg cm <sup>-2</sup>	Rise	Fall	Duration S	Structure
11099084484 48	48209302	1- 9886	131	9 9	(property)	3 13 + 0 14	9 50	5 11 + 0 20	-311+015	3 10 ± 0 30	6 34 + 0.61	19.60	3 08	17.94	M
			- 21	20.0		$3.58 \pm 0.23$	$2.77 \pm 0.22$	4 +	$-2.92 \pm 0.20$	+ +	1 +1	2.68	8.14	11.65	E o
_				8.2		$3.26 \pm 0.20$	$2.24 \pm 0.16$	+	$-2.91 \pm 0.17$	+	$3.86 \pm 0.47$	4.08	13.79	18.67	S
11092354011 48	48438016 2		-37.2	4.9			$2.02 \pm 0.11$	$6.78 \pm 0.38$	$-3.21 \pm 0.15$	$2.33 \pm 0.16$	$7.80 \pm 0.55$	15.87	12.81	30.66	M
				24.8		$2.93 \pm 0.25$	$2.12 \pm 0.22$	$2.59 \pm 0.27$	$-3.24 \pm 0.29$	+	$2.98 \pm 0.50$	2.51	14.32	17.16	S
				6.3			+	$6.69 \pm 0.28$	3.33 ±	# -	$7.86 \pm 0.50$	3.02	26.29	29.94	M
				5.5			$2.22 \pm 0.16$	$2.72 \pm 0.20$	$-2.88 \pm 0.18$	<b>H</b> -	$3.65 \pm 0.47$	2.95	6.21	9.70	w o
				1.3			₩ -	$34.2 \pm 1.2$	$-3.372 \pm 0.097$	<b>H</b> -	$38.4 \pm 1.7$	26.17	47.10	90.94	so ;
11092867902 48	48883915 2		-53.9	2.7		$3.61 \pm 0.17$	$2.56 \pm 0.14$	$6.28 \pm 0.35$	$-2.95 \pm 0.14$	$3.24 \pm 0.32$	$7.93 \pm 0.80$	3.07	18.88	23.15	Ν
		2- C.447.		1.0.1		0.50 H 0.57	2.38 H 0.22	2.11 ± 0.18	-3.06 H 0.23	3.11 ± 0.46	2.34 H 0.37	17.7	1.99	19.20	מ מ
				14.3			H +	2.04 ± 0.21 6.01 ± 0.33	$-3.08 \pm 0.21$	H +	H +	1.00 3.77	10.40	10.39	ก บ
_			cc	4 - 1			ΗН	2 00 ± 0.33	9.14 101	НН		0.00	10.40	11.91	מ מ
				7.0			2.45 ± 0.17	H +	H 4	H +	5.95 H 0.44	0.03	14.07	18.43	ממ
				11.0			2.64 + 0.15	++	4 +	++	4 92 + 0 46	10 91	7.46	19.30	2 ≥
				9.4			$2.43 \pm 0.22$	+	+	+	3.44 + 0.40	3.09	5.03	9.42	, v
				4.7		0	+ 1	+	+	+	5.04 + 0.80	2.63	16.91	20.26	o oc
				7.6			$3.81 \pm 0.22$	1 +1	1 +1	1 +1	$3.71 \pm 0.32$	2.65	10.84	14.04	ı w
			-59.2	17.3		$3.17 \pm 0.15$	+	+	+	+	+	2.98	60.6	12.57	S
11100554159 49	49474962 2	244.8 -3	-35.6 1	16.4		$2.93 \pm 0.23$	$2.35 \pm 0.23$	$2.88 \pm 0.28$	$-3.22 \pm 0.26$	$2.76 \pm 0.45$	$3.37 \pm 0.56$	2.80	9.04	13.91	S
				8.3			$2.75 \pm 0.16$	$4.48\pm0.27$	#	$\mathbb{H}$	$5.00\pm0.42$	2.80	14.67	18.95	w
				16.1			$3.51 \pm 0.34$	$4.29 \pm 0.42$	+	+	$4.85 \pm 0.55$	2.70	9.75	13.25	w
				17.2			$2.33 \pm 0.28$	$4.75 \pm 0.57$	$-3.20 \pm 0.31$	+	$6.0 \pm 1.1$	2.69	10.85	14.30	w
				12.0			$2.40 \pm 0.19$	$2.93 \pm 0.24$	#	+	$3.33 \pm 0.41$	3.86	11.05	23.12	Ω.
				12.1				$2.36 \pm 0.19$	+	+	$3.08 \pm 0.42$	1.87	89.9	10.61	w
				16.2			+	$5.17 \pm 0.38$	+1 -	# -	$6.63 \pm 0.76$	8.24	16.49	25.47	M
				4.9			$2.99 \pm 0.16$	$7.34 \pm 0.39$	₩.	₩.	$7.96 \pm 0.54$	8.12	7.09	16.71	Ω.
				12.2		0 0	₩-	$3.40 \pm 0.29$	H -	# -	$4.97 \pm 0.78$	6.72	5.31	12.80	w c
				15.0			Н-	$17.83 \pm 0.97$	$-3.60 \pm 0.16$	H -	$21.5 \pm 1.2$	6.31	35.89	50.51	n ;
11101322622 50	50134602 2	289.9 -I	-14.5 :	2.5		2.687 ± 0.090	$2.95 \pm 0.11$	8.42 ± 0.33	$-3.61 \pm 0.12$	3.31 ± 0.18	$9.46 \pm 0.52$	18.71	8.90 7.7E	31.60	Z o
				4.4			+	6.33 + 0.39	1 + K	+	7 17 + 0 68	3.24	17.50	21.62	ი ≽
				10.1				$4.54 \pm 0.26$	+	+	$5.76 \pm 0.55$	2.24	14.09	16.87	s o
					MAXI_J1421-613		+	8.77 ± 0.68	1 +1	1 +1	$13.0 \pm 2.0$	2.36	50.36	53.95	ι Ω
_							1	$10.05 \pm 0.84$	1	1	$10.6 \pm 1.0$	13.52	20.97	42.54	· so
				4.5	XB_1940-04		+	$6.30 \pm 0.34$	$-2.67\pm0.12$	+	$9.2 \pm 1.0$	11.81	28.63	50.11	w
				10.2			#	$7.04 \pm 0.88$	+1	+	$7.9 \pm 1.3$	34.38	10.18	46.18	M
		_		10.8			₩.	$4.37 \pm 0.27$	$-3.19 \pm 0.17$	+ -	$5.07 \pm 0.51$	1.08	12.57	13.92	w i
۰.				1.0			₩-	$39.06 \pm 0.61$	H -	# -	$47.7 \pm 1.2$	14.42	33.44	53.68	w c
11102314077 50	50990080 2	280.9	5.6- 5.6-	. y.: 		$3.48 \pm 0.27$	$2.65 \pm 0.25$	$2.16 \pm 0.20$	$-2.88 \pm 0.22$	$3.64 \pm 0.63$	$2.97 \pm 0.52$	3.79	5.43	9.64	n c
				7.1			+ +	$3.62 \pm 0.43$	$-3.03 \pm 0.21$ $-3.07 \pm 0.30$	$2.76 \pm 0.55$	4.50 + 0.90	1.06	10.52	12.00	o vo
				6.7			1	$5.09 \pm 0.30$	1 +1	1	$6.20 \pm 0.58$	12.18	11.21	25.69	M
11102735801 51	51357400 8	84.3 15	12.3	5.0		$2.911 \pm 0.094$	$3.76 \pm 0.14$	$7.69 \pm 0.30$	$-3.22\pm0.10$	$4.65 \pm 0.29$	$9.49 \pm 0.61$	4.36	7.24	12.93	Ø
				3.4			+	.18 ±	+	+	$6.91 \pm 0.43$	12.68	16.71	30.37	M
				15.0			+ -	$2.32 \pm 0.18$	$-2.84 \pm 0.19$	+ -	$3.05 \pm 0.46$	3.61	4.01	8.54	w i
				13.9	40-0614+09		H -	H -	$-3.28 \pm 0.11$	H -	$6.45 \pm 0.42$	2.46	10.06	12.99	ν;
				5.6			H -	$8.42 \pm 0.32$		H -	$10.36 \pm 0.63$	2.65	12.36	17.26	Σū
11110159342 51 11110279675 51	51812949 2	247.1 -5	-59.0 I	12.2		3.43 ± 0.19	3.00 ± 0.23	5.97 ± 0.38	$-3.11 \pm 0.17$	$4.52 \pm 0.46$ $2.03 \pm 0.46$	7.38 ± 0.75	4.98 8 1.70	20.08	13.48	υŽ
				14.9	IIW Crb		+ +	++	3 63	++	$4.70 \pm 0.37$	4 96	7.73	13.70	Z V
				6.6	20-	· -	+	+	4 +	+ +	$4.37 \pm 0.05$	3.19	6.31	10.96	o vo
				6.9		$2.65 \pm 0.21$	Н	$4.83 \pm 0.47$	$-3.37 \pm 0.25$	1 +1	$5.89 \pm 0.79$	1.30	18.23	21.61	M
			~	4.7			+	$6.46\pm0.21$	$-3.129 \pm 0.091$	+	$8.14 \pm 0.42$	5.60	9.30	17.48	w
				9.8	4U0614+09		+	$10.59 \pm 0.27$	+	+	$12.66 \pm 0.52$	5.22	11.26	17.05	w
				7.5		0	+ -	+ -	$-2.93 \pm 0.21$	+ -	+1 -	1.77	12.59	14.60	M
11111220220 52	52724232 2	248.4 -3	-35.4	13.7		$2.69 \pm 0.20$	$4.87 \pm 0.42$	$5.96 \pm 0.52$	$-3.87 \pm 0.30$	$5.35 \pm 0.55$	$6.55 \pm 0.68$	2.77	11.95	17.01	$\infty$

Table 5:: GBM Type 1 Events continued from previous page

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Π	Feak	Ка	Dec	Error	Name(distance) (sigma)	$^{ m keV}$	$10-8 \text{ erg cm}^{-2} \text{ s}^{-1}$	$_{ m L}$ $_{ m 10^{-7}erg~cm^{-2}}$	FL index	$^{ m PL\ Flux}_{10^{-8}{ m erg\ cm}^{-2}\ { m s}^{-1}}$	$^{1}$ 10 <sup>-7</sup> erg cm <sup>-2</sup>	Kise	Fall	Duration Sec	Structure
11111244138	52748136	267.4	-10.9	10.1		$3.16 \pm 0.30$	1.62	$1.98 \pm 0.22$	$-3.17 \pm 0.28$	$2.08 \pm 0.35$	$2.54 \pm 0.42$	2.65	4.80	8.04	Ø
		245.4	-45.3	6.7		$3.02 \pm 0.16$	$2.57 \pm 0.17$	$5.25 \pm 0.35$	$-3.14 \pm 0.17$	+	$6.32 \pm 0.70$	4.42	11.52	16.47	Ø
		257.8	-26.3	10.3		$3.51 \pm 0.30$	$3.75 \pm 0.37$	$3.06 \pm 0.30$	$-3.46 \pm 0.32$	$3.62 \pm 0.59$	$2.95 \pm 0.48$	1.68	6.57	8.52	Ø
	_	287.4	-3.9	10.6		$3.31 \pm 0.15$	$2.70 \pm 0.14$	$5.52 \pm 0.30$	$-2.99 \pm 0.14$	$3.17 \pm 0.23$	$6.47 \pm 0.48$	12.70	12.29	25.57	M
11112341768	53696181	278.2	-11.1	00 u		$3.01 \pm 0.14$	$3.19 \pm 0.17$	$6.51 \pm 0.36$	$-3.26 \pm 0.15$	$3.64 \pm 0.26$	$7.43 \pm 0.53$	5.13	8.03	21.19	w >
		213.5	-25.5	1.2		$1.735 \pm 0.047$	$6.49 \pm 0.25$	$13.24 \pm 0.52$	$-3.38 \pm 0.10$ $-4.84 \pm 0.12$	$4.83 \pm 0.21$ $6.80 \pm 0.31$	$13.86 \pm 0.65$	8.59	33.13	44.17	Z S
		277.5	6.6-	7.1		$3.01 \pm 0.11$	$4.17 \pm 0.18$	$6.82 \pm 0.30$	$-3.25 \pm 0.12$	1	$8.29 \pm 0.58$	4.43	8.08	14.43	ω
		247.8	-54.1	7.7		$3.37 \pm 0.14$	$4.25\pm0.21$	$3.47 \pm 0.17$	$-3.08\pm0.13$	+	$4.25\pm0.37$	2.80	11.26	14.68	w
		296.3	-12.3	12.8			+	$6.40 \pm 0.31$	$-3.34 \pm 0.13$	+ +	$7.13 \pm 0.56$	2.51	14.55	17.62	Ø.
		270.8	-37.6	5.5		$2.894 \pm 0.069$	+ -	$16.69 \pm 0.48$	$-3.369 \pm 0.094$	+ -	$20.5 \pm 1.0$	17.43	5.06	31.49	oo o
		106.2	-14.4	11.7			₩ -	$3.98 \pm 0.18$	$-3.27 \pm 0.12$	# -	$4.90 \pm 0.34$	2.01	12.47	14.92	w o
		290.1	-66.2	15.5		$3.57 \pm 0.33$	₩ -	$1.73 \pm 0.19$	$-2.85 \pm 0.26$	H -	$2.39 \pm 0.49$	2.62	5.04	8.36	v c
11121262935	55358948	1.082	-14.2	о с 2 т	4II 0614±09	$2.96 \pm 0.10$ $3.290 \pm 0.065$	$3.04 \pm 0.12$ $7.36 \pm 0.17$	7.44 ± 0.31	$-3.27 \pm 0.11$ $-3.120 \pm 0.063$	$3.70 \pm 0.25$ $9.05 \pm 0.35$	$9.07 \pm 0.61$	12.40 3.85	12.59	17.87	ນ ທ
		324.5	-14.6	19.0			1 +1	$3.14 \pm 0.22$	$-2.98 \pm 0.17$	+ +	$3.73 \pm 0.35$	7.87	5.33	14.03	⊇⊠
11121631212	55672803	227.9	-60.8	5.5	Cir_X-1 (0.7) MAXI 11421_613 (1.1)	$3.056 \pm 0.097$		$9.78 \pm 0.38$	$-3.080 \pm 0.098$	$3.84 \pm 0.25$	$12.54 \pm 0.81$	2.38	20.35	24.08	M
11122006005	55993207	252.8	-53.6	° 65.	(T:T) 010-1741 0-1777	$3.12 \pm 0.12$	$4.42 \pm 0.21$	$5.41 \pm 0.26$	$-3.16 \pm 0.12$	$5.22 \pm 0.41$	$6.38 \pm 0.50$	5.62	8.39	15.48	Ø
11122409854		142.6	-62.7	3.5				+	$-3.03 \pm 0.10$	$4.35 \pm 0.28$		38.54	45.61	92.63	w
11122561717	56480921	255.8	-50.9	9.6		$3.31 \pm 0.22$	$4.39 \pm 0.33$	$3.58 \pm 0.27$	$-3.25 \pm 0.22$	$4.91 \pm 0.47$	$4.01 \pm 0.38$	2.57	7.35	10.88	Ø
		242.3	-63.5	2.9			#	+	+1	+	+	96.6	5.69	17.24	M
		261.7	-22.4	2.2			+ -	+1 -	$-3.24 \pm 0.10$	+ -	₩.	4.95	7.84	14.30	w :
		241.0	3.5	12.4	0 W-Crb	$3.83 \pm 0.40$	₩ -	$2.23 \pm 0.26$	$-2.77 \pm 0.28$	H -		2.41	7.85	10.68	n;
12010125204	57049204	253.0	-53.9	T. 6		$3.32 \pm 0.13$	$2.67 \pm 0.12$	$5.45 \pm 0.24$	$-3.01 \pm 0.11$	$3.22 \pm 0.18$	$6.57 \pm 0.38$	10.90	1.74	20.00	M o
		0.147	-01.0 28.5	0 0 0 7			H +	H +		H +	ΗН	60.4	14.00 6.03	11.75	ממ
_		253.8	-47.7	4. 6.			H +	H +	H +	H +	H +	2.45	10.13	14.61	ο σ
		233.8	-54.2	8.6			1 +1	1 +1	1	1 +1	1 +1	2.30	8.25	11.53	o o
		208.2	-60.9	14.0		$3.46 \pm 0.19$	1	1	1	1	1	2.50	7.20	10.75	ß
		287.2	9.9	6.6			+	+	+	+	+	14.44	20.23	38.30	SO
		282.9	-24.6	6.1		$2.66 \pm 0.13$	+1 -	+ -	$-3.49 \pm 0.17$	+1 -	$7.28 \pm 0.55$	9.68	18.55	29.56	ωo
12012420137	59031336	252.3	-41.7	20.8			$2.03 \pm 0.12$	H -	$-3.20 \pm 0.16$	2.18 ± 0.18	$6.24 \pm 0.52$	3.50	15.26	19.41	w c
		271.3	-21.0	11.9		$3.46 \pm 0.26$	$2.35 \pm 0.22$ $2.18 \pm 0.20$	2.97 + 0.27	$-3.18 \pm 0.20$ $-2.88 \pm 0.21$	H +	H +	3.01	9.25	12.97	o o
		244.6	-48.0	4.9			$3.38 \pm 0.11$	$8.29 \pm 0.28$	$-3.279 \pm 0.096$	1 +1	$9.91 \pm 0.53$	17.18	7.76	25.68	M
		270.4	-12.5	8.1			+	+1	# .	+1	$6.76 \pm 0.58$	13.34	9.44	23.10	M
		242.5	-24.3	21.9				$2.97 \pm 0.23$	$-2.97 \pm 0.19$	+ -	$4.01 \pm 0.55$	2.51	5.16	8.27	o o
12020448807	60010410	240.6	2.62-	16.4		$3.37 \pm 0.27$ $3.61 \pm 0.31$	$1.98 \pm 0.18$	$2.42 \pm 0.22$ $2.50 \pm 0.25$	$-3.17 \pm 0.26$ $-2.92 \pm 0.25$	$2.12 \pm 0.27$ $3.93 \pm 0.69$	$2.60 \pm 0.33$ $3.21 \pm 0.56$	4.57	3.47	8 80	מ מ
		223.5	-51.7	20.3			+ +	$2.30 \pm 0.26$ 2.30 + 0.26	$-2.68 \pm 0.27$	+	$2.57 \pm 0.38$	1.61	9.21	11.06	o oo
		234.2	-28.3	19.0				$2.97 \pm 0.22$	$-2.99 \pm 0.18$	1 +1	$3.41 \pm 0.33$	2.83	5.66	10.79	υx
	61223312	91.5	5.4	6.2	4U0614+09	$3.194 \pm 0.084$	+	$10.18 \pm 0.31$	$-3.130 \pm 0.083$	+	$12.54 \pm 0.65$	6.31	9.90	16.99	w
		249.0	-36.0	11.2			+	$3.89 \pm 0.20$	+1	+1	$5.07 \pm 0.47$	3.61	8.08	14.21	Ø
		237.9	-55.1	0.0			$2.75 \pm 0.17$	+ -	+1 -	+ -	$5.45 \pm 0.54$	1.62	10.08	12.05	o o
12022029147	613/3132	1.022	-53.4 FOR	ж. о		3.45 ± 0.18	3.03 ± 0.18	$2.47 \pm 0.14$	$-2.93 \pm 0.16$	$4.10 \pm 0.45$	$3.34 \pm 0.37$	67.7	7 64	9.35	ט מ
		259.4	-60.5	o o			H +	H +	H +	H +	H +	1.63	46.7	11.79	n w
		279.2	-15.6	8.4			1 +1	1 #1	1 +1	+	1 +1	2.52	13.85	16.75	M
	61650380	222.4	-68.6	2.9		Ö		$^{\rm H}$	$-3.67 \pm 0.44$	+	+	1.96	8.51	10.80	M
		254.6	-56.9	6.9		$3.20 \pm 0.14$	+	+	+	+	+	5.28	6.41	12.21	M
		237.6	-58.1	8.7		0	+ -	+ +	₩.	₩.	+ -	2.03	4.27	8.43	Ω :
12022745314	61994114	237.1	-54.8	4.9		3.55 ± 0.15	$4.92 \pm 0.26$	$4.01 \pm 0.21$	$-2.95 \pm 0.13$	$6.39 \pm 0.57$	$5.21 \pm 0.46$	2.63	8.87	13.60	ω o
_		271.9	-17.4	5. 00			H +	H +	H +	H +	H +	12.63	6.03	19.09	ი ≽
		244.0	-44.3	2.9		$2.82 \pm 0.15$	1 +1	+	+	1	$6.34 \pm 0.61$	2.26	11.39	14.07	ß
12022938463	62160066	100.5	6.1	6.1	4U0614+09	$3.43 \pm 0.13$	$3.88 \pm 0.17$	$4.76\pm0.21$		$5.20 \pm 0.44$	$6.37 \pm 0.54$	1.59	10.83	12.73	S

Table 5:: GBM Type 1 Events continued from previous page

f	-	ţ		ľ		ţ	ţ	ţ		į	ţ				
ì	S	110	3		(sigma)	keV	$10-8 \text{ erg cm}^{-2} \text{ s}^{-1}$	$10^{-7} {\rm erg \ cm^{-2}}$	7	$10^{-8} \text{erg cm}^{-2} \text{s}^{-1}$	$10^{-7} \text{ erg cm}^{-2}$	sec	sec	sec	o inconico
12022979307 6	62200915	246.0	-46.8	8.6		$3.43 \pm 0.29$		$2.16 \pm 0.22$	$-3.00 \pm 0.25$	$3.50 \pm 0.58$	$2.86 \pm 0.48$	1.40	8.99	10.60	S
	62337973		-44.8	5.2		$3.29 \pm 0.13$	$4.51 \pm 0.21$	+	$-3.03 \pm 0.12$	+	+	1.39	11.35	13.04	M
			-38.9	11.1		$3.45 \pm 0.22$	$2.28 \pm 0.17$	$^{\rm H}$	$-2.90\pm0.19$	$3.00 \pm 0.40$	$3.68 \pm 0.49$	1.98	4.80	10.86	w
			-52.5	6.3		$3.48 \pm 0.16$	+	+	2.90	$3.69 \pm 0.39$	+	1.45	8.27	26.6	S
_		01	-53.4	8.5		$2.77 \pm 0.10$	+ -	+ -	$-3.51 \pm 0.14$	+ -	$5.11 \pm 0.34$	3.33	7.32	11.45	Ω i
12030804891 6	62817689	7.4	-55.6	12.4		$1.90 \pm 0.13$	$1.84 \pm 0.16$	$3.01 \pm 0.26$	$-4.45 \pm 0.30$	$2.11 \pm 0.22$	$3.44 \pm 0.36$ 2 02 $\pm$ 0 24	3.10 7.00	8.91 15.05	12.74	n o
		257.1	-19.3	7.2		3.51 + 0.43	H +	H +	H +	H +	H +	9 00	21.39	26.14	ი ≽
		208.4	-45.6	5.4		$2.86 \pm 0.14$	+	+	$-3.39 \pm 0.17$	+	4.74 + 0.43	2.96	80.00	10.67	, o
	_	271.7	-19.0	5.7		$2.99 \pm 0.12$	1 11	+	3.34 #	+	$8.48 \pm 0.63$	10.82	7.43	19.02	Σ
		234.7	-15.3	16.2		$3.95 \pm 0.44$	$2.30 \pm 0.30$	+	$-2.74 \pm 0.30$	+	$2.50 \pm 0.60$	1.02	11.03	12.14	w
		237.0	-39.6	8.6		$3.48 \pm 0.17$	+	+	$-2.97 \pm 0.15$	+	$6.45 \pm 0.68$	2.06	11.40	19.52	M
12031853042 6	63729842	6.96	6.7	5.4	4U0614+09	$3.474 \pm 0.085$	$5.74 \pm 0.16$	$9.37 \pm 0.26$	$-3.046 \pm 0.076$	$7.23 \pm 0.34$	$11.80 \pm 0.56$	1.60	12.48	16.34	Ø
12032639643 6	64407644	257.4	-55.0	9.4			$2.58 \pm 0.16$	$5.26 \pm 0.34$	$-3.47 \pm 0.19$	$2.93 \pm 0.28$	$5.99 \pm 0.58$	13.43	1.90	16.71	M
		281.4	-31.3	11.8		0	+	$\mathbb{H}$	$-3.00 \pm 0.20$	$2.67 \pm 0.36$	$4.37 \pm 0.60$	4.31	4.42	14.89	M
		235.3	-39.9	10.6			$3.15 \pm 0.22$	+	+	+	$4.78 \pm 0.59$	2.29	29.9	9.45	w
		277.2	-5.9	7.8		0	+	+	+	+	+	3.46	6.50	18.13	M
		284.1	-19.7	9.2			+	+	$-3.18 \pm 0.20$	+	$2.96 \pm 0.37$	1.90	2.12	8.00	w
			-48.5	9.9		Ö	+	+	#	+	+	3.27	9.26	13.29	S.
_			-60.2	11.9			$2.49 \pm 0.23$	₩.	$-3.33 \pm 0.24$	# -	$2.52 \pm 0.35$	2.61	3.73	7.38	w j
			-16.4	00 00			₩.	₩.	$-3.22 \pm 0.12$	₩.	$7.40 \pm 0.51$	9.57	5.21	17.11	M
		259.0	-29.2	ος ( ος (			Н-	H -	Н-	H -		3.18	5.45	9.51	w c
			-44.9	180.0			$2.46 \pm 0.17$	H -	H -	H -	$3.71 \pm 0.44$	6.37	4.15	11.04	v c
			-15.7	10.7			$2.14 \pm 0.18$	₩.	$-2.45 \pm 0.17$	₩.	$4.61 \pm 0.94$	5.66	3.95	7.61	v i
			-60.7	4. 0		$3.13 \pm 0.15$		# -	$-3.16 \pm 0.15$	# -	$3.61 \pm 0.32$	2.16	8.72	11.23	w ;
			-18.4	xo 0		$2.85 \pm 0.14$	H -	Н-	н.	Н-	$6.90 \pm 0.69$	9.61	6.74	18.24	M
			-51.2	0.6			$5.81 \pm 0.21$	H -	$-2.716 \pm 0.081$	H -	$7.10 \pm 0.49$	2.55	6.91	12.44	w c
			-47.9	10.8		$3.09 \pm 0.19$	₩ -	H -	Н-	H -	$3.15 \pm 0.40$	2.80	4.65	8.00	N O
			0.2-	7.0	200	$3.12 \pm 0.10$	$4.05 \pm 0.15$	Н-	Н-	Н-	6.50 ± 0.39	2.63	10.94	14.05	N C
12042226680 6	66727495	150.1	-59.5	χ. τ.	2S_0918-549	3.05 ± 0.13	$2.64 \pm 0.14$	$4.32 \pm 0.23$	$-3.21 \pm 0.14$	$3.20 \pm 0.28$	$5.23 \pm 0.47$	6.10	16.52	23.86	n S
		0.017	2:1-	15.0 7.3		2.95 H 0.17	2.05 ± 0.20 2.046 ± 0.094	4.57 H 0.55	H +	3.13 H 0.39	3.21 ± 0.63	16.00	17.10 7.10 7.10	3E.01	M
		262.6	-40.2	5.1			$2.36 \pm 0.034$	++	$-2.329 \pm 0.092$ $-2.74 \pm 0.14$	++	4.18 + 0.52	2.39	6.06	9.73	Z V
		252.4	-49.5	6.9			+ +	+	+	+	3.89 + 0.35	2.95	5.24	10.61	o vo
			-15.7	5.9			1 +1	+	$-3.17 \pm 0.14$	+	$8.28 \pm 0.76$	2.08	13.29	15.82	Σ
			-52.8	4.1			1	+	$-3.01 \pm 0.13$	+	$4.81 \pm 0.41$	6.84	5.22	13.85	S
			-14.2	11.4			+	+	$-3.53 \pm 0.12$	+		2.61	16.18	19.23	M
12050706763 6			19.4	6.9			+	$4.58 \pm 0.16$	$-3.32 \pm 0.10$	$^{\rm H}$	$5.65 \pm 0.31$	2.67	5.57	8.75	w
			-39.5	18.4		$3.25 \pm 0.25$	+	+	$^{+}$	+	+	1.53	7.58	10.89	S
			-22.1	∞ i			₩.	# -	+ -	# -	$7.13 \pm 0.62$	3.30	5.87	17.34	M
12052156800 6	69263197	273.8	-12.3	7.0		<i>-</i>	$0.80 \pm 0.12$		-3.88 ± 0.46	H -	# -	11.09	8.29	20.22	Z Z
		250.4	43.0	0.01		3.37 H 0.23	H +	2.09 H 0.22 3 30 + 0 23	$-2.08 \pm 0.17$ $-3.24 \pm 0.20$	3.10 ± 0.78	4.10 H 0.03 3.75 + 0.43	13.32	15.20	16.70	N >
		265.9	2.86	22.2			2.73 + 0.26	+ +	+ +	+ +	2.87 + 0.48	6.36	3.93	10.65	
		248.6	-47.8	9.3			$2.60 \pm 0.13$	1	$-3.17 \pm 0.14$	1	1 +1	6.83	12.05	19.72	M
		272.5	-36.5	14.6			+	+	$-2.82\pm0.30$	+	+	7.20	3.46	11.62	M
		271.4	-6.7	7.8			$4.18 \pm 0.19$	$\mathbb{H}$	$-3.25 \pm 0.12$	+	$6.04 \pm 0.42$	1.38	12.19	15.51	M
		267.6	-50.6	9.3			#	#	#	+	#	4.23	11.63	16.79	Ω.
		271.6	-38.6	18.5			# -	# -	+ -	+ -	+ -	4.87	7.39	13.38	w c
		232.7	-31.9	11.5		<u> </u>	H -	Н-	Н-	H -	H -	3.13	6.69	10.72	n ;
12060479182 7	70495187	272.0	-34.3	11.6		$2.68 \pm 0.15$	$2.17 \pm 0.15$	$3.54 \pm 0.25$	$-3.34 \pm 0.19$	$2.64 \pm 0.28$ 6 0 ± 1 0		11.53	5.37	17.58	Ν
-			-51.9	16.0			ΗН	ΗН	ΗН	ΗН	4.35 H 0.66	1.90	14.04	10.00 7 FF	ם מ
		264.1	-29.1	7.3		3.82 + 0.26	+ +	+	+	+	+	1.40	8.17	00°-0	n w
		267.5	-13.4	0.6			#	1	1	1	1	13.11	5.81	19.20	M
12060876579 7	70838192	267.7	-43.8	6.4		$2.92 \pm 0.19$	$3.18 \pm 0.25$	$3.89 \pm 0.31$	$-3.48 \pm 0.24$	$3.50 \pm 0.41$	$4.28 \pm 0.51$	4.37	5.86	10.92	S
		259.1	-31.4	9.6		$3.91 \pm 0.23$	+	+	$-2.84\pm0.16$	$4.12 \pm 0.50$	$3.36\pm0.41$	1.67	15.16	17.04	ω

Table 5:: GBM Type 1 Events continued from previous page

							i			i				
ID Pe	Peak Ra s	» Dec	Error	Name(distance) (sigma)	$^{ m BB\ temp}_{ m keV}$	$^{10-8} erg cm^{-2} s^{-1}$	$_{10^{-7}\mathrm{erg~cm}^{-2}}$	PL index	$_{ m 10^{-8}erg~cm^{-2}~s^{-1}}$	PL Flnc $10^{-7}  \mathrm{erg \ cm^{-2}}$	Rise	Fall	Duration S	Structure
12060978999 7092	70927004 296.3	1	8.6		$3.35 \pm 0.20$	2.48	$2.02 \pm 0.14$	$-2.94 \pm 0.18$	$3.44 \pm 0.46$	$-\infty$	5.27	5.87	11.73	w
		.1 -34.2			$3.96 \pm 0.26$	Н	1 #1	$-2.84 \pm 0.18$	1 +1	$3.37 \pm 0.47$	1.17	1.50	66.9	S
					$4.38 \pm 0.65$	$2.84 \pm 0.49$	#	#	$5.8 \pm 2.0$	$2.39 \pm 0.84$	2.16	3.62	6.30	w
						$2.88 \pm 0.27$	#	#	$3.98 \pm 0.64$	$3.24 \pm 0.52$	2.87	3.55	7.32	S
					$3.01 \pm 0.16$	$2.25 \pm 0.14$	$2.76 \pm 0.17$	$-3.34 \pm 0.19$	+1 -	$3.30 \pm 0.33$	6.57	99.6	16.72	SO I
					$3.01 \pm 0.19$	+ +	$2.83 \pm 0.21$	$-3.31 \pm 0.21$	$2.60 \pm 0.31$	3.18 # 0.38	2.64	11.57	15.00	s >
120617046397 7154	715438192 289.7	1.0.4	T.T.		4.00 ± 0.32	$2.83 \pm 0.26$	2.31 ± 0.21 5 50 ± 0.30	$-2.79 \pm 0.23$	3.61 ± 0.66 5.77 ± 0.51	7.04 H 0.54	10.37	11.80	16.64	M
_					3.30 H 0.13	4.30 ± 0.24 3.36 ± 0.13	3.70 ± 0.30	H +	H +	7.00 H 0.02	10.04	4. 00 2. 17 2. 12	10.04	Ţυ
					$4.62 \pm 0.18$	+ +	+	2.57 +	4 +	2.56 + 0.62	2.50	2.00	5.26	n w
				4U_0614+09	$3.008 \pm 0.086$	1 +	$14.44 \pm 0.48$	$-3.279 \pm 0.092$	+	$18.15 \pm 0.90$	8.61	11.74	24.37	o vo
				4U_0614+09		1	$5.78 \pm 0.22$	1	1	$7.49 \pm 0.48$	1.39	7.45	10.93	w
12062060270 7185	71858667 254.4	.4 -26.3	3 12.2		$3.30 \pm 0.24$	$2.74 \pm 0.23$	$2.24 \pm 0.19$	$-3.04 \pm 0.22$	+	$2.79 \pm 0.40$	2.21	4.54	7.05	Ø
12062073109 7187	71871513 243.6	.6 -8.9			$3.17 \pm 0.18$	+	$^{\rm H}$	$-3.39 \pm 0.20$	+	$2.80 \pm 0.26$	2.67	3.51	6.82	w
		.0 -23.8				$2.21 \pm 0.27$	$1.80 \pm 0.22$	#	$\mathbb{H}$	$2.16 \pm 0.43$	2.93	6.94	10.48	S
		'			$3.33 \pm 0.21$	+	+	+	+	+	14.30	13.05	30.16	Ω
						+	+	+	+	+	1.99	5.83	18.05	Ø
						+	+	2.71 土	+	+	2.72	3.83	8.22	w
						+	.19 ±	+	+	$3.38 \pm 0.57$	2.77	5.54	8.71	S
						+	+	+	+	$3.10 \pm 0.78$	1.26	5.93	2.66	w
						+	.56 ±	+	+	+	5.56	4.27	10.08	w
	_					+	+	+	+	+	8.28	16.07	26.42	w
						$\mathbb{H}$	$\mathbb{H}$		+	+	8.35	16.70	25.71	w
12071115416 7362	73628227 166.4	.4 -52.2				$2.80 \pm 0.15$	$\mathbb{H}$	$-3.44 \pm 0.16$	$\mathbb{H}$	$5.20 \pm 0.42$	9.81	6.16	16.76	ß
12071729405 7416	74160607 149.9	.9 -44.4	4 11.9	2S_0918-549 (1.1)	$4.96 \pm 0.77$	$0.74 \pm 0.13$	$2.44 \pm 0.42$	$-2.40 \pm 0.33$	$1.33 \pm 0.48$	$4.3 \pm 1.5$	10.22	19.20	31.49	S
				GS_0836-429 (1.3)								_		
				$4U_{-}0614+09$		+	+ 1	+	+	+ .	6.22	15.69	23.69	w
_						₩ .	₩.	₩.	₩.	₩.	1.33	3.29	8.78	Ω į
						+	+	+	+	+	8.32	10.44	20.13	S
						+	#	41 -	#	+	1.62	21.33	23.47	M
•						₩ -	₩ -		+ -	₩ -	2.80	5.53	8.81	ω ;
						H -	H -	Н-	H -	$5.81 \pm 0.62$	12.72	7.41	20.64	Z,
					$2.92 \pm 0.20$	₩-	H -	H -	H -	$4.34 \pm 0.56$	8.37	5.77	14.59	Z <sup>©</sup>
							H -	H -	H -	$2.33 \pm 0.98$	5.64	4.69	10.99	n ;
					$3.19 \pm 0.19$	H -	H -	Н-	H -	$6.19 \pm 0.43$	10.80	4.91	16.38	Z Z
	75367781 245.6				$3.52 \pm 0.18$	$2.80 \pm 0.17$	$3.43 \pm 0.21$	2.90 ±	H -	$4.56 \pm 0.52$	7.11	3.11	10.57	Σū
12080230593 7554	75544208 287.8 76664696 196 9	0.62-8.	7.4		$3.34 \pm 0.23$ $2.11 \pm 0.24$	ΗН	ΗН	$-2.94 \pm 0.21$	2.63 ± 0.40	7.5 H 1.1	14.95	11.52	30.70	ט מ
Ì					3.11 H 0.24	ΗН	ΗН	-4.00 H 0.34	НН	0.04 H 0.39	60.1	11.02	24.53	υ Σ
						+ +	+ +	3 39 +	+ +	$7.20 \pm 0.45$	2.43	7.15	17.94	
Ī.					$3.71 \pm 0.40$	1 +	1 +	1 +	1	$2.14 \pm 0.54$	2.95	3.05	7.03	ß
	_				$2.93 \pm 0.12$	+	$5.64 \pm 0.29$	3.44 ±	1	$6.25 \pm 0.51$	1.20	68.9	20.64	M
12080974583 7619	76192976 272.8	.8 -23.4	4 7.2		$2.74 \pm 0.14$	+	$5.09 \pm 0.32$	$-3.54 \pm 0.19$	$2.29 \pm 0.21$	$5.62 \pm 0.53$	10.87	69.7	19.17	M
						+	$3.06 \pm 0.27$	+	+	$3.56 \pm 0.48$	12.03	4.34	17.29	M
٠.						+	$2.89 \pm 0.17$	$\mathbb{H}$	+	$3.64 \pm 0.37$	2.94	4.37	8.89	w
						₩.	$5.77 \pm 0.37$	+ -	+ -	$6.99 \pm 0.78$	2.32	23.11	26.42	M
						₩.	$116.6 \pm 2.8$	# -	₩.	$136.9 \pm 4.9$	70.93	66.24	157.61	w ;
						$6.13 \pm 0.25$	$10.01 \pm 0.41$	+1 -	# -	$12.47 \pm 0.82$	11.35	9.72	21.61	W W
						H -	H -	Н-	H -	$3.92 \pm 0.48$	14.01	10.79	25.92	ν;
· 1					$3.32 \pm 0.11$		H -	$-3.10 \pm 0.10$	5.38 ± 0.35	$6.59 \pm 0.43$	1.15	12.89	14.28	M o
12082549804 (750	77560018 254.8	- 29.1 - 4	2.8		20 ± 7.7.2	4.05 H 0.26	4.95 ± 0.31	-3.00 ± 0.19	4.42 ± 0.38	$0.41 \pm 0.40$	6.03	5.04	10.30	Ω >
						+ +	++	4 +	+ +	3 89 + 0 88	230	7.55	10.21	Z V
						+	+	2.67 +	+	5.28 + 0.96	1.39	6.70	10.22	o oo
						1	1	1	1	1	15.89	9.70	26.47	M
12090880535 7879				4U_0614+09			+	+	+	+	2.18	7.61	10.97	Ø
12090947623 7884	78844430 295.5	5 -3.4	15.9		$3.46\pm0.28$	$1.74 \pm 0.16$	$3.55\pm0.34$	$-3.28\pm0.28$	$1.89 \pm 0.30$	$3.85 \pm 0.62$	14.10	5.15	19.87	M

Table 5:: GBM Type 1 Events continued from previous page

ID	Peak	Ra	Dec	Error	Name(distance)	BB temp	BB flux		PL index	PL Flux		Rise	Fall	Duration	Structure
	œ				(sigma)	keV	$10-8 \text{ erg cm}^{-2} \text{ s}^{-1}$	$10^{-7} \rm erg \ cm^{-2}$		$10^{-8} \mathrm{erg} \ \mathrm{cm}^{-2} \ \mathrm{s}^{-1}$	$10^{-7} \text{ erg cm}^{-2}$	sec	sec	sec	
12090966157 7	78862958	230.5	-28.7	13.6		$3.21 \pm 0.30$	$4.11 \pm 0.37$	$3.35 \pm 0.30$	$-4.14 \pm 0.43$	$5.34 \pm 0.48$	$4.36 \pm 0.39$	2.86	4.68	9.45	S
12091235378 7	79091385	272.3	-35.5	7.2		$3.11 \pm 0.17$	$3.14 \pm 0.21$	$3.85 \pm 0.26$	$-3.19 \pm 0.18$	$3.77 \pm 0.40$	$4.61 \pm 0.49$	5.37	8.81	15.39	S
12091323339 7	79165740	267.0	-18.9	8.9		$3.09 \pm 0.17$	$2.64 \pm 0.18$	$5.39 \pm 0.37$	$-3.28 \pm 0.19$	$2.99 \pm 0.34$	$6.11 \pm 0.70$	14.11	3.66	18.31	M
12091828018 7	79602421	91.5	7.2	11.9	$4U_{-}0614+09$	$3.11 \pm 0.13$	$3.28 \pm 0.16$	$4.02 \pm 0.20$	$-3.19 \pm 0.14$	$4.05 \pm 0.34$	$4.96 \pm 0.42$	1.26	16.21	17.81	Ø
12092148059 7	79881673	258.2	-52.0	9.6		$3.58 \pm 0.21$	$5.19 \pm 0.35$	$2.11 \pm 0.14$	$-2.98 \pm 0.17$	$6.61 \pm 0.75$	$2.69 \pm 0.30$	4.45	5.55	10.39	S
12092217967 7	79937968	269.5	-41.0	6.6		$3.11 \pm 0.17$	$2.49 \pm 0.16$	$3.05 \pm 0.20$	$-3.14 \pm 0.17$	$3.13 \pm 0.33$	$3.84 \pm 0.41$	4.75	4.82	10.21	S
12100281652 8	80865648	138.6	-72.9	0.9		$2.96 \pm 0.11$	$2.22 \pm 0.10$	$6.35 \pm 0.28$	$-3.19 \pm 0.12$	$2.78 \pm 0.21$	$7.95 \pm 0.60$	10.14	15.02	26.49	S
12100862493 8	81364893	188.6	-17.2	1.7		$3.16 \pm 0.16$	$4.29 \pm 0.20$	$10.50 \pm 0.50$	$-3.96 \pm 0.21$	$5.66 \pm 0.26$	$13.86 \pm 0.63$	4.01	11.25	15.98	S
12101875093 8	82241502	274.1	-23.0	9.6		$2.99 \pm 0.12$	$4.38 \pm 0.23$	$5.36 \pm 0.28$	$-3.20 \pm 0.13$	$5.30 \pm 0.43$	$6.48 \pm 0.53$	9.03	3.54	13.31	M
12101900479 8	82253284	241.1	-42.8	12.4		$3.23 \pm 0.28$	$2.72 \pm 0.28$	$2.22 \pm 0.23$	$-3.20 \pm 0.28$	$3.02 \pm 0.50$	$2.46 \pm 0.40$	3.16	8.25	11.90	S
12102423481 8	82708271	274.0	7.7-	9.9		$3.36 \pm 0.16$	$4.02 \pm 0.19$	$4.93 \pm 0.23$	$-3.72 \pm 0.20$	$5.02 \pm 0.26$	$6.16 \pm 0.32$	9.18	3.46	14.07	M
12102585037 8	82856237	238.1	-68.7	11.3		$3.28 \pm 0.32$	$3.72 \pm 0.35$	$3.04 \pm 0.28$	$-4.03 \pm 0.43$	$4.90 \pm 0.45$	$3.99 \pm 0.37$	3.14	5.60	9.21	S
12103137028 8	83326622	274.6	-10.7	6.5		$3.41 \pm 0.11$	$4.82 \pm 0.19$	$5.91 \pm 0.23$	$-2.994 \pm 0.097$	$6.03 \pm 0.40$	$7.39 \pm 0.49$	8.98	4.39	13.86	M
12110910913 8	84078119	273.4	-4.0	8.9		$2.820 \pm 0.098$		$6.48 \pm 0.27$	$-3.41 \pm 0.12$	$3.72 \pm 0.23$	$7.59 \pm 0.48$	1.49	9.31	21.10	M
12111324503 8	84437310	259.3	-18.0	9.7		$2.74 \pm 0.18$	$3.91 \pm 0.27$	$6.39 \pm 0.44$	$-4.54 \pm 0.33$	$5.01 \pm 0.35$	$8.18 \pm 0.57$	1.54	5.57	17.37	M
12111517342 8	84602951	275.8	-27.9	9.6		$3.17 \pm 0.12$	$3.78 \pm 0.18$	$6.17 \pm 0.30$	$-3.09 \pm 0.12$	$4.79 \pm 0.37$	$7.82 \pm 0.61$	1.79	3.20	15.34	M
12112127678 8	85131694	256.5	-35.3	9.7		$3.22 \pm 0.18$	$3.13 \pm 0.20$	$2.56 \pm 0.16$	$-3.25 \pm 0.18$	$3.76 \pm 0.36$	$3.07 \pm 0.30$	5.02	9.04	14.97	S
12121216694 8	86935093	284.0	6.9	6.6		$3.21 \pm 0.11$	$2.75 \pm 0.12$	$6.74 \pm 0.30$	$-3.12 \pm 0.11$	$3.29 \pm 0.24$	$8.07 \pm 0.60$	2.76	15.88	19.22	M
12121222136 8	86940537	153.4	-61.0	8.9	$2S_{-}0918-549$	$2.89 \pm 0.12$	$2.49 \pm 0.13$	$4.07 \pm 0.21$	$-3.21 \pm 0.14$	$3.17 \pm 0.27$	$5.18 \pm 0.44$	4.25	15.82	21.14	S
12121382175 8	87086961	94.6	7.4	1.7	$4U_{-}0614+09$	$3.219 \pm 0.048$	$4.042 \pm 0.071$	$34.65 \pm 0.61$	$-3.150 \pm 0.048$	$4.90 \pm 0.14$	$42.0 \pm 1.2$	5.74	43.14	51.37	S
13010467917 8	88973529	264.8	-15.6	4.8		$3.12 \pm 0.10$	$3.22 \pm 0.12$	$7.89 \pm 0.31$	$-3.24 \pm 0.11$	$3.89 \pm 0.23$	$9.52 \pm 0.57$	11.34	13.46	25.75	M
13012372764 9	90619972	257.4	-38.8	1.1		$3.132 \pm 0.073$	$2.664 \pm 0.074$	$39.1 \pm 1.0$	$-3.167 \pm 0.074$	$3.22 \pm 0.14$	$47.4 \pm 2.1$	34.91	116.28	183.08	M
13012978737 9	91144343	97.3	-3.1	10.3	$4U_{-}0614+09$	$3.17 \pm 0.13$	$2.97 \pm 0.15$	$4.85 \pm 0.25$	$-3.03 \pm 0.13$	$3.91 \pm 0.36$	$6.39 \pm 0.59$	2.80	17.43	20.55	S
13020216053 9	91427258	276.7	-19.4	8.5		$3.03 \pm 0.12$	$3.53 \pm 0.18$	$5.77 \pm 0.29$	$-3.25 \pm 0.13$	$4.21 \pm 0.32$	$6.88 \pm 0.53$	12.11	4.04	16.89	M
13020579710 9	91750115	102.3	-3.1	9.2	$4U_{-}0614+09$	$3.64 \pm 0.13$	$5.19 \pm 0.22$	$6.36 \pm 0.27$	$-2.93 \pm 0.10$	$6.71 \pm 0.52$	$8.22 \pm 0.63$	2.63	8.93	12.14	S
13021350602 9	92412210	275.5	-5.9	12.2		$2.81 \pm 0.21$	$1.90 \pm 0.18$	$2.32 \pm 0.22$	$-3.26 \pm 0.25$	$2.28 \pm 0.33$	$2.79 \pm 0.41$	11.76	9.68	24.56	S
13022239487 9	93178695	272.6	-22.5	17.3		$3.38 \pm 0.19$	$2.65 \pm 0.18$	$3.24 \pm 0.22$	$-3.03 \pm 0.18$	$3.30 \pm 0.39$	$4.05 \pm 0.48$	7.61	9.55	17.51	M

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## A. aFXPs Light Curves

See http://gammaray.msfc.nasa.gov/gbm/science/xrb.html

## B. uGRBs Light Curves

See http://gammaray.msfc.nasa.gov/gbm/science/xrb.html

## C. tXRBs Light Curves

See http://gammaray.msfc.nasa.gov/gbm/science/xrb.html